

International Astrophysical Consortium for High Energy Calibration

Galaxy Clusters WG Summary

Eric Miller (MIT)
19 May 2021

Clusters WG Recent History

- Jukka left IACHEC shortly before 2019 Workshop in Japan. Eric took over the Clusters WG and Herman took over the Contamination WG.
- The Clusters WG has been dormant for two years.
- WG was previously very active:
 - "Cross-calibrating X-ray detectors with clusters of galaxies: an IACHEC study" Nevalainen, David & Guainazzi 2010, A&A, 523, 22
 - "Cross-calibration of Suzaku XIS and XMM-Newton EPIC using clusters of galaxies" Kettula, Nevalainen & Miller 2013, A&A, 552, 47
 - "XMM-Newton and Chandra Cross Calibration Using HIFLUGCS Galaxy Clusters" Schellenberger et al. 2015, A&A, 575, 30
- It should be again!
 - eROSITA, XRISM, ATHENA will do lots of cluster science and use cluster calibration targets.
 - The Multi-Mission Study is in progress and promising useful results.

WG Membership

- E. Miller (chair, XRISM, Hitomi, Suzaku/XIS)
- A. Beardmore (Swift/XRT)
- M. Bonamente
- Y. Chen (HXMT)
- Y.-P. Chen (HXMT)
- L. David (Chandra)
- J. de Plaa
- K. Forster (NuSTAR)
- C. Grant (Chandra/ACIS)
- S. Jia
- C. Li

- K. Madsen (NuSTAR)
- H. Matsumoto (XRISM/Xtend, Hitomi/SXI, Suzaku/XIS)
- N. Ota (XRISM, Hitomi, Suzaku/XIS)
- A. Read (XMM-Newton/EPIC-MOS)
- G. Schellenberger (XMM-Newton/EPIC, Chandra/ACIS)
- S. Snowden (XMM-Newton/EPIC-MOS)
- M. Stuhlinger (XMM-Newton/EPIC)
- I. Valtchanov (XMM-Newton/EPIC)
- N-J. Westergaard (NuSTAR)
- H. Zhao (HXMT)

WG Membership at meeting last Tuesday

- E. Miller (chair, XRISM, Hitomi, Suzaku/XIS)
- A. Beardmore (Swift/XRT)
- M. Bonamente
- Y. Chen (HXMT)
- Y.-P. Chen (HXMT)
- L. David (Chandra)
- J. de Plaa
- K. Forster (NuSTAR)
- C. Grant (Chandra/ACIS)
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WG Meeting

WG Meeting #1: 2021-05-11 (virtual)

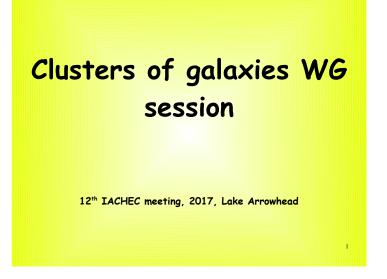
Solution Signature Si

Meeting information

- Date: May 11, 2021 at 07:00 PDT = 10:00 EDT = 14:00 UTC = 16:00 CEST = 22:00 CST = 23:00 JST
- · Connection: Zoom connection circulated by email
- · Attendees: Eric Miller, Catherine Grant, Andy Beardmore, Chen Yong, Gerrit Schellenberger, Hiro Matsumoto, Ivan Valtchanov, Karl Forster, Kristin Madsen, Larry David
- Next meeting: TBD

Agenda

Where the Clusters WG left off





Where do we go from here?



Program

- 1) Erosita cross-calibration with clusters
- 2) Multi-mission review from last year
- 3) Updates on the Multi-Mission Study

(A/I Eric) Contact Konrad about eROSITA cluster calibration targets and observations (A1795, A2029, ...)

Described below.

Multi-Mission Study (MMS)

Review 2015

- ★ Comparison of cluster spectra measured with XMM-Newton/EPIC, Chandra/ACIS, Swift/XRT, Suzaku/XIS, ROSAT/PSPC i.e. 5 missions, 10 instruments
- * Residual ratios to evaluate the effective area cross-calibration:
 - We use EPIC-pn as a reference. (Try also ACIS, TBD)
 - For instrument i we calculate the mean of the ratio

$$R_{i/pn} = \frac{data_i}{model_{pn} \otimes resp_i} \times \frac{model_{pn} \otimes resp_{pn}}{data_{pn}}$$

The latter term corrects for deviations btw. pn model and pn data which cannot be produced by the model (no point in comparing reference instrument with another using a model which does not fit the reference instrument data)

Multi-Mission Study (MMS)

Review 2015

Sample

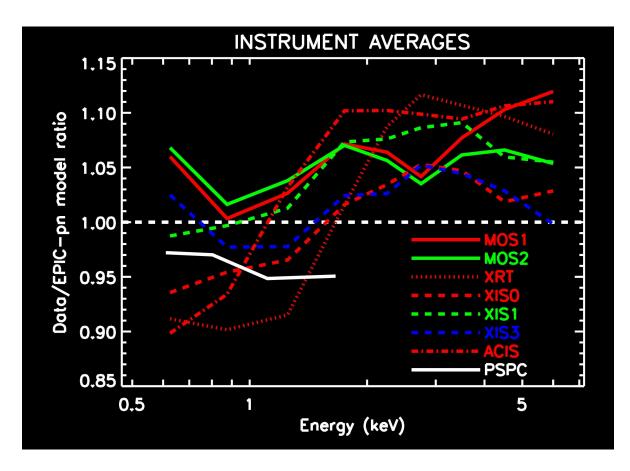
* Currently the sample consists of

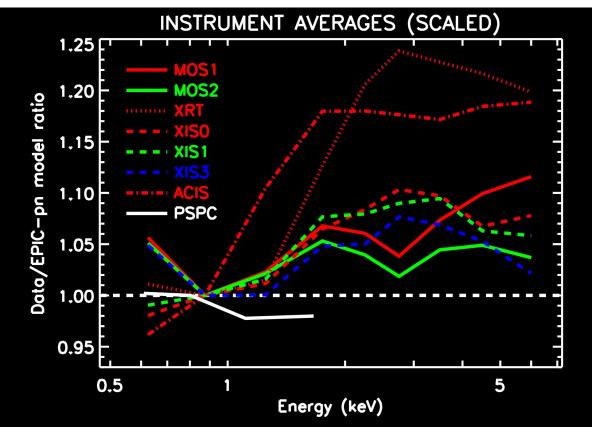
A1795, A2029, Coma and PKS 0745-19

	Satellite/instrument	Date of processing	Software/CALDB
	XMM-Newton/EPIC	April 2014	xmmsas_20131209_1901- 13.5.0
	Chandra/ACIS	May 2014	ciao-4.6
	Swift/XRT	April 2014	
	Suzaku/XIS	May 2014	xissimarfgen 2010-11-05 ae_xi0_contami_20130813.fits
	ROSAT/PSPC-B	May 2013	

At the moment the results apply to calibration status on May IACHEC 2021 — Galaxy Clusters WG Sumpry 14

Multi-Mission Study (MMS) – Review

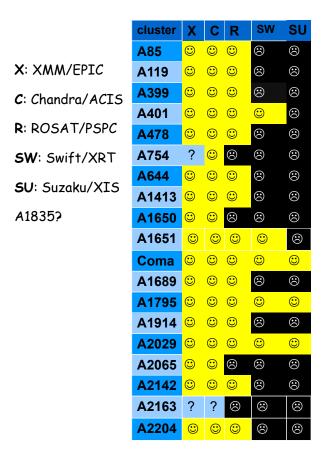


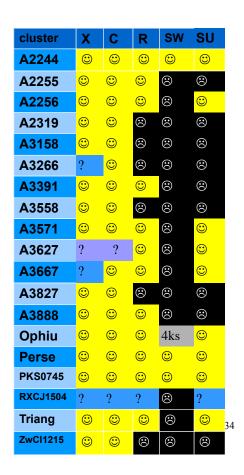


Multi-Mission Study (MMS) – Review

- 2017 TASK 1: Define the extraction radius
 - Statistics requirement and PSF minimization prefer bigger values
 - BG and vignetting minimization prefer smaller values
 - Extraction radius = 6 arcmin
- 2017 TASK 2: Define suitable clusters
 - Hot enough (minimize 1 keV line emission, better source/BG ratio at 7 keV)
 - Not too distant to yield enough photons
 - Prefer low NH, but no requirement for NH at the moment
 - kT > 6 keV, z < 0.1
- 2017 TASK 3: Define suitable observations
 - Long enough for statistics.
 - Cluster centered on-axis.
 - Use single observations. If too constraining, we will discuss merging observations.
 - 100,000 cts in central 6 arcmin (40,000 for PSPC)
 - Center < 3 arcmin off-axis

Multi-Mission Study (MMS) – Review





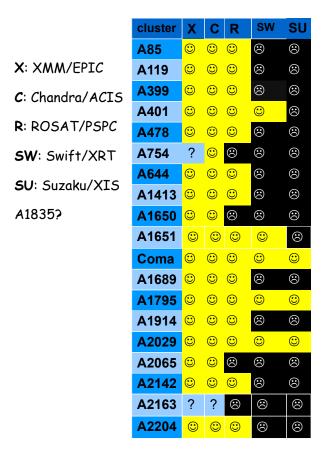
- Cluster sample is hot HIFLUGCS (+ others) that meet these criteria (plus Perseus), and:
 - Offset btw. the cluster center and pointing FOV center < 3 arcmin
 - Exposure > 10 ks in the available data

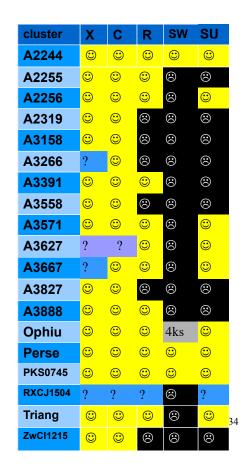
Multi-Mission Study (MMS) - Review

- 2017 TASK 4: Compile a list of available clusters and obs. ID:s fulfilling our criteria: Larry (Chandra), Eric (Suzaku), Andy B. (Swift), Steven Snowden (ROSAT), Jukka (XMM) Deadline end of April
- 2017 TASK 5: Extract and process data with May 2017 calibration information. Deadline end of June
- 2017 TASK 6: Jukka will do the stack residuals ratio analysis.

 Only XMM-ROSAT comparison has been done, for ~12 clusters, presented at 2018 IACHEC.

Multi-Mission Study (MMS) – Update





- Swift now has lots of data on PKS0745.
- Chandra has lots of data on A1795.
- Ivan has worked on A1795 work flow.
- Jukka, Ivan et al. worked on chip gap/bad pixel correction on pn using image from MOS
 - Reduces the effect on the ARF to 0.1% (TN: https://arxiv.org/abs/2103.01753)

Multi-Mission Study (MMS) – Update

- Action items from 2018
 - 2018 TASK 1: Check ROSAT PSPC calibration using one of our clusters (Jukka & M. Freyberg) (Postpone)
 - 2018 TASK 2: Check one cluster with Konrad's methods. Needs isothermal region for simple and accurate modelling. (Postpone)
 - 2018 TASK 3: Swift XRT flux weighting of ARFs (A/I Andy B.) XRT mkarf doesn't flux-weight ARFs, so Andy has to figure out how to do this by hand.
 - 2018 TASK 4: Draft ready by next IACHEC (Postpone)
- Where do we go from here?
 - Complete multi-mission study
 - Supply data to the concordance effort (Statistics WG)

Multi-Mission Study (MMS) – Update

Instrument	Contact Person	Notes
XMM EPIC MOS & pn	Ivan	
Chandra ACIS	Larry & Gerrit	
Suzaku XIS	Eric	
Swift XRT	Andy B.	
ROSAT PSPC	?	(A/I Eric) Ask Konrad and Michael F. about ROSAT cluster person.
NuSTAR	Karl and Dan Wik	NuSTAR LP, A2029, A478, A1795, A2199, all completed according to Karl (A/I Eric) contact Dan and Karl about NuSTAR data
eROSITA	?	(A/I Eric) Ask Konrad and Michael F. about eROSITA cluster person and data set.
ASTROSAT	?	(A/I Eric) Ask ASTROSAT team about cluster person and data set.
HXMT	Chen Yong	Only two clusters have been observed: Coma, Perseus
NICER	?	(A/I Eric) Ask about cluster observations/calibration, interest in cross-calibration, contact.

Summary and Future Plans

- CluWG is still useful, will be more so for future missions.
- We have a path forward for the MMS.
 - Assign WG member for each mission. Include additional missions.
 - Identify and gather existing data.
 - Select ~4 clusters to include.
 - Update calibration.
 - Extract spectra and responses, provide to Eric for residual ratio determination.
- Provide data for Calstats WG concordance effort.

Two CluWG meetings planned before September IACHEC Workshop.
 If in person, we hope to work real-time on this effort.