Observing the Auroral Oval Group 2



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What to Investigate?



Experiment Setup

- Experiment time: 18-20 UT
- ESR (MLT=UT+2:45)
 - 42m: along B (el 81.6)
 - 32m: toward geomagnetic south, el 30
- EISCAT Tromso (MLT=UT+2:30)
 - UHF: along B
 - VHF: toward geomagnetic north, el 30





ACE data (SWPC)





Global Magnetic indices

Plot of Kp, Dst, AE indices from 2019-08-13 to 2019-08-14



http://isgi.unistra.fr/

H-component magnetic field data for August 13th 2019.



- Ø Interval of radar observation (~ 18 20 UT).
- Ø A decrease ~15:45 and ~16:45 (high lat. Stns), a likely enhanced westward electrojet a substorm signature.
- Ø Possible Substorm-enhanced particle population associated with the burst 1 hr before the time of radar

A comparison of H-component magnetic field and AE data for August 13th, 2019.



Ø Variation of AE index confirms a possible magnetic activity

during the interval of radar observation (~ 18 – 20 UT)

Ovation Predictive Modeling

Ovation predicts total energy and number flux, and average energy of precipitating polar-region ions and electrons.

<u>08/13/2019, 1800-2000 UT:</u> No G1, R1, S1 (Minor) or greater geomagnetic storms, transients or recurrent solar wind features expected.



Seasonal Trends vs Acute Space Weather

The NOAA SWPC database does not store its predictive Nowcast results; NASA OMNIweb level 2 radar measurements are published with significant latency

Model accuracy requires consideration of:

- Geomagnetic storm hysteresis
- IMF conditions
- F10.7

Six year (2013-2018) electron energy flux summary is shown to the right





F10.7cm data does not seem indicate relatively higher electron energy fluxes in 2017 vs 2015





Credit: NOAA Space Weather Prediction Center (SWPC)



Seasonal Trends vs Acute Space Weather

EISCAT data





Field aligned
No precipitation signatures!





• Low elevation south





- Precipitation reaches low altitudes
- Still moving southward





- Auroral precipitation near the radar site
- Moving southward





32m VHF TRO

- Periodic features in ion temperature (Ti)
- ~ 5 10 min periodicity
- Enhancement of ~ 2000 K

TIDs or ULF, or aurora streamers ?



32m VHF TRO - aurora streamers?





lan's presentation - large structures

32m VHF TRO - TIDs?



Borries et al., 2017



20 November 2003

UHF, 32m, pointing along B **EISCAT Scientific Association EISCAT UHF RADAR** CP, uhfa, beata, 13 August 2019 DLAPTOP-H7OVKQLA, 16-Aug-2019 Not for publication - see Rules-of-the-road 400 Altitude (km) 300 km 18:00 19:30 19:45 20:00 3000 Mittude (km) 000 2000 1000툲 Altitude (F 18.00 19.00 19:45 20:00 <u>(ال</u> 300 لغ ji 200 1000 ត្តី 20:00 de (km) (m) 200 18:30 19:00 19:30 20:00 18:00 Universal time

32m UHF TRO

- Nothing particular
- F-region/ E-region in Ne

Conclusion

- Relatively quiet solar conditions small increase in sw velocity and a substorm at our observations time
- Detection of the auroral oval in ESR 32m pointing southward
- Observations of ionospheric features in Tromsø 32m VHF (TIDs, ULF waves or aurora streamers ?)



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