

# ISR Theory 2

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- 1 Thompson Scattering
- 2 Scatter from Multiple Electrons
- 3 Collective Interactions

# Warning on Notation

## Engineering Notation

$$e^{j\omega t - j\mathbf{k}\cdot\mathbf{r}}$$

- Used in most antenna theory textbooks
- Used in the Kudeki and Milla [2011] IEEE review paper on ISR theory

Conversion between notations is  $j = -i$

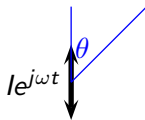
## Physicist Notation

$$e^{-i\omega t + i\mathbf{k}\cdot\mathbf{r}}$$

- Used in Jackson E&M textbook
- Used in most plasma physics textbooks
- Used for solution to Landau problem

# Hertzian Dipole Antenna

Consider an infinitesimal dipole antenna of length  $d\ell$  carrying current a sinusoidal current  $I$



$$\mathbf{J} = Id\ell \delta(\mathbf{x}) \hat{\mathbf{z}} e^{j\omega t}$$

Far Field Solution ( $\eta_0 \equiv \sqrt{\frac{\mu_0}{\epsilon_0}} = 377 \Omega$ )

$$\mathbf{E}_{ff} = \frac{jk_0\eta_0 Id\ell}{4\pi r} \sin\theta e^{j\omega t - jk_0 r} \hat{\theta}$$

$$\mathbf{B}_{ff} = \frac{jk_0\mu_0 Id\ell}{4\pi r} \sin\theta e^{j\omega t - jk_0 r} \hat{\phi}$$

Far Field Radiated Power

$$S = \frac{1}{2\mu_0} \Re \{ \mathbf{E} \times \mathbf{B}^* \} = \frac{1}{2\eta_0} |E|^2 = \frac{\eta_0}{2} \left( \frac{k_0 Id\ell}{4\pi r} \right)^2 \sin^2\theta$$

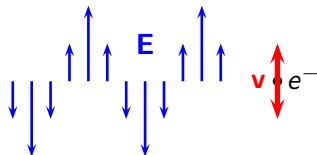
# Thompson Scatter from One Electron

Incident wave:

$$\mathbf{E} = \hat{z} E_0 e^{j\omega t - j\mathbf{k}_0 \cdot \mathbf{r}}$$

Motion of the electron:

$$j\omega m_e \mathbf{v} = -e\mathbf{E} \rightarrow \mathbf{v} = \frac{je}{\omega m_e} E_0 \hat{z}$$



Effective Hertzian Dipole with  $Idl \rightarrow e\mathbf{v}$  (also note  $\omega/k = c$ )

$$\mathbf{E}_{scat} = \frac{-\eta_0 e^2}{4\pi r m_e c} E_0 \sin \theta e^{j\omega t - j\mathbf{k}_0 \cdot \mathbf{r}} \hat{\theta} = -\frac{r_e}{r} E_0 e^{j\omega t - j\mathbf{k}_0 \cdot \mathbf{r}} \hat{\theta}$$

Where the classical electron radius is

$$r_e = \frac{\eta_0 e^2}{4\pi m_e c} = \frac{e^2}{4\pi \epsilon_0 m_e c^2} \approx 2.818 \times 10^{-15} \text{ m}$$

# Thompson Scatter Cross section

Total Cross Section:

$$\sigma_t \equiv \frac{P_{\text{tot}}}{\frac{1}{2\eta_0} |E_0|^2} = \frac{8\pi}{3} r_e^2 \quad \text{Where } P_{\text{tot}} \equiv \int_0^{2\pi} \int_0^\pi S_{\text{scat}} r^2 \sin \theta d\theta d\phi$$

Radar Cross Section:

$$\sigma = \sigma_t D_s$$

Directivity of scattering in the direction towards the radar:

$$D_s \equiv \frac{S_{\text{scat}}(\text{at the radar})}{\frac{P_{\text{tot}}}{4\pi r^2}}$$

For a Herztian dipole,  $S \propto \frac{\sin^2 \theta}{r^2}$ ,  $D_s(\theta, \phi) = \frac{3}{2} \sin^2 \theta$ .

For backscatter  $\theta = 180^\circ$ , so the radar cross section of one electron is

$$\sigma = 4\pi r_e^2 \approx 10^{-28} \text{ m}^2 \quad (\sim 0.9979 \times 10^{-28} \text{ m}^2)$$

# Why Can We Ignore the Ions?

$$\sigma_e \propto \frac{1}{m_e^2}$$

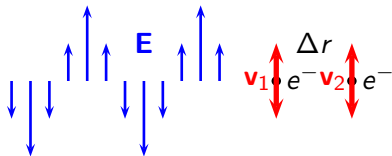
The scattering cross section of an ion is

$$\sigma_i = \frac{m_e^2}{m_i^2} \sigma_e$$

For an  $O^+$  plasma

$$\frac{m_e^2}{m_i^2} = 1.16 \times 10^{-9}$$

# Scatter from Two Electrons



Incident on first electron:

$$E_1 = E_0 e^{j\omega t}$$

Scattered from first electron:

$$\begin{aligned} E_{s1} &= -\frac{r_e}{r} E_1 e^{-jk_0 r} \\ &= -\frac{r_e}{r} E_0 e^{j\omega t - jk_0 r} \end{aligned}$$

In the far field  $\frac{1}{r+\Delta r} \approx \frac{1}{r}$ , so the sum of the fields is

Incident on second electron:

$$E_2 = E_0 e^{j\omega t - jk_0 \Delta r}$$

Scattered from second electron:

$$\begin{aligned} E_{s2} &= -\frac{r_e}{r + \Delta r} E_2 e^{-jk(r + \Delta r)} \\ &= -\frac{r_e}{r} E_0 e^{j\omega t - jk_0 r - j2k_0 \Delta r} \end{aligned}$$

$$E_{s1} + E_{s2} = -\frac{r_e}{r} E_0 e^{j\omega t - jk_0 r} \left( 1 + e^{-j2k_0 \Delta r} \right)$$



# Bragg Wavelength

For scatter from two electrons

$$|E_{s1} + E_{s2}|^2 \propto |1 + e^{-j2k_0\Delta r}|^2 = 4 \cos^2(k_0\Delta r)$$

- If  $\Delta r = \frac{\lambda}{2}$ ,  $k_0\Delta r = \pi$ , and the factor is 4 (perfect constructive interference)
- If  $\Delta r = \frac{\lambda}{4}$ ,  $k_0\Delta r = \frac{\pi}{2}$ , and the factor is 0 (perfect destructive interference)
- If  $\Delta r$  is a random number, the expected value of the factor is 2.

The Bragg wavelength  $\lambda_b = \frac{\lambda_0}{2}$  is the preferred spacing where the scatter adds constructively.

Define the Bragg wavenumber (for backscatter) as  $k_b = \frac{2\pi}{\lambda_b} = \frac{4\pi}{\lambda_0} = 2k_0$ .

# Scatter from Many Electrons

$$\begin{aligned} E_s &= -\frac{r_e}{r} E_0 e^{j\omega t - jkr} \left( \sum_{p=0}^{N-1} e^{-j2\mathbf{k}_0 \cdot \Delta \mathbf{r}_p} \right) \\ &= -\frac{r_e}{r} E_0 e^{j\omega t - jk_0 r} \int n_e(\mathbf{r}) e^{-j2\mathbf{k}_0 \cdot \Delta \mathbf{r}_p} d^3r \end{aligned}$$

where the microscopic electron density is

$$n_e(\mathbf{r}) \equiv \sum_{p=0}^{N-1} \delta(\mathbf{r} - \Delta \mathbf{r}_p)$$

This looks like a spatial Fourier transform evaluated at the Bragg wavenumber  $\mathbf{k}_b = 2\mathbf{k}_0$ .

The scatter is most sensitive to density structures at the Bragg wavelength.

# Coherent vs Incoherent Scatter

- **Coherent Scatter:** If the plasma is unstable and full of irregularities at the Bragg wavelength, lots of constructive interference will occur, and the radar will receive lots of signal.
- **Incoherent Scatter:** The plasma is disorganized:

$$\left| \sum_{p=0}^{N-1} e^{-j2\mathbf{k}_0 \cdot \Delta \mathbf{r}_p} \right|^2 \approx N$$

The pathological case where no scatter is received due to perfect destructive interference will almost surely never happen with a large number of electrons.

# Rough Detectability Calculations

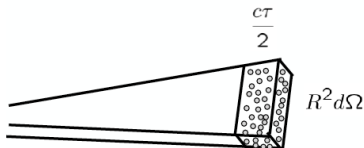
Radar Equation:

$$P_r = P_t \frac{G}{4\pi r^2} \sigma \frac{A_{eff}}{4\pi r^2}$$

For a distribution of electrons:

$$\sigma = \sigma_e N_e V \approx \sigma_e N_e r^2 \frac{c\tau_p}{2} \frac{4\pi}{G}$$

$$P_r \approx P_t \sigma_e N_e \frac{c\tau}{2} \frac{A_{eff}}{4\pi r^2}$$



Approximate beam solid angle:

$$d\Omega \approx \frac{4\pi}{G}$$

For  $P_t = 1$  MW,  $N_e = 10^{11} \text{ m}^{-3}$ ,  $\tau = 500 \mu\text{s}$ ,  $r = 300$  km,  $A_{eff} \approx 0.6 A_{geo}$ ,  $A_{geo} = \frac{\pi}{4} d^2$ , and a dish diameter of  $d = 300$  m, this gives:

$$P_r = 2.81 \times 10^{-14} \text{ W}$$

For a smaller radar with  $d = 30$  m,  $P_r = 2.81 \times 10^{-16} \text{ W}$

# Radio Noise

Nyquist Noise Theorem:  $P_N = k_B T_{sys} B$

- A good UHF receiver will have a  $T_{sys} \approx 125$  K.
- B is the receiver bandwidth.

Doppler shift from electron thermal motion:

$$\Delta f = \frac{2}{c} f_{Tx} v \approx \frac{2}{c} f_{Tx} \sqrt{\frac{k_B T_e}{m_e}}$$

Let's assume we need to capture  $B = 4\Delta f$  to get the full spectrum.

For  $f_{Tx} = 450$  MHz and  $T_e = 1000$  K:

$$B = 1.48 \text{ MHz} \Rightarrow P_N = 2.55 \times 10^{-15} \text{ W}$$

What if instead the bandwidth is related to the ion motion?

$$v_i = \sqrt{\frac{m_e}{m_i}} v_e \Rightarrow v_i = 5.83 \times 10^{-3} v_e \text{ for } O^+$$

The same numbers would yield

$$B = 8.63 \text{ kHz} \Rightarrow P_N = 1.48 \times 10^{-17} \text{ W}$$

# Autocorrelation Functions and Power Spectra

What determines the bandwidth of the received signal?

The electrons are moving, such that  $\Delta \mathbf{r}_p(t)$  is a function of time.

Autocorrelation function between scatter at two different times,  $t$  and  $t + \tau$ :

$$\begin{aligned} \langle E_s^*(t) E_s(t + \tau) \rangle &= \frac{r_e^2}{r^2} |E_0|^2 \left\langle \left[ \sum_{p=0}^{N-1} e^{+j2\mathbf{k}_0 \cdot \Delta \mathbf{r}_p(t)} \right] \left[ \sum_{q=0}^{N-1} e^{-j2\mathbf{k}_0 \cdot \Delta \mathbf{r}_q(t+\tau)} \right] \right\rangle \\ &= \frac{r_e^2}{r^2} |E_0|^2 \left\langle \sum_{p=0}^{N-1} \sum_{q=0}^{N-1} e^{j2\mathbf{k}_0 \cdot (\Delta \mathbf{r}_p(t) - \Delta \mathbf{r}_q(t+\tau))} \right\rangle \end{aligned}$$

Power spectrum of scatter:

$$\begin{aligned} \langle |E_s(\omega)|^2 \rangle &\equiv \int_{-\infty}^{\infty} \langle E_s^*(t) E_s(t + \tau) \rangle e^{-j\omega\tau} d\tau \\ &= \frac{r_e^2}{r^2} |E_0|^2 \langle |n_e(\mathbf{k}_b, \omega)|^2 \rangle \end{aligned}$$

# Simplest Possible Case

- The particles are non interacting
  - $\langle e^{j2\mathbf{k}_0 \cdot (\Delta\mathbf{r}_p(t) - \Delta\mathbf{r}_q(t+\tau))} \rangle = 0$  for  $p \neq q$ .
- The particles move in straight lines at constant velocities  $\mathbf{v}_p$  (positive away from radar)
  - $\Delta\mathbf{r}_p(t) = \Delta\mathbf{r}_p(0) + \mathbf{v}_p t$

$$\begin{aligned}
 \langle E_s^*(t) E_s(t + \tau) \rangle &= \frac{r_e^2}{r^2} |E_0|^2 \left\langle \sum_{p=0}^{N-1} \sum_{q=0}^{N-1} e^{j2\mathbf{k}_0 \cdot (\Delta\mathbf{r}_p(t) - \Delta\mathbf{r}_q(t+\tau))} \right\rangle \\
 &= \frac{r_e^2}{r^2} |E_0|^2 \left\langle \sum_{p=0}^{N-1} e^{j2\mathbf{k}_0 \cdot (\Delta\mathbf{r}_p(t) - \Delta\mathbf{r}_p(t+\tau))} \right\rangle \\
 &= \frac{r_e^2}{r^2} |E_0|^2 \left\langle \sum_{p=0}^{N-1} e^{j2\mathbf{k}_0 \cdot (-\mathbf{v}_p \tau)} \right\rangle \\
 &= \frac{r_e^2}{r^2} |E_0|^2 \int f(\mathbf{v}) e^{-j2\mathbf{k}_0 \cdot \mathbf{v} \tau} d^3v
 \end{aligned}$$

# Doppler Spectrum in Simplest Possible Case

$$\begin{aligned}
 \langle |E_s(\omega)|^2 \rangle &\equiv \int_{-\infty}^{\infty} \langle E_s^*(t) E_s(t+\tau) \rangle e^{-j\omega\tau} d\tau \\
 &= \frac{r_e^2}{r^2} |E_0|^2 \int_{-\infty}^{\infty} d\tau e^{-j\omega\tau} \int d^3v f(\mathbf{v}) e^{-j2\mathbf{k}_0 \cdot \mathbf{v}\tau} \\
 &= \frac{r_e^2}{r^2} |E_0|^2 \int d^3v f(\mathbf{v}) \delta(-\omega - 2\mathbf{k}_0 \cdot \mathbf{v}) \\
 &= \frac{r_e^2}{r^2} |E_0|^2 f\left(-\frac{\omega}{2k_0} \hat{k}\right)
 \end{aligned}$$

A particle moving at velocity  $\mathbf{v}$  backscatters with a Doppler shift of  $\Delta\omega = -2\mathbf{k}_0 \cdot \mathbf{v}$ .

Spectrum of received frequencies is directly related to spectrum of particle velocities.



# Particle Trajectories in Plasmas

Factors complicating particle trajectories

- Background magnetic field
- Collisions with neutrals
- Coulomb collisions with other particles
- Forces from the self-consistent fields generated by all other particles

The first three effects can be treated by deriving more complicated  $\Delta\mathbf{r}_p(t)$  for each particle.

The **collective** effects, however, are much more complicated to treat.

# Debye Length

At what scale are collective effects important?

- Characteristic particle velocity: Electron thermal speed

$$v_{te} = \sqrt{\frac{k_B T_e}{m_e}}$$

- Characteristic time scale for collective interactions: Inverse electron plasma frequency

$$\tau_e = \frac{1}{\omega_{pe}} = \sqrt{\frac{m_e \epsilon_0}{e^2 N_e}}$$

- Characteristic length scale: Debye length

$$\lambda_{De} = v_{te} \tau_e = \sqrt{\frac{\epsilon_0 k_B T_e}{e^2 N_e}}$$

Collective effects will have a significant affect on the particle trajectories over a Bragg wavelength if  $\lambda_b > \lambda_{De}$

# Particle-in-Cell Simulations

Electrostatic PIC equations:

$$m_s \frac{dv_s}{dt} = q_s \mathbf{E}$$

$$\mathbf{E} = -\nabla \phi$$

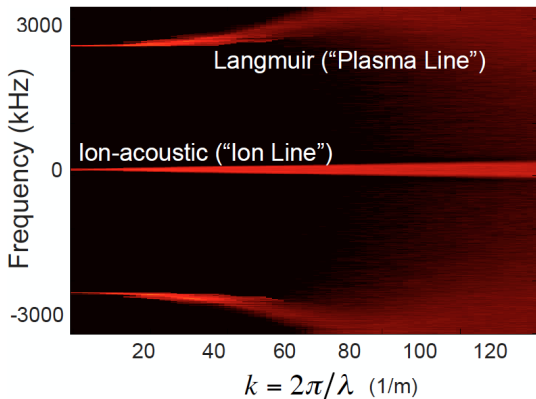
$$\nabla^2 \phi = -\frac{\rho_c}{\epsilon_0}$$

Diaz et al., RS [2008].

An ISR would pick out one slice of this spectrum at  $\mathbf{k} = \mathbf{k}_b$ .

Computed density fluctuation spectrum

$$\langle |n_e(\mathbf{k}, \omega)|^2 \rangle$$



# Dressed Particle Theory

- Imagine a set of “test particles” that move without experiencing collective effects.
- As the charged test particles move they create macroscopic disturbances in the rest of the plasma.

$$\left\langle |n_e(\mathbf{k}, \omega)|^2 \right\rangle \approx \left\langle |n_{te}(\mathbf{k}, \omega) + \delta N_e(\mathbf{k}, \omega)|^2 \right\rangle$$

Where

- $n_{te}$  is the microscopic density function of the test electrons
- $\delta N_e$  are the macroscopic density fluctuations in response to the test particles (ions and electrons)

# Plasma Response Functions

Plasma as a generalized conductor:

$$\mathbf{J}_e = \sigma_e(\mathbf{k}, \omega) \mathbf{E}$$

Plasma as a generalized dielectric:

$$\mathbf{P}_e = \epsilon_0 \chi_e(\mathbf{k}, \omega) \mathbf{E}$$

Bound charge density and electron density disturbance

$$\rho_c = -\nabla \cdot \mathbf{P} \rightarrow \delta N_e(\mathbf{k}, \omega) = +\frac{\epsilon_0}{e} \nabla \cdot [\chi_e(\mathbf{k}, \omega) \mathbf{E}]$$

These are equivalent descriptions

$$\mathbf{J}_e = \frac{\partial}{\partial t} \mathbf{P}_e \rightarrow \chi_e(\mathbf{k}, \omega) = \frac{\sigma_e(\mathbf{k}, \omega)}{j\omega\epsilon_0}$$

# Gauss' Law

$$\epsilon_0 \nabla \cdot \mathbf{E} = \rho_c$$

$$\epsilon_0 \nabla \cdot \mathbf{E} = -\nabla \cdot (\mathbf{P}_i + \mathbf{P}_e) + en_{ti} - en_{te}$$

$$\epsilon_0 \nabla \cdot \mathbf{E} = -\epsilon_0 \nabla \cdot [(\chi_i + \chi_e) \mathbf{E}] + en_{ti} - en_{te}$$

$$\epsilon_0 \nabla \cdot [(1 + \chi_i + \chi_e) \mathbf{E}] = en_{ti} - en_{te}$$

$$-j\epsilon_0 (1 + \chi_i + \chi_e) \mathbf{k} \cdot \mathbf{E} = en_{ti} - en_{te}$$

$$\mathbf{k} \cdot \mathbf{E} = \frac{en_{ti} - en_{te}}{-j\epsilon_0 (1 + \chi_i + \chi_e)}$$

Dielectric permittivity of the plasma:

$$\epsilon = \epsilon_0 (1 + \chi_i + \chi_e)$$

# Effective electron density fluctuations

$$\begin{aligned}
 n_e &= n_{te} + \delta N_e \\
 &= n_{te} + \frac{\epsilon_0}{e} \nabla \cdot [\chi_e \mathbf{E}] \\
 &= n_{te} - j \frac{\epsilon_0}{e} \chi_e \mathbf{k} \cdot \mathbf{E} \\
 &= \frac{1 + \chi_i}{1 + \chi_i + \chi_e} n_{te} + \frac{\chi_e}{1 + \chi_i + \chi_e} n_{ti}
 \end{aligned}$$

Power spectrum (note  $n_{te}$  and  $n_{ti}$  are uncorrelated)

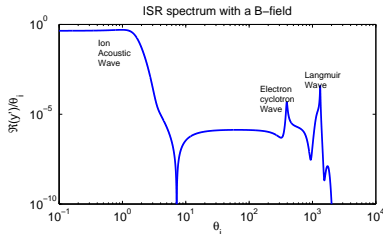
$$\begin{aligned}
 \langle |n_e(\mathbf{k}, \omega)|^2 \rangle &= \frac{|1 + \chi_i|^2}{|1 + \chi_i + \chi_e|^2} \langle |n_{te}(\mathbf{k}, \omega)|^2 \rangle \\
 &\quad + \frac{|\chi_e|^2}{|1 + \chi_i + \chi_e|^2} \langle |n_{ti}(\mathbf{k}, \omega)|^2 \rangle
 \end{aligned}$$

# Connection to Dispersion Relation for Electrostatic Waves

- If there were no test particle to drive fluctuations, Gauss' Law of the macroscopic plasma is

$$\epsilon_0 \nabla \cdot [(1 + \chi_i + \chi_e) \mathbf{E}] = 0$$

- If  $1 + \chi_i(\mathbf{k}, \omega) + \chi_e(\mathbf{k}, \omega) = 0$ , then  $\mathbf{E}$  could be anything.
- The set of  $\mathbf{k}$  and  $\omega$  that satisfy  $1 + \chi_i(\mathbf{k}, \omega) + \chi_e(\mathbf{k}, \omega) = 0$  define the **normal modes** in the plasma.
- Near a normal mode, the denominator of the ISR spectrum is nearly 0, so the spectrum has a peak.



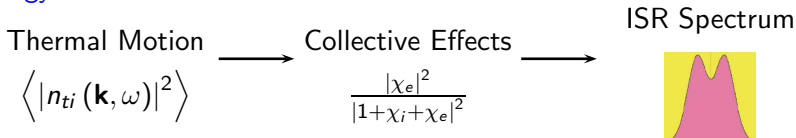


# Filtered Noise

Can you hear the ocean in a seashell?



Analogy to ISR:



# Linearized Vlasov Equations

Vlasov equation:

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \frac{\partial f}{\partial \mathbf{r}} + \frac{q}{m} \mathbf{E} \cdot \frac{\partial f}{\partial \mathbf{v}}$$

Linearization assumptions:  $f \approx f_0 + f_1 e^{j\omega t - j\mathbf{k} \cdot \mathbf{r}}$

$$j\omega f_1 - j\mathbf{v} \cdot \mathbf{k} f_1 + \frac{q}{m} \frac{\partial f_0}{\partial \mathbf{v}} \cdot \mathbf{E} = 0$$

$$f_1 = -\frac{q}{m j (\omega - \mathbf{k} \cdot \mathbf{v})} \frac{\partial f_0}{\partial \mathbf{v}} \cdot \mathbf{E}$$

Electron density perturbation:

$$\begin{aligned} \delta N &= \int f_1 d^3 v = -\frac{q}{m} \int d^3 v \frac{\partial f_0}{\partial \mathbf{v}} \cdot \mathbf{E} \\ &= j \frac{\epsilon_0}{q} \chi \mathbf{k} \cdot \mathbf{E} \end{aligned}$$

# Landau Problem

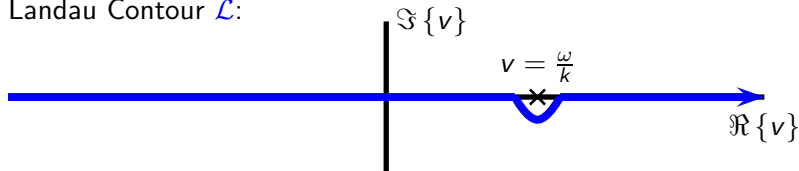
Susceptibility integral has singularity where  $\omega = \mathbf{k} \cdot \mathbf{v}$

$$\chi(\mathbf{k}, \omega) = \frac{q^2}{k^2 m \epsilon_0} \int d^3 v \frac{\mathbf{k} \cdot \frac{\partial f_0}{\partial \mathbf{v}}}{\omega - \mathbf{k} \cdot \mathbf{v}}$$

Solution: Treat the problem with Laplace transforms (Landau [1946])

$$\begin{aligned} \chi(\mathbf{k}, \omega) &= \frac{q^2}{k^2 m \epsilon_0} \int_{\mathcal{L}} d^3 v \frac{\mathbf{k} \cdot \frac{\partial f_0}{\partial \mathbf{v}}}{\omega - \mathbf{k} \cdot \mathbf{v}} \\ &= \frac{q^2}{k^2 m \epsilon_0} \left\{ \mathcal{P} \int d^3 v \frac{\mathbf{k} \cdot \frac{\partial f_0}{\partial \mathbf{v}}}{\omega - \mathbf{k} \cdot \mathbf{v}} + i\pi \mathbf{k} \cdot \frac{\partial f_0}{\partial \mathbf{v}} \Big|_{\mathbf{v}=\frac{\omega}{k} \hat{\mathbf{k}}} \right\} \end{aligned}$$

Landau Contour  $\mathcal{L}$ :



# Relationship of Spectrum to Density and Temperature

Everything is a function of the zeroth-order distribution functions

- Susceptibilities:

$$\chi(\mathbf{k}, \omega) = \frac{q^2}{k^2 m \epsilon_0} \int_{\mathcal{L}} d^3 v \frac{\mathbf{k} \cdot \frac{\partial f_0}{\partial \mathbf{v}}}{(\omega - \mathbf{k} \cdot \mathbf{v})}$$

- Test particle spectra:

$$\left\langle |n_t(\mathbf{k}, \omega)|^2 \right\rangle = f_0 \left( -\frac{\omega}{k} \hat{k} \right)$$

If you assume  $f_0$  is Maxwellian, then  $f_0$  is a function of density, bulk velocity, and temperature. Thus the whole ISR spectrum can be written a function of  $N_e$ ,  $T_e$ ,  $T_i$ , and  $\mathbf{u}$ .

# Total Scattered Power

$$\begin{aligned} \langle |n_e(\mathbf{k}, \omega)|^2 \rangle &= \frac{|1 + \chi_i|^2}{|1 + \chi_i + \chi_e|^2} \langle |n_{te}(\mathbf{k}, \omega)|^2 \rangle && \text{Electron Line} \\ &+ \frac{|\chi_e|^2}{|1 + \chi_i + \chi_e|^2} \langle |n_{ti}(\mathbf{k}, \omega)|^2 \rangle && \text{Ion Line} \end{aligned}$$

Area under spectrum

$$\int \langle |n_e(\mathbf{k}, \omega)|^2 \rangle \frac{d\omega}{2\pi} = \frac{k^2 \lambda_{De}^2 N_e}{1 + k^2 \lambda_{De}^2} + \frac{N_e}{(1 + k^2 \lambda_{De}^2) \left(1 + k^2 \lambda_{De}^2 + \frac{T_e}{T_i}\right)}$$

In the limit  $k^2 \lambda_{De}^2 \gg 1$ , electron line dominates (**wide bandwidth**)

In the limit  $k^2 \lambda_{De}^2 \ll 1$ , the ion line dominates (**narrow bandwidth**)

$$\int \langle |n_e(\mathbf{k}, \omega)|^2 \rangle \frac{d\omega}{2\pi} = N_e$$

$$\int \langle |n_e(\mathbf{k}, \omega)|^2 \rangle \frac{d\omega}{2\pi} = \frac{N_e}{1 + \frac{T_e}{T_i}}$$

# Big Takeaways

- Radars only scatter from electrons
- Bragg scattering effectively picks out a single wavenumber
- Even in a homogeneous, stable plasma, the scatter is never zero
- Ions encode information into the spectrum through collective effects
- ISR theory is intimately connected to dispersion relationships for electrostatic plasma waves
- Plasma normal modes  $\Rightarrow$  peaks in spectrum