

# Phased Arrays for Atmospheric and Geospace Science

Roger H. Varney

Center for Geospace Studies  
SRI International

July 28, 2017

- 1 Phased Array Fundamentals
- 2 Passive and Active Arrays
- 3 Interferometry
- 4 Digital Beam Forming

# Superposition Principle

Maxwell's Equations are Linear:

$$\mathbf{J}_1 = \frac{1}{\mu_0} \nabla \times (\mathbf{B}_1) - \epsilon_0 \frac{\partial}{\partial t} (\mathbf{E}_1)$$

$$0 = \nabla \times (\mathbf{E}_1) + \frac{\partial}{\partial t} (\mathbf{B}_1)$$

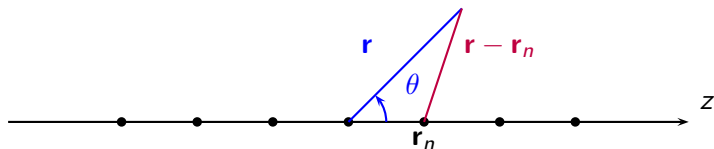
$$\mathbf{J}_2 = \frac{1}{\mu_0} \nabla \times (\mathbf{B}_2) - \epsilon_0 \frac{\partial}{\partial t} (\mathbf{E}_2)$$

$$0 = \nabla \times (\mathbf{E}_2) + \frac{\partial}{\partial t} (\mathbf{B}_2)$$

$$\mathbf{J}_1 + \mathbf{J}_2 = \frac{1}{\mu_0} \nabla \times (\mathbf{B}_1 + \mathbf{B}_2) - \epsilon_0 \frac{\partial}{\partial t} (\mathbf{E}_1 + \mathbf{E}_2)$$

$$0 = \nabla \times (\mathbf{E}_1 + \mathbf{E}_2) + \frac{\partial}{\partial t} (\mathbf{B}_1 + \mathbf{B}_2)$$

# Superposition Applied to Antenna Arrays



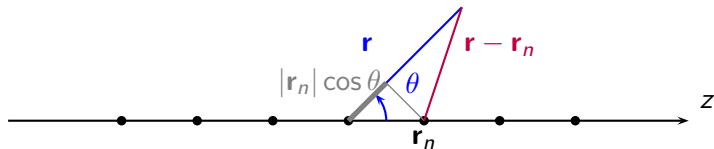
Fields radiated by single element at the origin with applied current  $I_0$ :

$$\mathbf{E} = \mathbf{E}_0 I_0 \frac{e^{-jk|\mathbf{r}|}}{|\mathbf{r}|}$$

Fields radiated by entire array:

$$\mathbf{E} = \mathbf{E}_0 \sum_{n=0}^{N-1} I_n \frac{e^{-jk|\mathbf{r} - \mathbf{r}_n|}}{|\mathbf{r} - \mathbf{r}_n|}$$

# Far Field Approximation (Fraunhofer Zone)



If  $\mathbf{r}$  and  $\mathbf{r} - \mathbf{r}_n$  are almost parallel lines:

$$\mathbf{r} - \mathbf{r}_n \approx \mathbf{r} - |\mathbf{r}_n| \cos \theta \hat{\mathbf{r}}$$

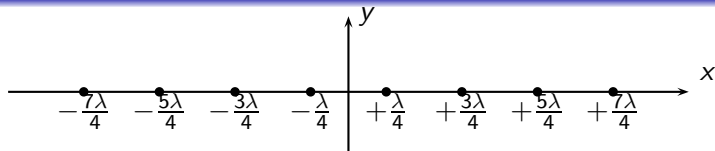
Assume  $|\mathbf{r}_n| \ll |\mathbf{r}|$ :

$$|\mathbf{r} - \mathbf{r}_n| \approx |\mathbf{r}| \text{ for demoninator terms}$$

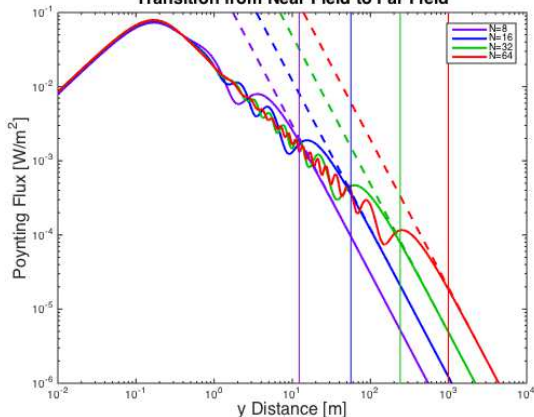
$$-jk |\mathbf{r} - \mathbf{r}_n| \approx -jk |\mathbf{r}| + jk |\mathbf{r}_n| \cos \theta$$

$$\mathbf{E} \approx \underbrace{\mathbf{E}_0 \frac{e^{-jk|\mathbf{r}|}}{|\mathbf{r}|}}_{\text{Element Factor}} \underbrace{\sum_{n=0}^{N-1} I_n e^{jk|\mathbf{r}_n| \cos \theta}}_{\text{Array Factor}}$$

# Distance to Far Field: Fresnel Numbers



Transition from Near Field to Far Field



Transition from near to far determined by the **Fresnel Number**:

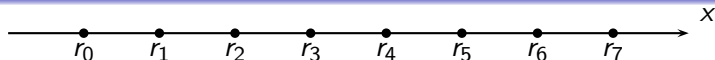
$$\frac{L^2}{r\lambda} \ll 1 \rightarrow \text{Far Field}$$

$$\frac{L^2}{r\lambda} > 1 \rightarrow \text{Near Field}$$

$L$  = Array length

$\lambda$  = wavelength

# 1-D Linear Phased Array



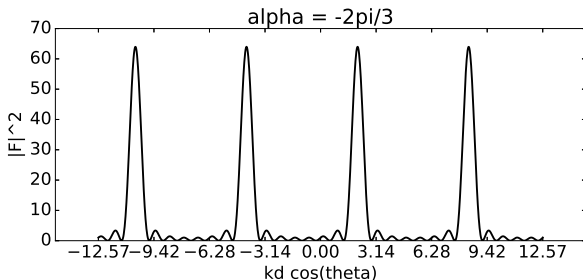
$$|\mathbf{r}_n| = nd \quad I_n = e^{jn\alpha}$$

Array Factor:

$$\begin{aligned}
 F &= \sum_{n=0}^{N-1} e^{jn\alpha} e^{jkn d \cos \theta} \\
 &= \frac{1 - e^{jN\alpha + jNkd \cos \theta}}{1 - e^{j\alpha + jkd \cos \theta}} \\
 &= e^{j\frac{(N-1)}{2}(kd \cos \theta + \alpha)} \frac{\sin \left[ \frac{N}{2} (kd \cos \theta + \alpha) \right]}{\sin \left[ \frac{1}{2} (kd \cos \theta + \alpha) \right]} \\
 |F|^2 &= \frac{\sin^2 \left[ \frac{N}{2} (kd \cos \theta + \alpha) \right]}{\sin^2 \left[ \frac{1}{2} (kd \cos \theta + \alpha) \right]}
 \end{aligned}$$

# 1-D Linear Phased Array Cont.

$$|F|^2 = \frac{\sin^2 \left[ \frac{N}{2} (kd \cos \theta + \alpha) \right]}{\sin^2 \left[ \frac{1}{2} (kd \cos \theta + \alpha) \right]}$$



Peak appears when  $kd \cos \theta = -\alpha$

Additional peaks could appear when  $kd \cos \theta = -\alpha + 2\pi m$  (Grating Lobes)

**Visible Region:**  $0 < \theta < \pi \rightarrow -kd < kd \cos \theta < kd$



# Grating Lobes

- $d < \lambda/2 \rightarrow kd < \pi$ : No grating lobes will ever appear
- $\lambda/2 < kd < \lambda \rightarrow \pi < kd < 2\pi$ : Grating lobes will only appear at some steering angles
- $d > \lambda \rightarrow kd > 2\pi$ : Grating lobes will always appear

Example of linear array with  $d = 2\lambda/3$  spacing

Movie

# Mutual Coupling

- The true element factor for antennas in an array is different from the same type of antennas in isolation
  - Scattering off of neighboring antennas
  - Inductive coupling involving antenna near-fields

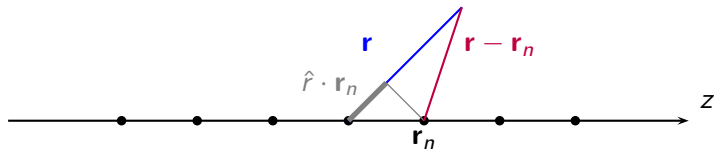
Two possible solutions

- Use a larger antenna separation and live with a limited grating-lobe free steering range.
- Use specially designed antennas to minimize coupling

EISCAT\_3D Prototype Drooped Dipole →



# Multi-Dimensional Arrays



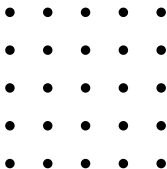
$$-jk |\mathbf{r} - \mathbf{r}_n| \approx -jk |\mathbf{r}| + jk (\hat{\mathbf{r}} \cdot \mathbf{r}_n)$$

$$\mathbf{E} \approx \underbrace{\mathbf{E}_0 \frac{e^{-jk|\mathbf{r}|}}{|\mathbf{r}|}}_{\text{Element Factor}} \underbrace{\sum_{n=0}^{N-1} I_n e^{jk(\hat{\mathbf{r}} \cdot \mathbf{r}_n)}}_{\text{Array Factor}}$$

In spherical coordinates:

$$\hat{\mathbf{r}} \cdot \mathbf{r}_n = x_n \cos \phi \sin \theta + y_n \sin \phi \sin \theta + z_n \cos \theta$$

## 2-D Rectangular Array

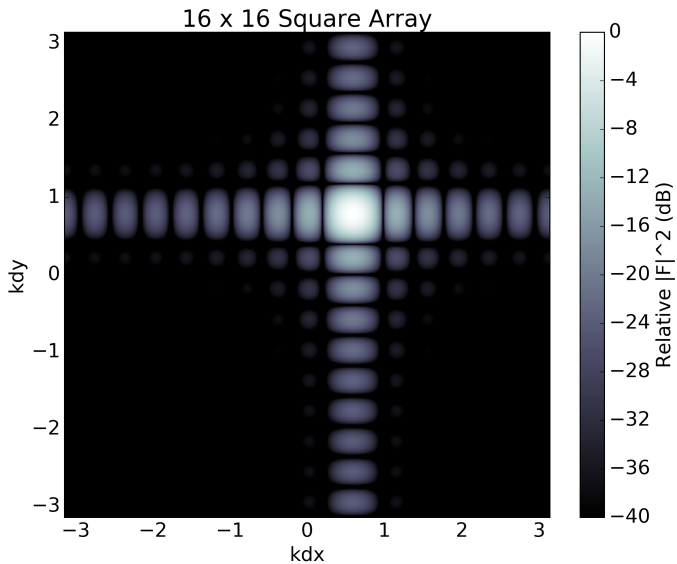


$$\mathbf{r}_{nm} = nd_x \hat{\mathbf{x}} + md_y \hat{\mathbf{y}} \quad I_{nm} = e^{j(n\alpha + m\beta)}$$

Array Factor:

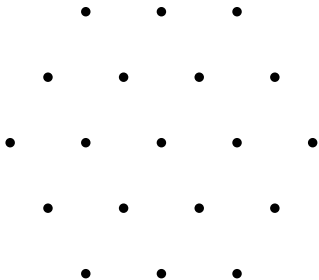
$$\begin{aligned}
 |F(\theta, \phi)|^2 &= \left| \sum_{n=0}^{N_x-1} \sum_{m=0}^{N_y-1} e^{j(nkd_x \cos \phi \sin \theta + n\alpha + mkd_y \sin \phi \sin \theta + m\beta)} \right|^2 \\
 &= \frac{\sin^2 \left[ \frac{N_x}{2} (kd_x \cos \phi \sin \theta + \alpha) \right]}{\sin^2 \left[ \frac{1}{2} (kd_x \cos \phi \sin \theta + \alpha) \right]} \frac{\sin^2 \left[ \frac{N_y}{2} (kd_y \sin \phi \sin \theta + \beta) \right]}{\sin^2 \left[ \frac{1}{2} (kd_y \sin \phi \sin \theta + \beta) \right]}
 \end{aligned}$$

# 2-D Rectangular Array

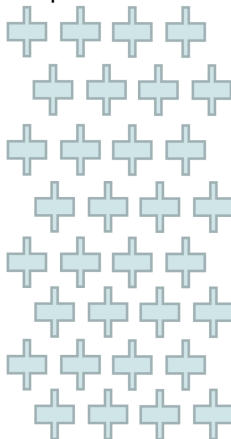


# Hexagonal Spacing

Hexagon



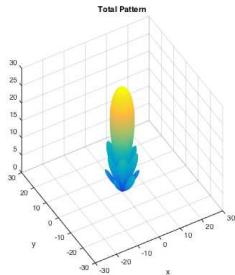
Honeycomb Rectangular Array  
One AMISR panel:



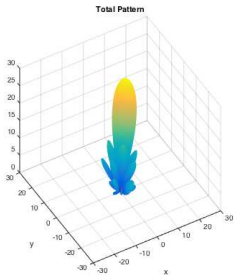
# Steering A Hexagonal Array

19-antenna hexagon with  $d = 3\lambda/4$ .

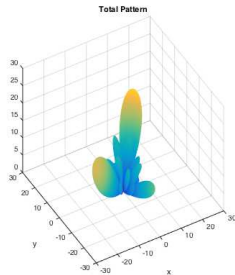
$$\theta = 0^\circ, \phi = 0^\circ$$



$$\theta = 20^\circ, \phi = 45^\circ$$

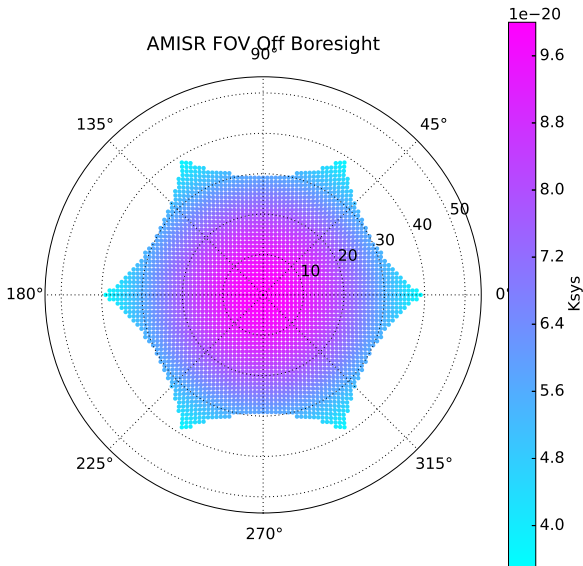


$$\theta = 45^\circ, \phi = 45^\circ$$



# AMISR Antenna Properties

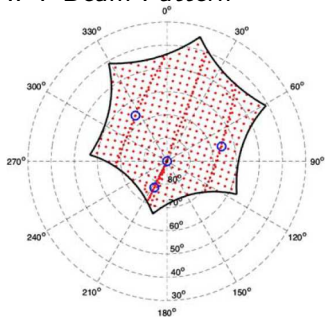
- FOV limited by grating lobe limit  $\sim 30^\circ - 40^\circ$
- Antenna gain decreases with steering angle off of boresight
- Antenna works best within  $\sim 25^\circ$  off of boresight





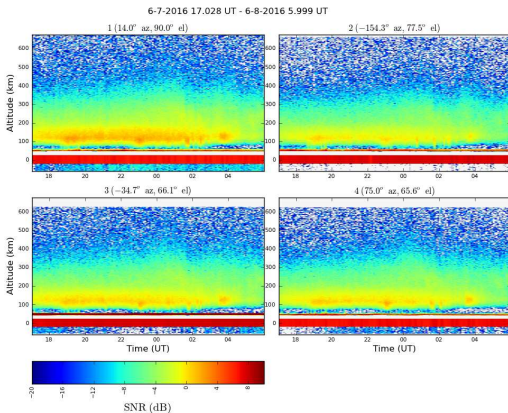
# The PFISR Up-B Compromise

## IPY Beam Pattern



The Up-B beam is close to the grating lobe limit, and therefore has reduced sensitivity.

## Reduced SNR in Up-B (Beam 2)



# Passive Phased Arrays: Jicamarca

- One transmitter feeds entire array



# Passive Phased Arrays: Jicamarca



# Manual Phasing (Jicamarca)



# Active Electronically Steerable Phased Arrays

## The AMISR UHF System

### AMISR AEU = Tx/Rx Unit

- 500 W solid state transmitter
- Phasing control
- Status monitoring
- 4096 AEU's/AMISR radar face

Antenna Element Unit (AEU)



32/panel

### AMISR Panel

- 32 Antenna Element Units arranged in hexagonal pattern
- 3.5 x 2 meters; 19.8 dBi / panel
- 16 kW peak power per panel
- Basic system building block for AMISR
- Embedded linux controller



Panel (with PCU)

Face



Utility Distribution Unit (UDU)



AMISR Control System (ACS)

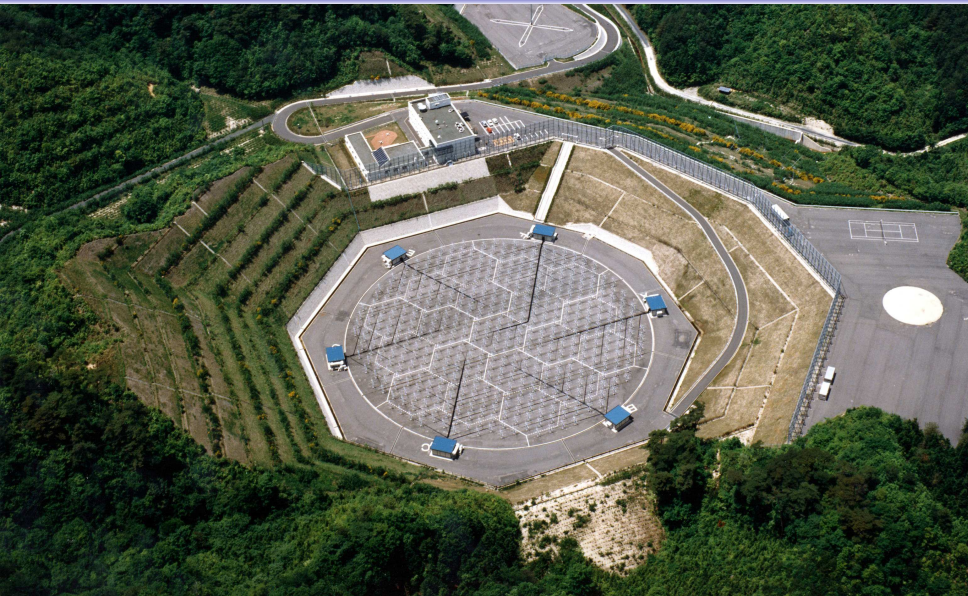
### AMISR UDU

- 400 Hz JetPower converters
- Remote power control units
- Fiber distribution system

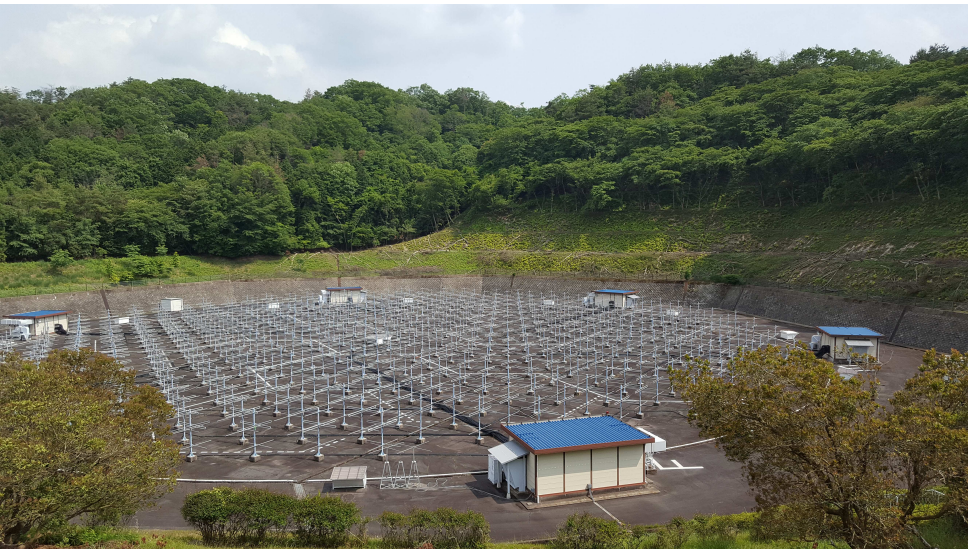
### AMISR ACS

- Flexible transmit and receive system
- Completely remotely controlled
- Experiments run off a scheduler

# MU Radar



# MU Radar



# MU Radar Cabling





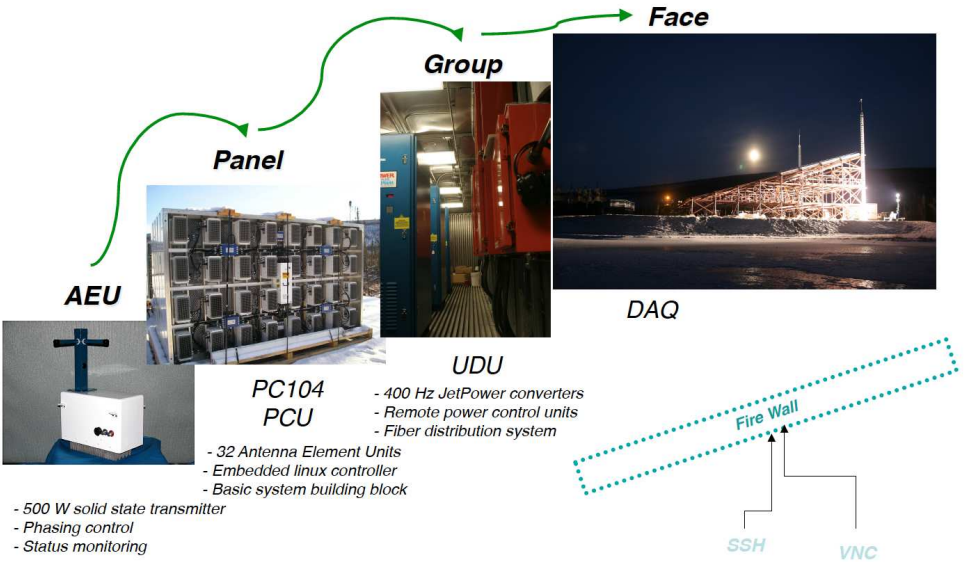
# MU Radar Cabling



# MU Radar Power Amplifiers



# Advanced Modular Incoherent Scatter Radar



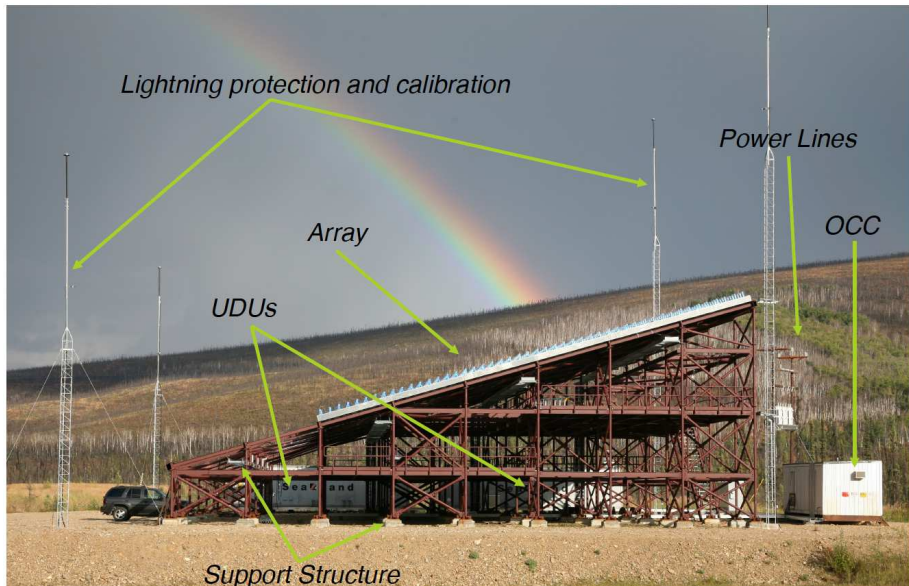
# Antenna Element Unit (AEU) Specifications

- **Distributed Solid State Power Amplifiers (SSPAs)**
- **430-450 MHz instantaneous bandwidth**
- **10% Maximum duty cycle**
- **Minimum PRF interval 500 usec**
- **Maximum pulsewidth 2 msec**
- **Passive cooling (no moving parts)**
- **400 Hz prime power**



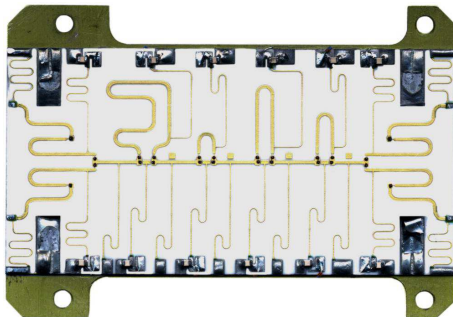
- **Crossed dipoles, circular polarization on axis**
- **Balun built into the antenna support shaft**
- **Constant impedance over bandwidth and scan angle**
- **Spacing is hexagonal for efficiency**
- **Tx/Rx polarizations are opposite and fixed (not measurable)**

# Poker Flat Incoherent Scatter Radar (PFISR)



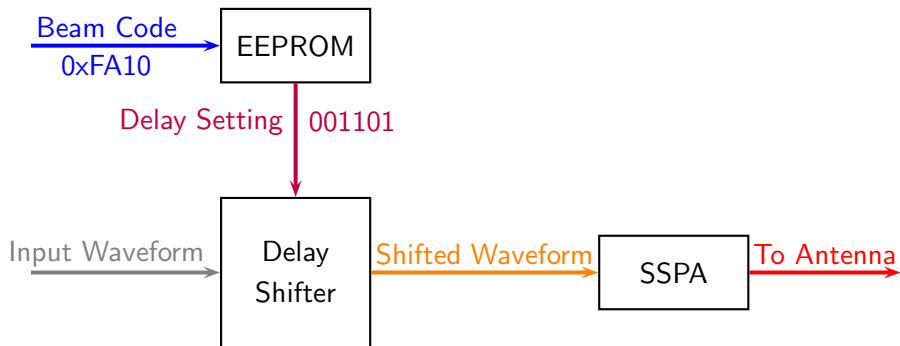
# Electronic Steering with Delay Shifters

Example 4-bit delay shifter:



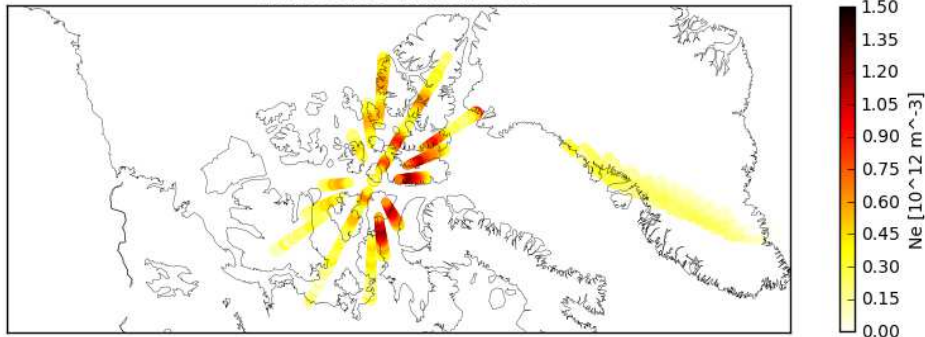
- AMISR uses 6-bit delay shifters
- $2^6 = 64$  steps spaced by  $\pi/32 = 5.625^\circ$

# Conceptual Diagram of Steering with AMISR



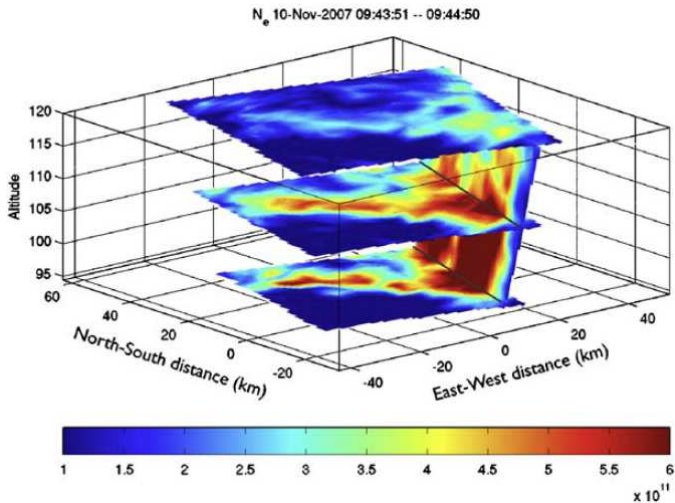
# Differences Between AMISR and Scanning Radars

2016-Oct-13 16:59:24 UTC

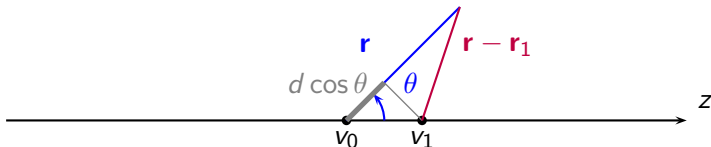




# Imaging Auroral Structure [Semeter et al. (2009)]



# Two Antenna Interferometry



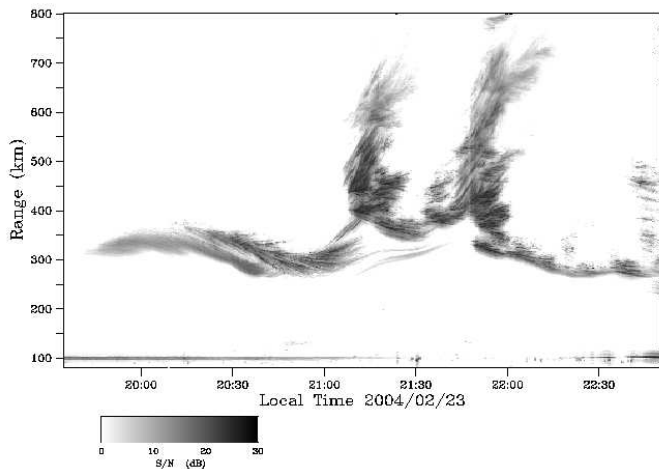
$$\langle v_0 v_1^* \rangle = e^{jkd \cos \theta}$$

- Measure  $v_0$  and  $v_1$  separately and estimate the angle of arrival  $\theta$ .
- Larger baseline  $\rightarrow$  more precise angle estimates.
- Baselines where  $kd > \pi$  suffer from  $2\pi$  ambiguity issues. Related to grating lobe problem.

# Interferometric Imaging Configuration at Jicamarca



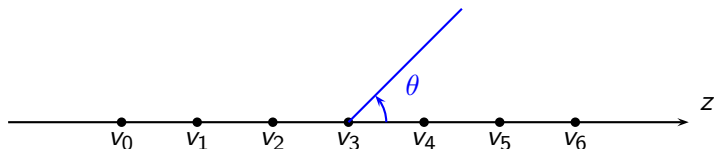
# Interferometric Images of Coherent Scatter



Movie

<http://landau.geo.cornell.edu/image.html>

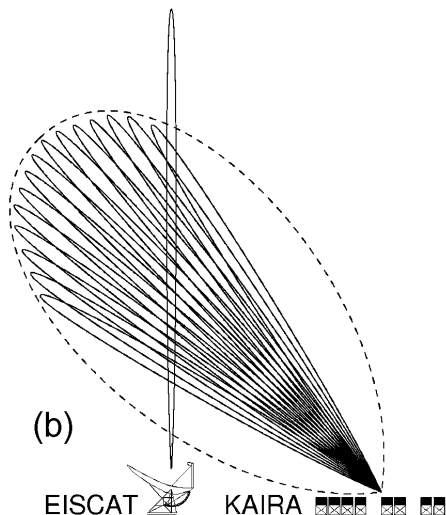
# Digital Beam Forming



On reception:

- Digitize the signals  $v_n$  on every antenna (**expensive!**)
- Synthesize any beams you want by forming different linear combinations in software/firmware (**computationally intense!**)
- Allows you to form custom beam patterns  $\rightarrow$  look at signals of interest while nulling interference.
- Allows you to form any number of different radiation patterns  $\rightarrow$  look in multiple directions at once.

# Digital Beam Forming in Multi-static Radar Experiments

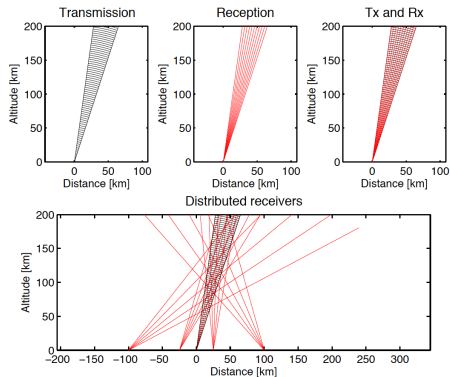
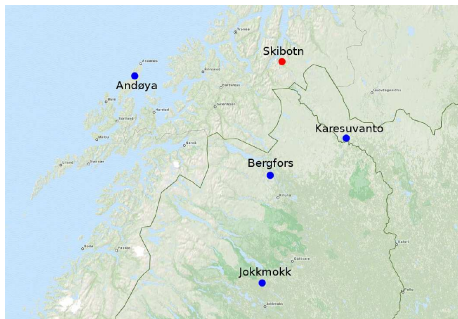


KAIRA = The Kilpisjärvi  
Atmospheric Imaging  
Receiver Array

McKay et al. (2015)  
10.1109/TGRS.2014.2342252

## EISCAT\_3D

Major planned facility:



Operational 2021?