



EISCAT_3D

Craig Heinselman

EISCAT Scientific Association



EISCAT_3D Design Study

- 5 partners, 30 man years
 - EISCAT, University of Tromsø, Luleå University of Technology, Rutherford Appleton Laboratory, Swedish Institute of Space Physics

- TI budgeted volume 2.8 MEUR

- EU FP6 support 2 ME

WP1: Project Management

WP2: Evaluation of design performance goals

WP3: Evaluation of options for the active element

WP4: Phased array receivers

WP5: Interferometric receivers

WP6: Active element

WP7: Distributed control and monitoring and Observation scheme

WP8: Data Archiving and Distribution

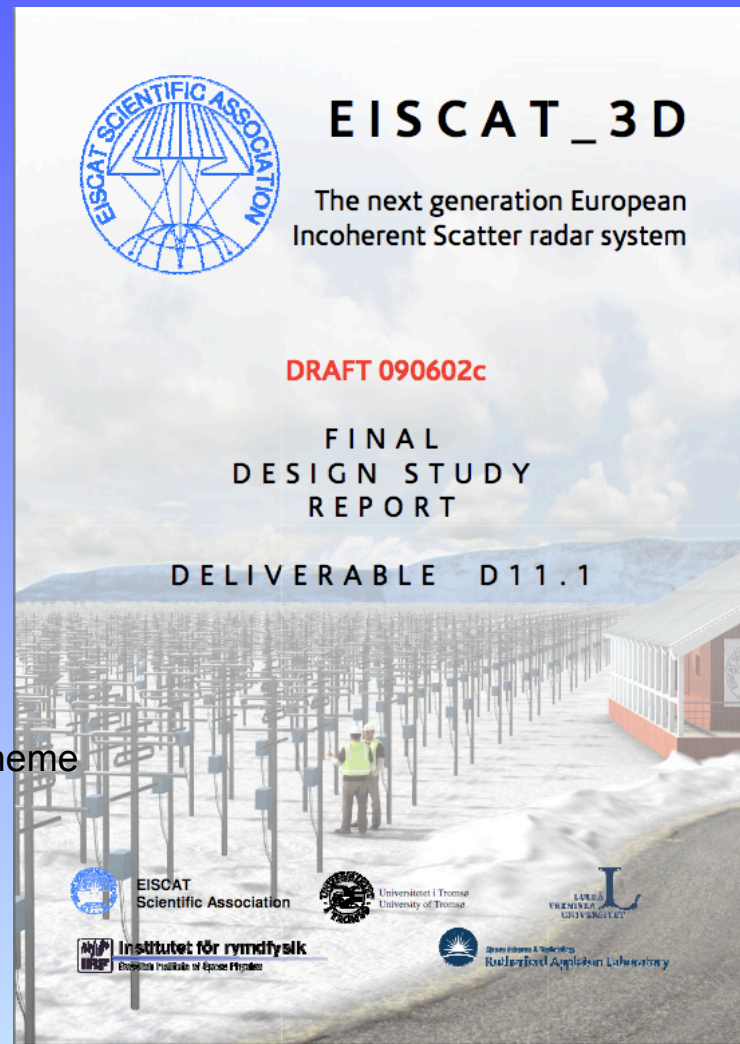
WP9: Signal Processing

WP10: New uses

WP11: Implementation Blueprint

WP12: Time and frequency distribution

WP13: Enabling procedures





EISCAT_3D Preparatory Phase

14 work packages:

- WP1: Management and reporting
- WP2: Legal and logistical issues
- WP3: Science planning
- WP4: Outreach activities
- WP5: Consortium building
- WP6: Performance specification
- WP7: Signal processing
- WP8: Antenna, front end and timing
- WP9: Transmitter development
- WP10: Aperture synthesis imaging
- WP11: Software theory & implementation
- WP12: System control
- WP13: Data handling & distribution
- WP14: Mass-production & reliability



EISCAT_3D

A European Three-Dimensional Imaging Radar for
Atmospheric and Geospace Research

ESFRI Roadmap Project

State of the art today: AMISR



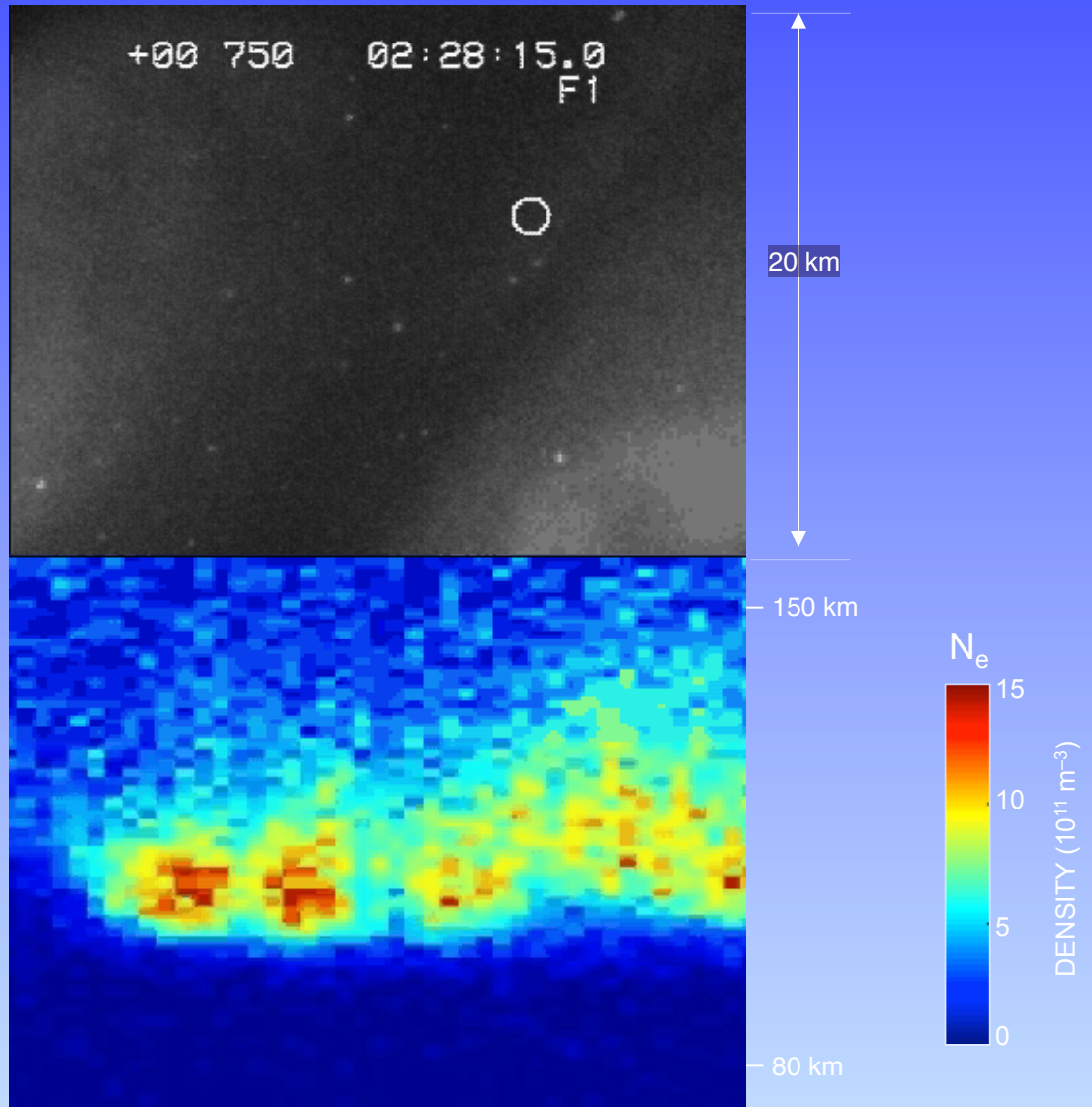
Resolving Temporal/Spatial Ambiguities

High Speed Intensified Auroral Imaging

- **Narrow-field camera**
 - 25 frames/sec
 - > 640 nm
 - 21 x 26 km at 110 km

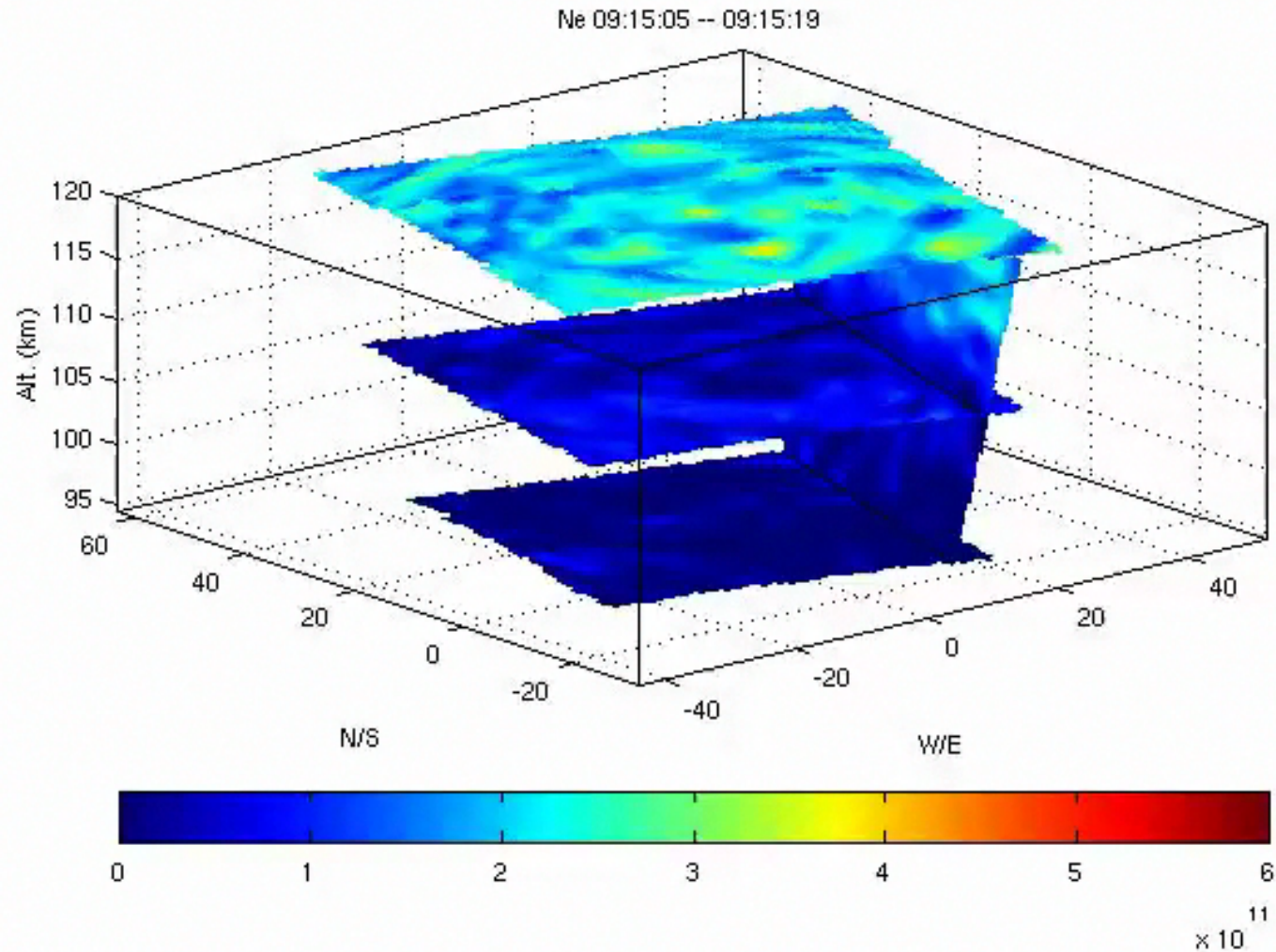
- **Sondrestrom IS radar**
 - Electron density
 - 1 km x 1.2 sec

J. Semeter





AMISR view of an aurora





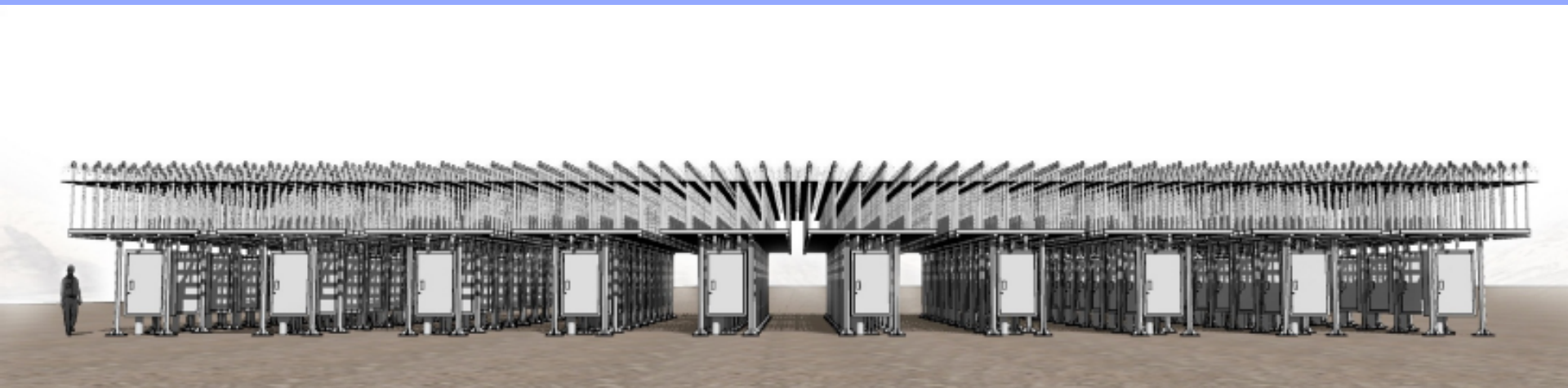
EISCAT_3D

EISCAT_3D will be a volumetric vector-imaging radar for studying the geospace environment

It represents a revolutionary upgrade to the existing EISCAT mainland facilities, utilizing multi-static, phased-array technologies

It will support continuous measurements of the space environment via unattended operations

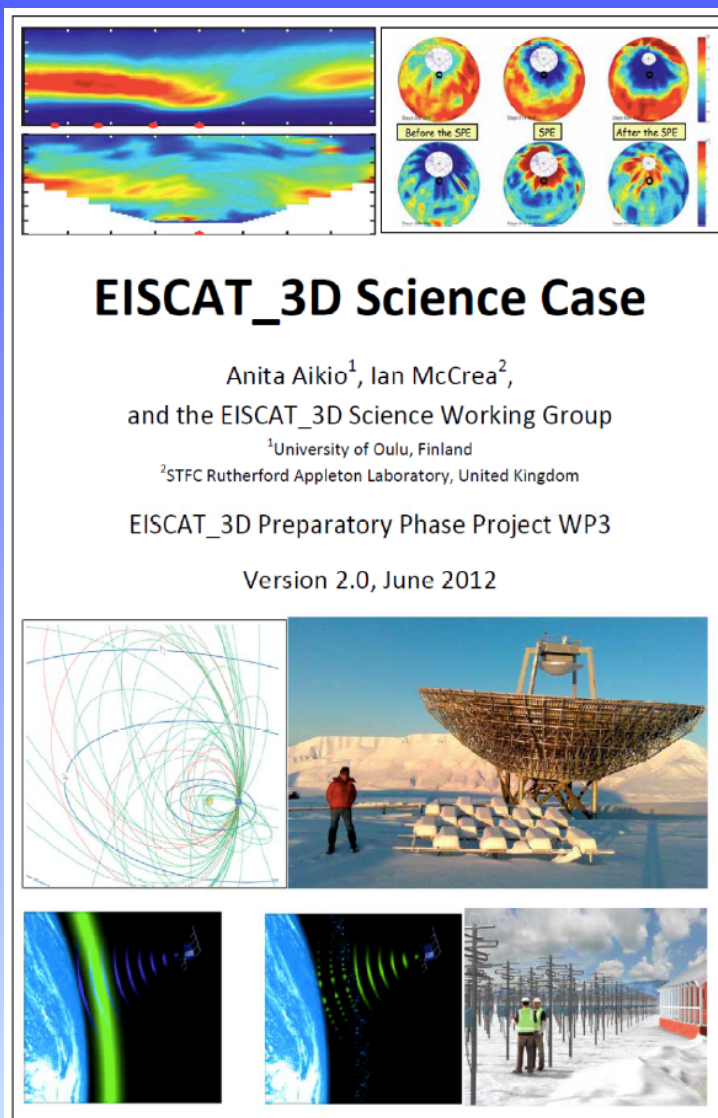
EISCAT_3D will have the sensitivity needed for ionospheric measurements at better than 100 msec time scales and 50-100 meter spatial scales (order of magnitude improvements over current systems)





EISCAT_3D Science

- An extremely versatile and largely software-defined instrument
- Specific science plans are developed within national user communities
 - Norway's proposal poses 5 major questions
 - Sweden's proposal has 5 active areas for research
 - Finland's roadmap proposal has 5 active areas for research
- Easy expansion to new fields
- Easy inclusion of new nations



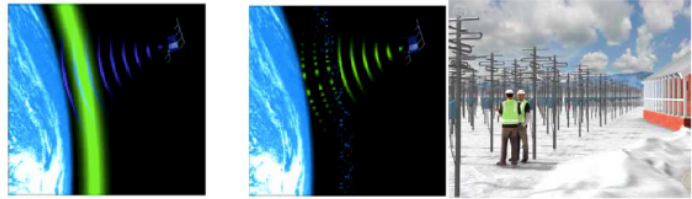

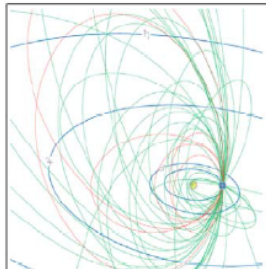
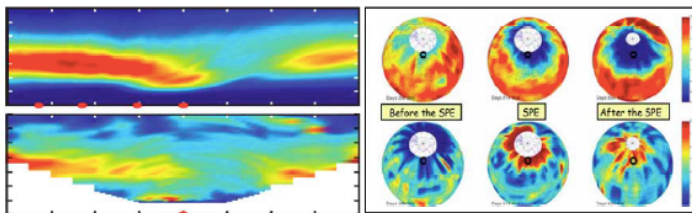
EISCAT_3D Science Case

Anita Aikio¹, Ian McCrea²,
and the EISCAT_3D Science Working Group

¹University of Oulu, Finland
²STFC Rutherford Appleton Laboratory, United Kingdom

EISCAT_3D Preparatory Phase Project WP3

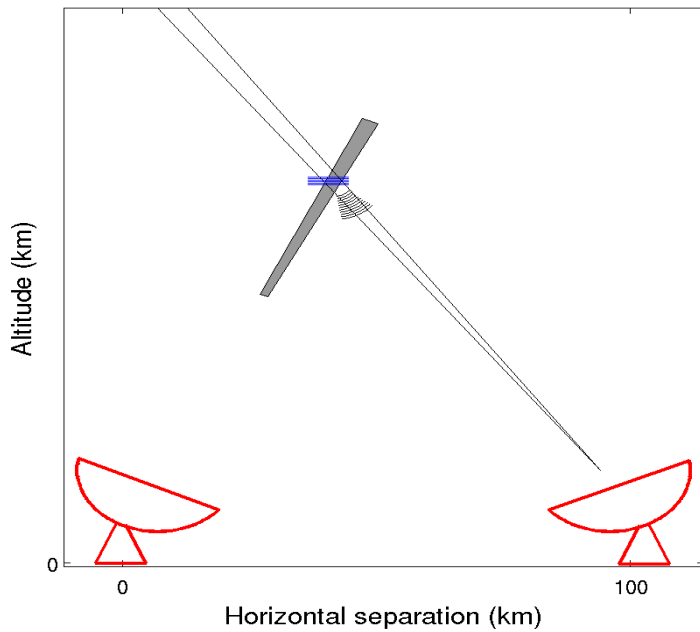
Version 2.0, June 2012



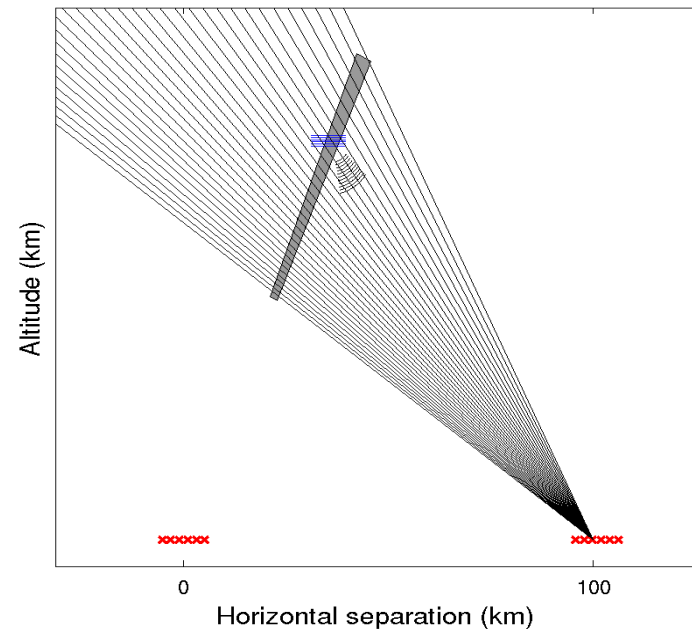


EISCAT 3D – the next generation ISR

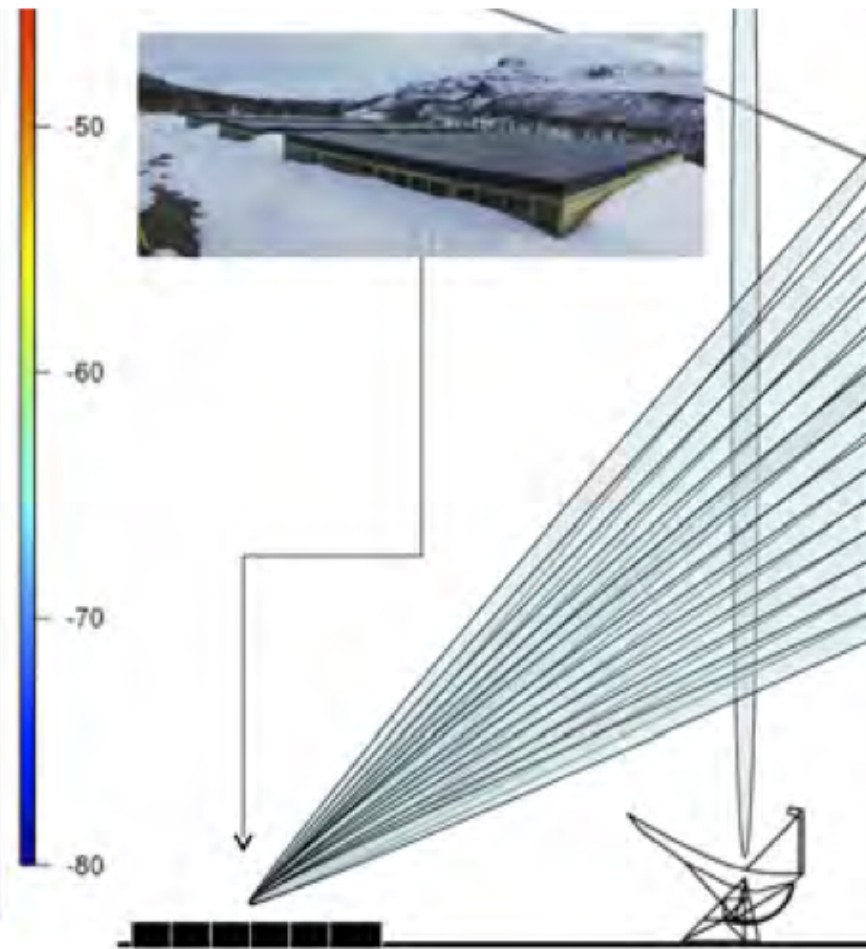
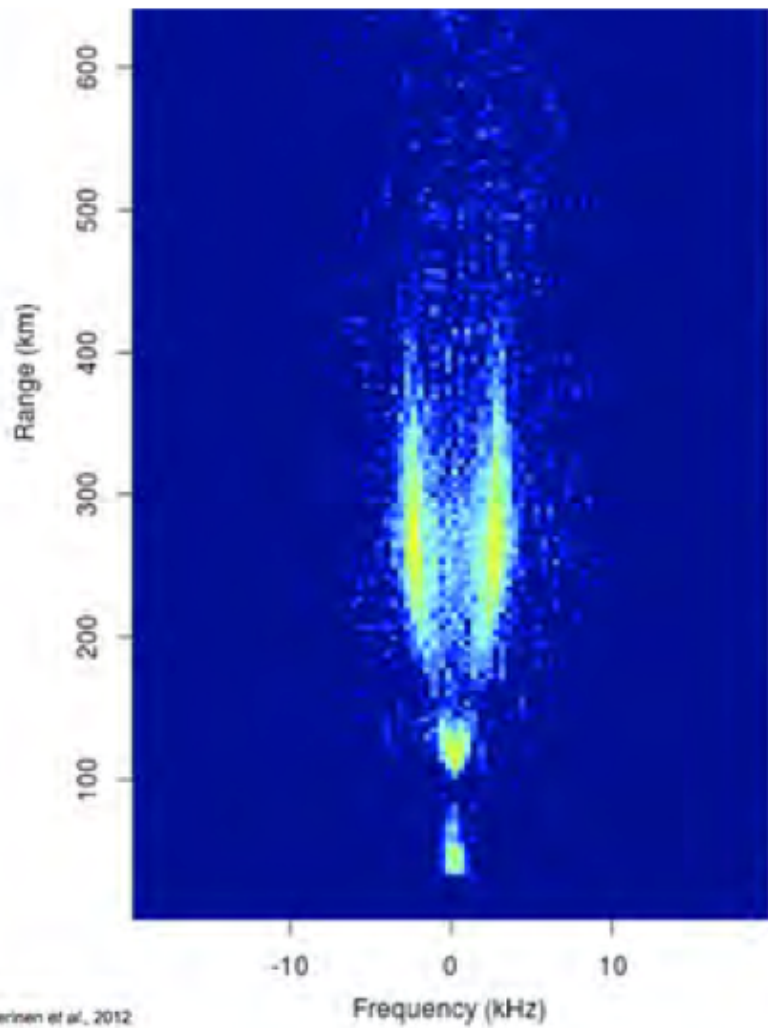
Bistatic dish antenna system



Bistatic phased array system



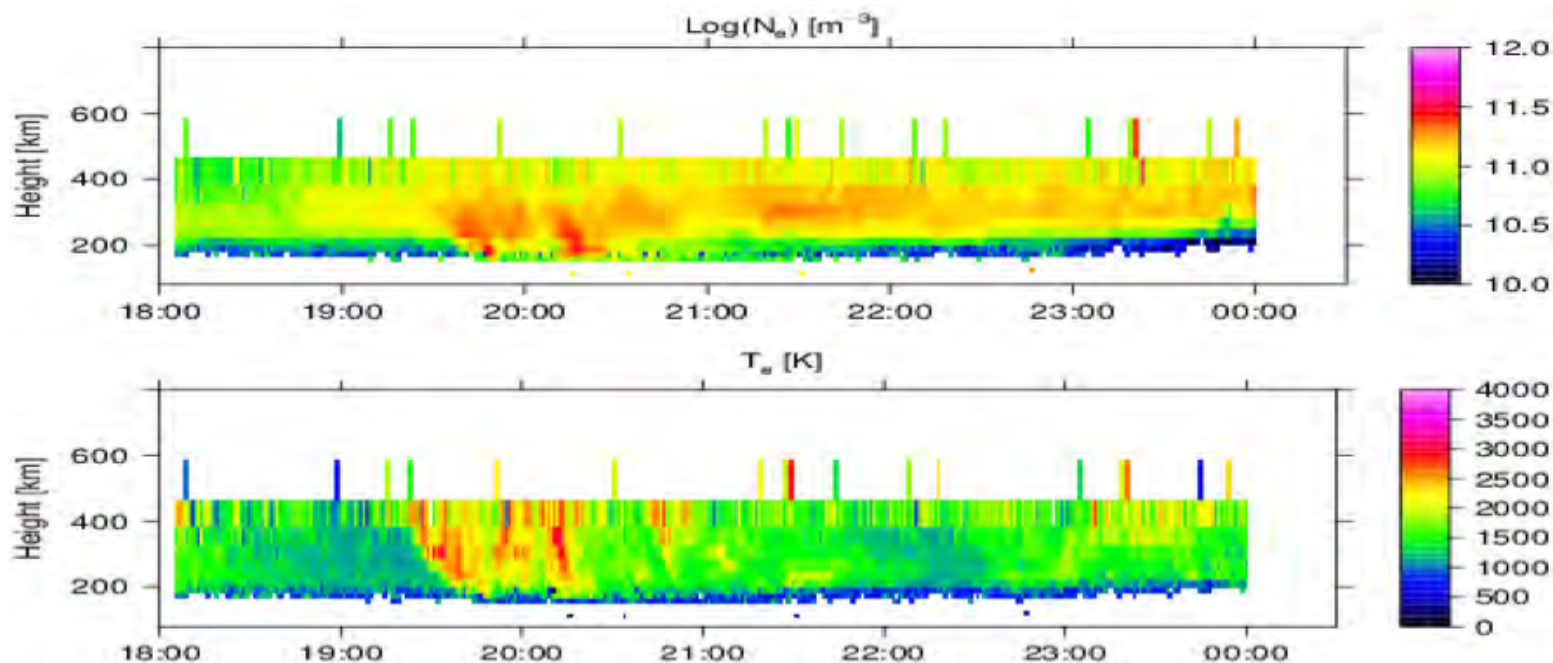
First multibeam receiver data analysis, KAIRA receiving EISCAT VHF



Verinen et al., 2012



Bistatic KAIRA compared to monostatic EISCAT VHF



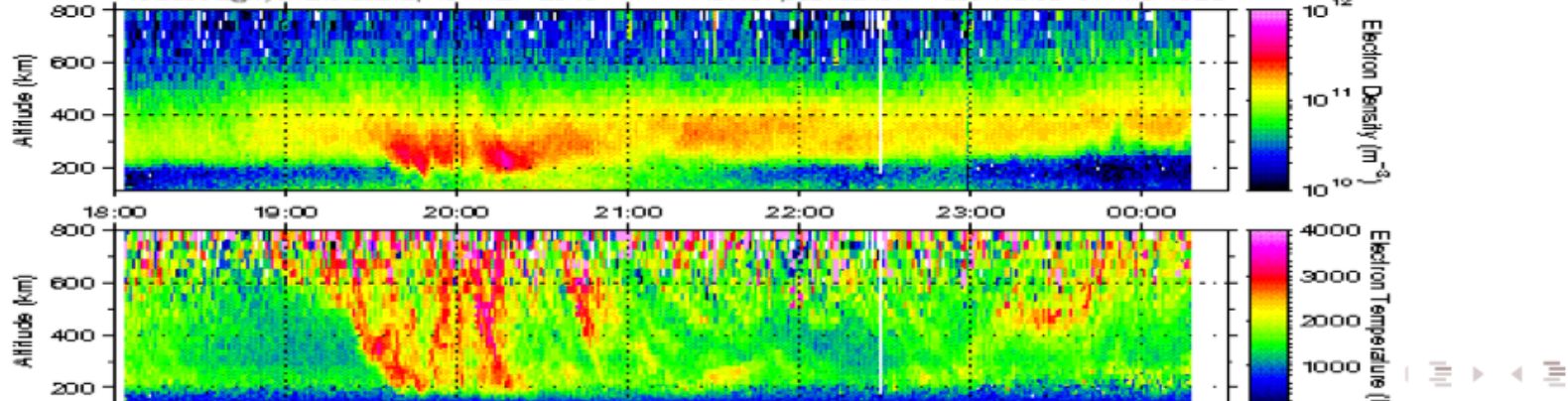
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EISCAT VHF RADAR

RT, vhf, bella, 14–15 March 2013

Produced @ phys-ateiscat12, 15-Mar-2013

Not for publication - see Rules-of-the-road



Embedded Antenna Element

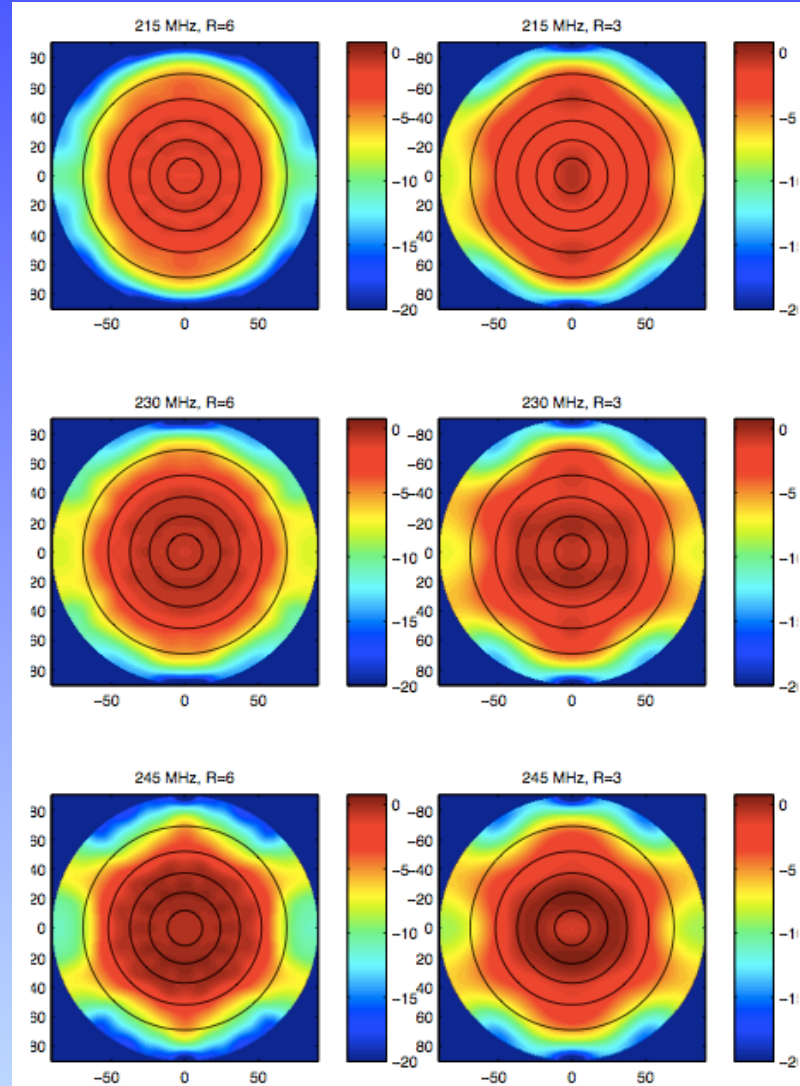
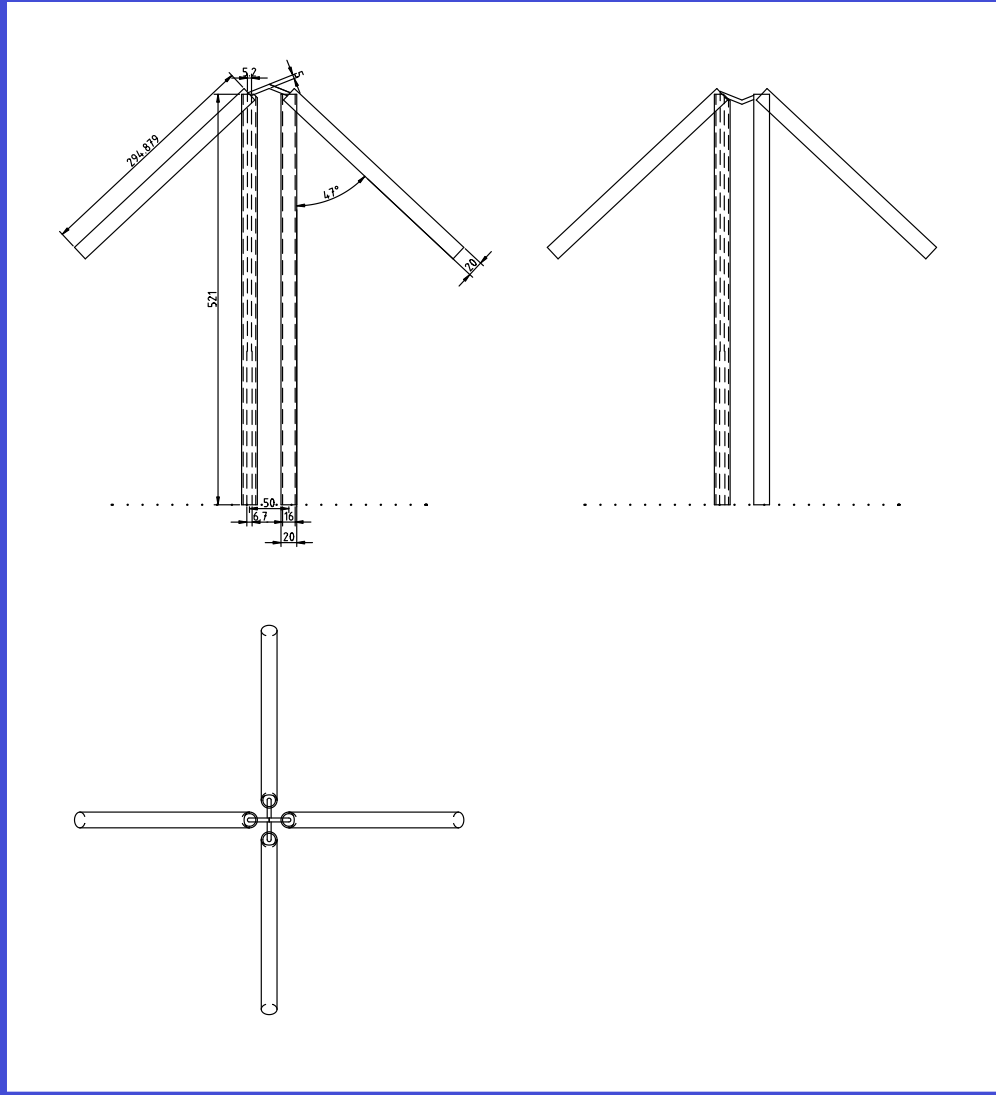
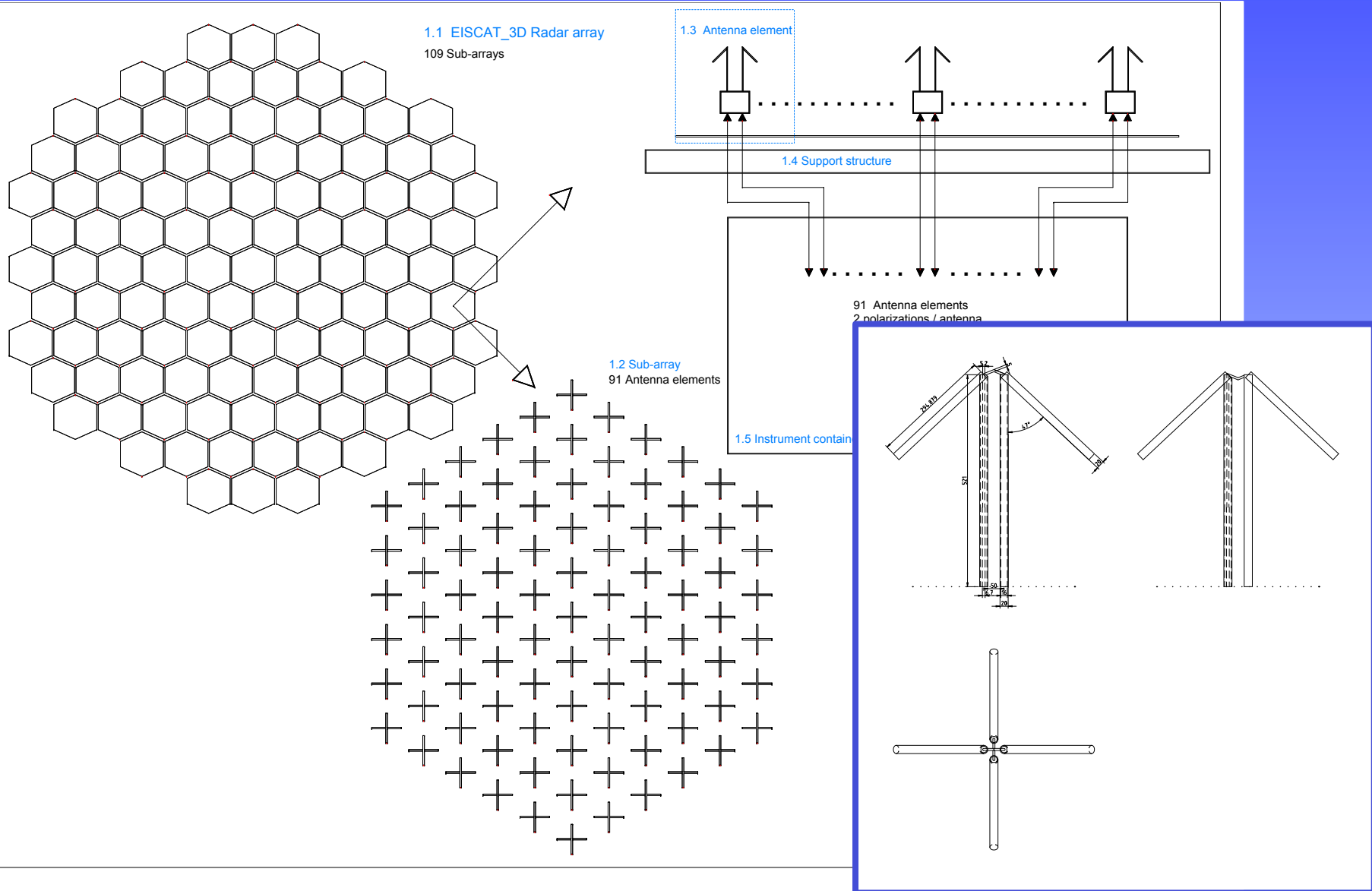


Figure 10: Gain of the selected antenna



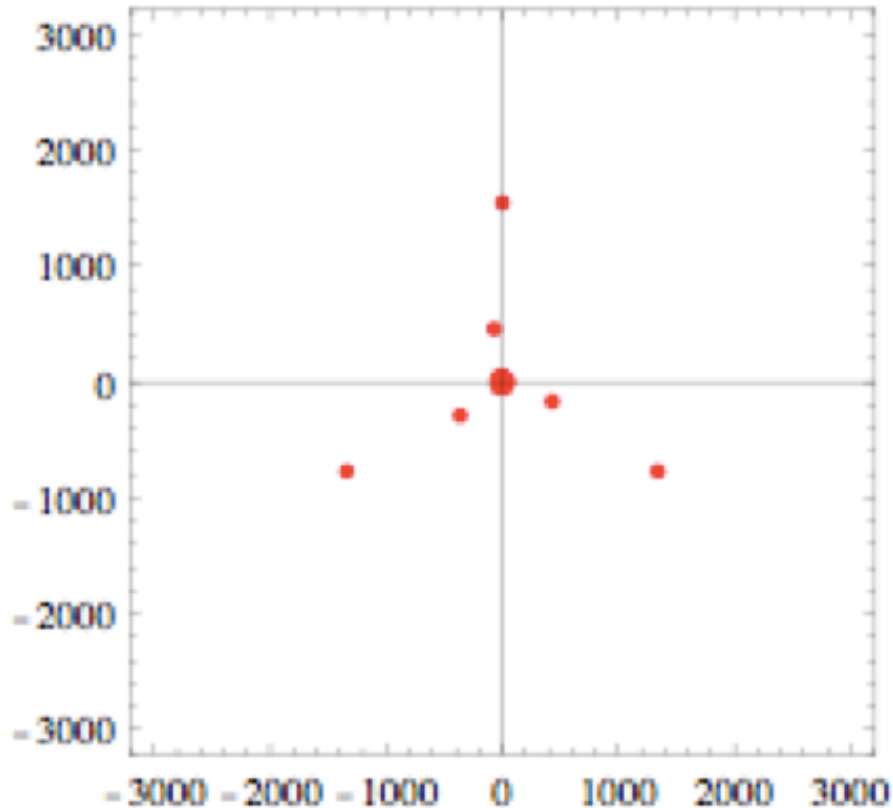
EISCAT_3D Site Dense Core



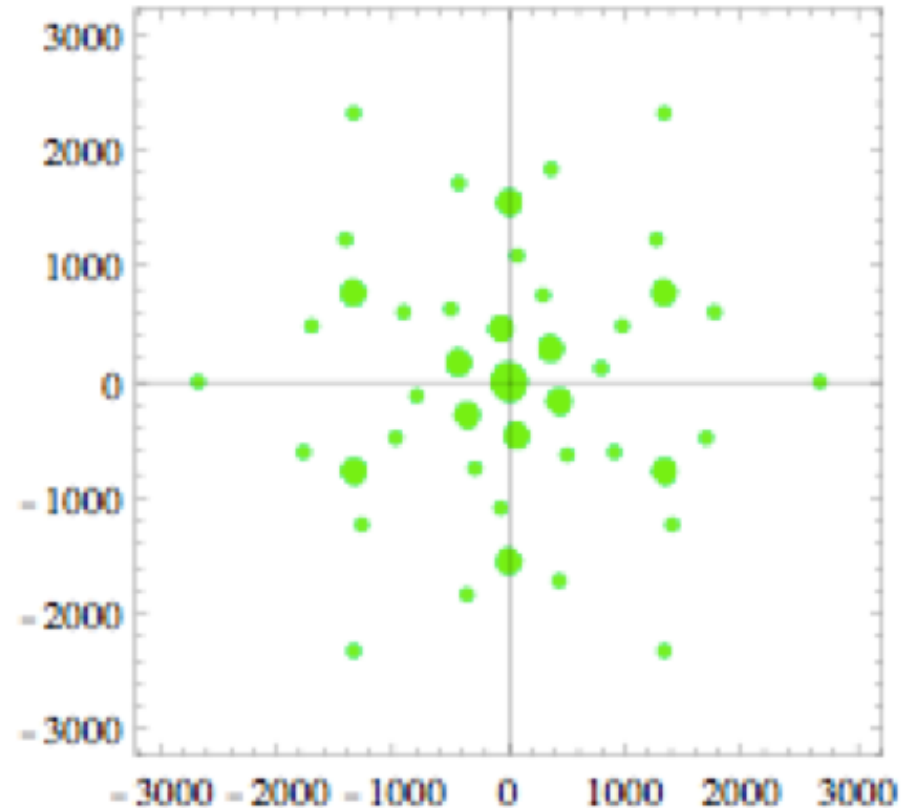


Aperture Synthesis Imaging

Antenna positions #16L

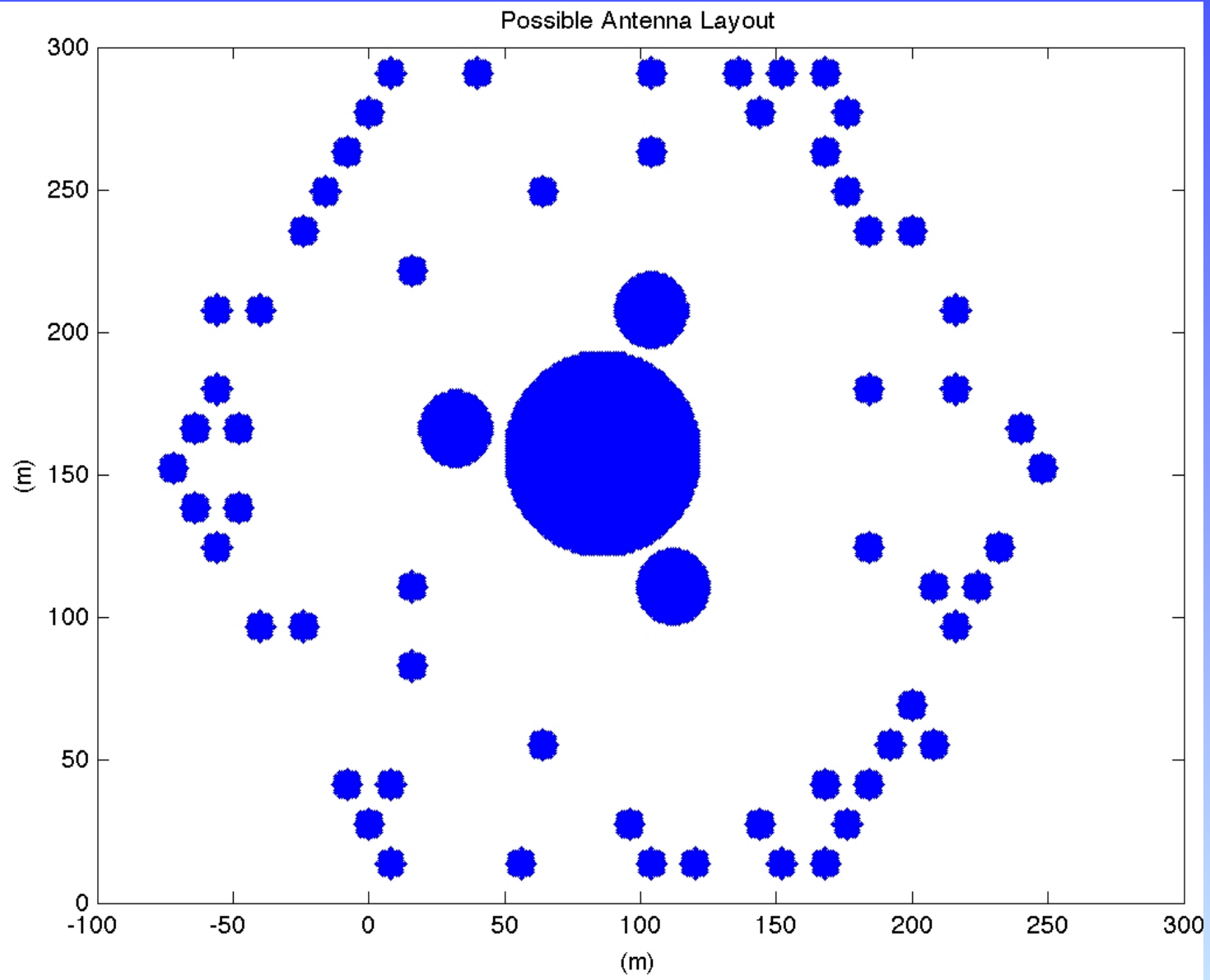


Baselines #120L





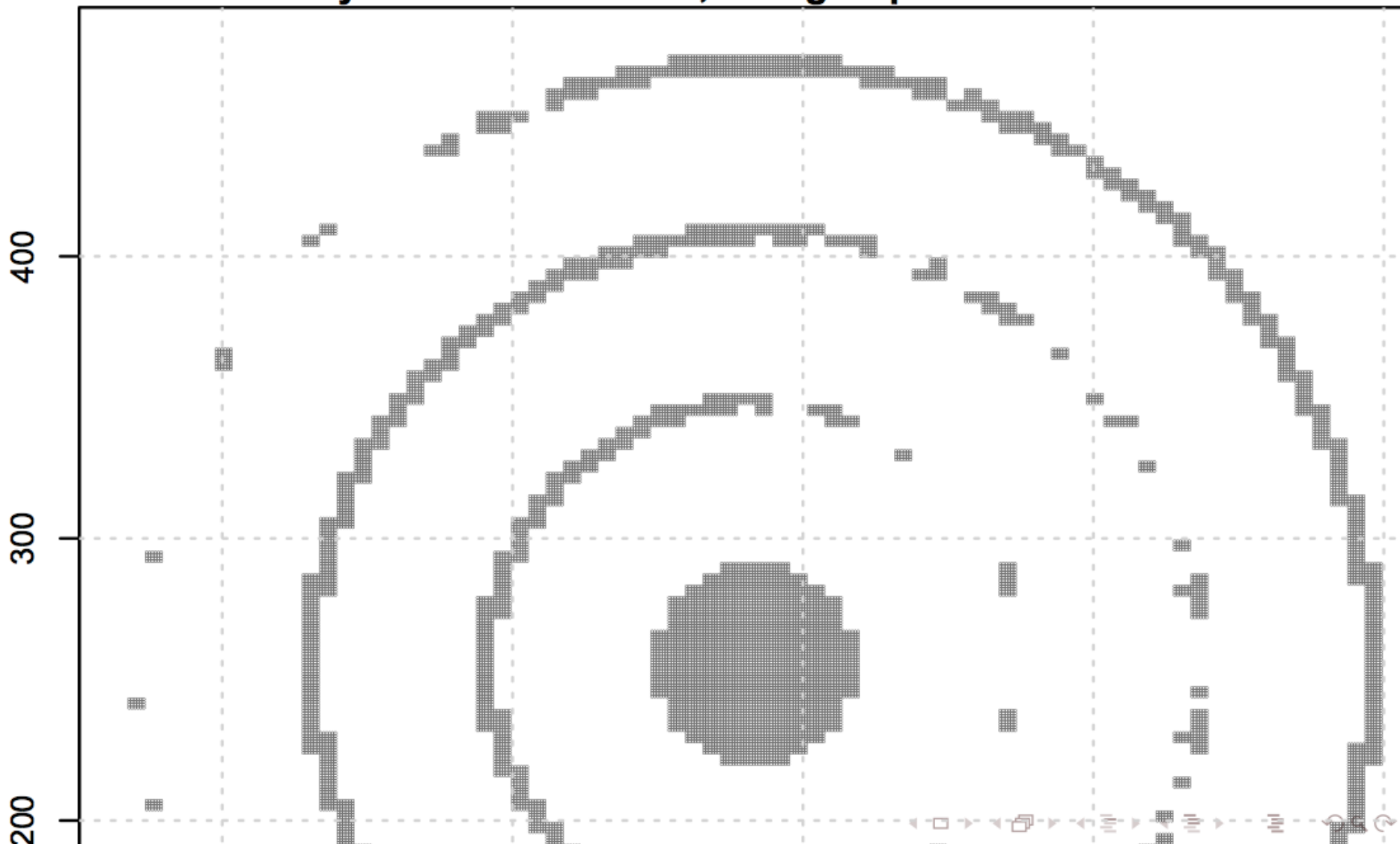
Another Possible Layout





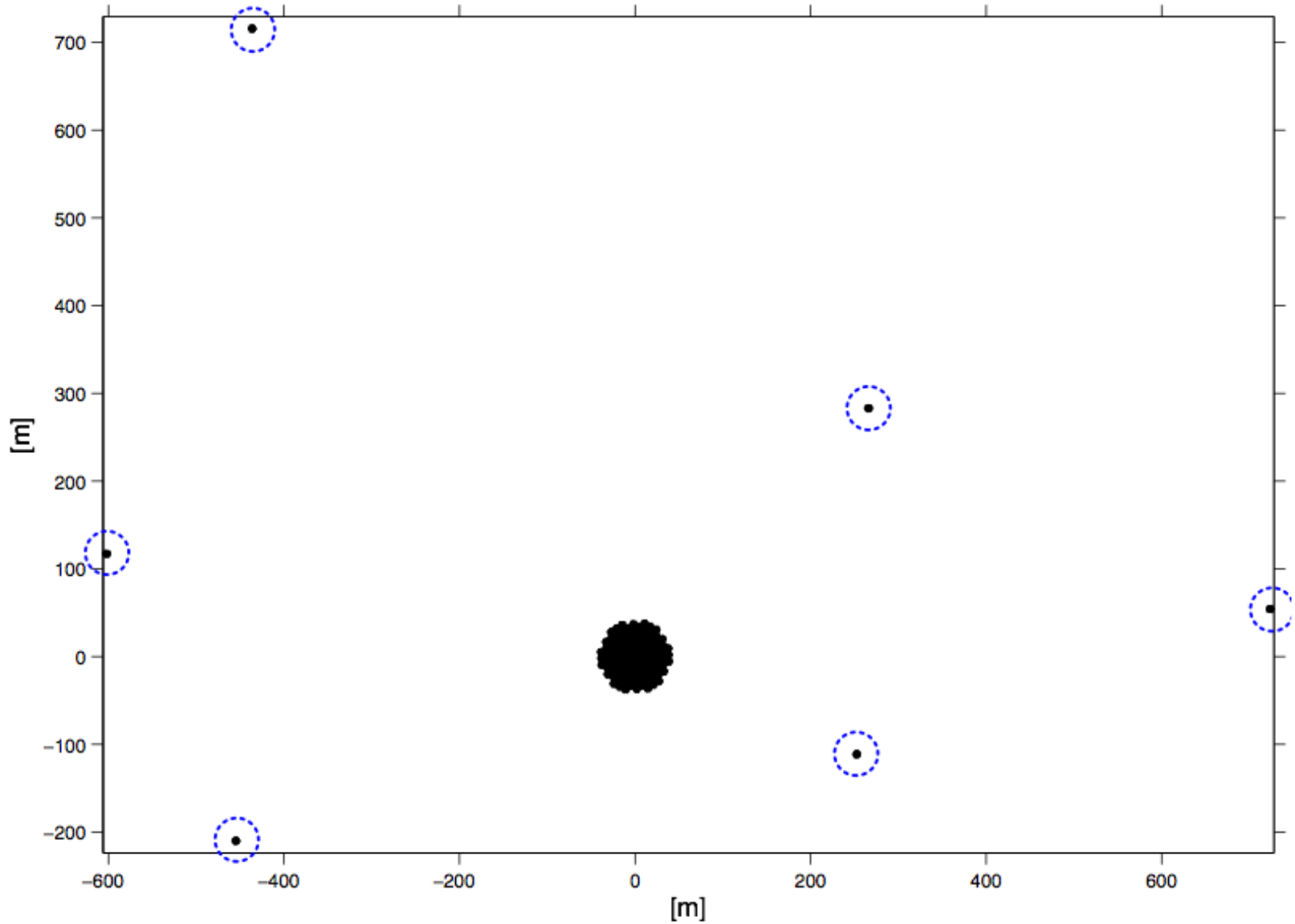
Yet Another Possible Layout

Array of 16896 antennas, 704 groups of 4 x 6 elements



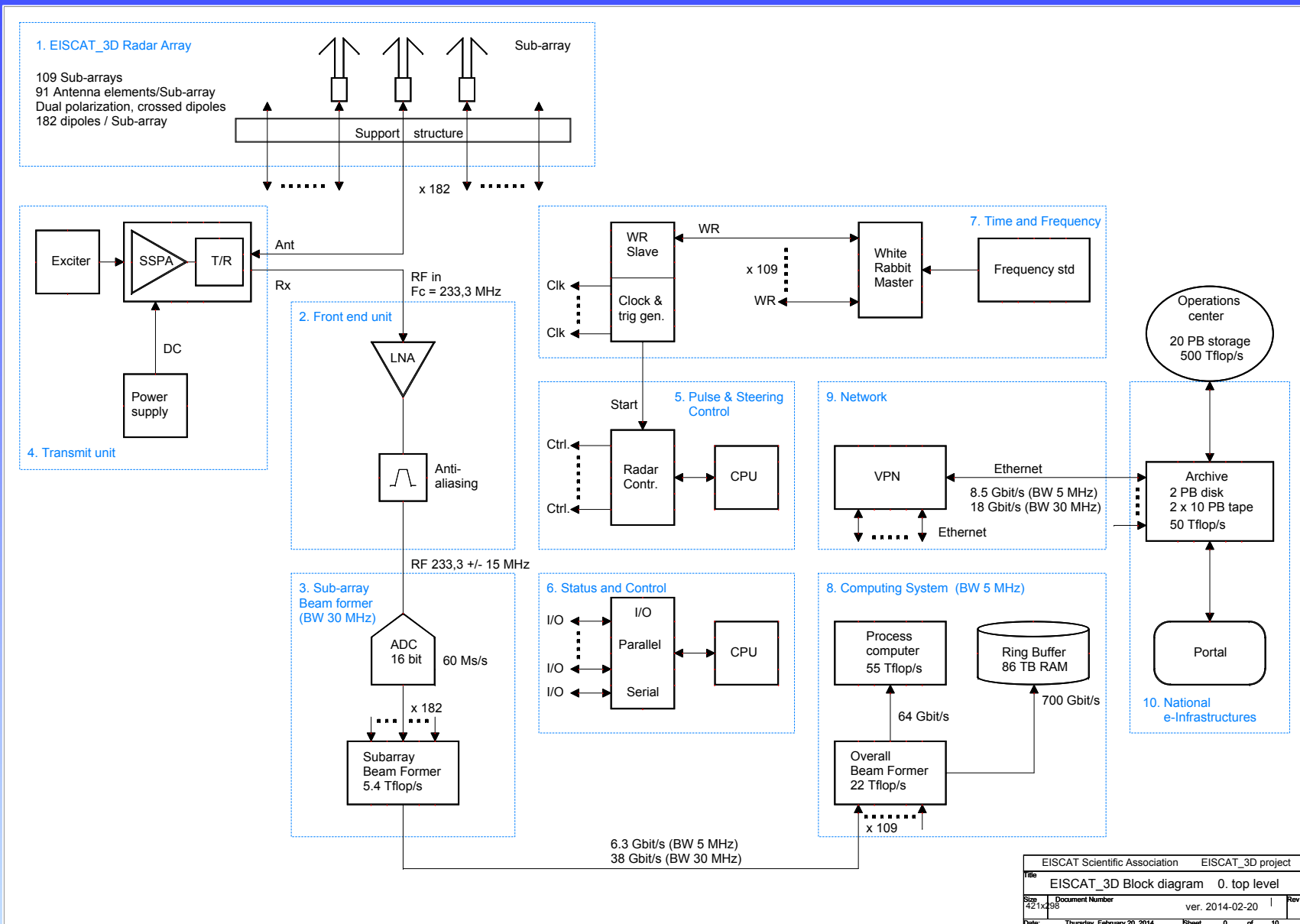


What we may end up with





Component Modules





EISCAT_3D Instrument

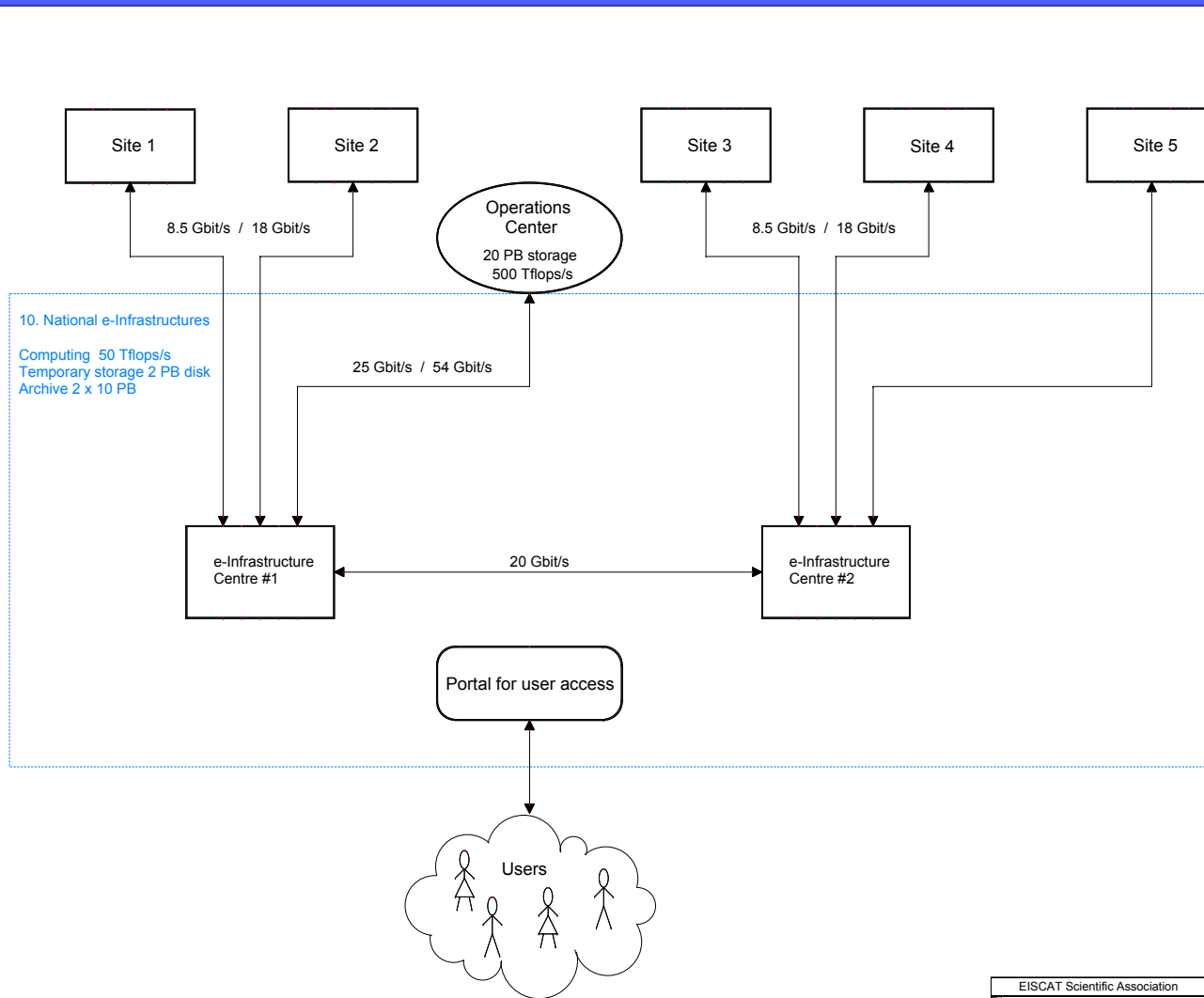
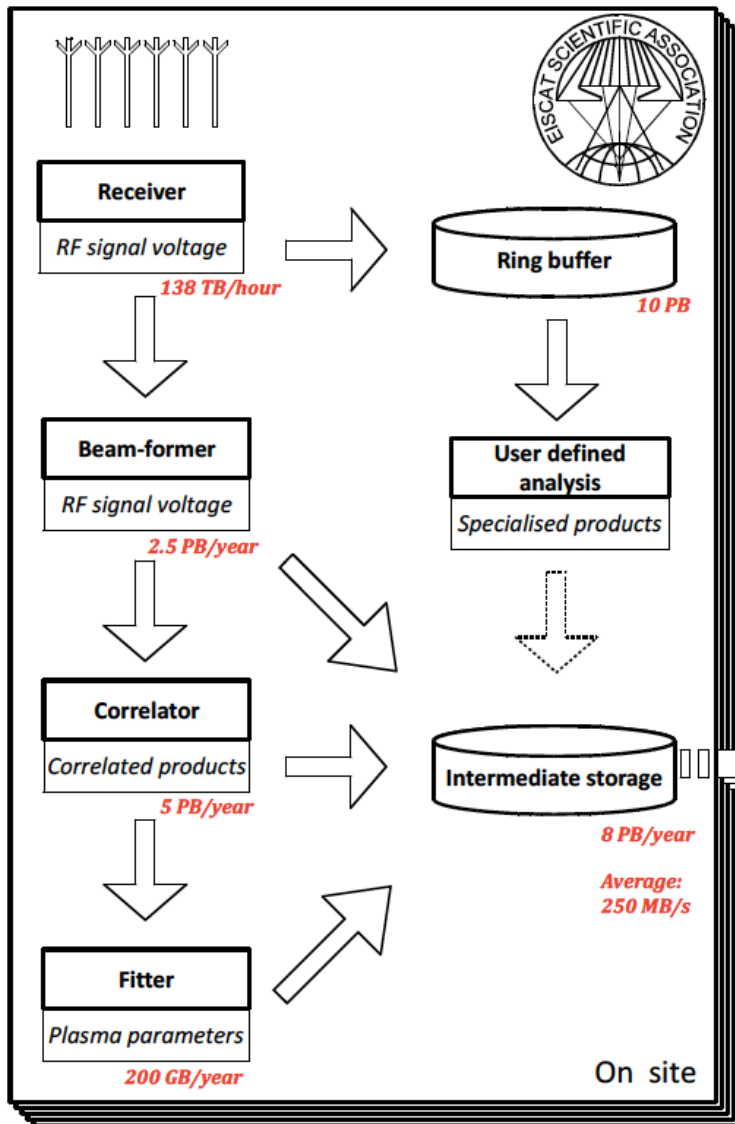


Table 1**EISCAT_3D Specifications**

Transmitter			Features		
1	Type	Pulsed	-	1	3-dimensional vector and scalar imaging
2	Duty Cycle	25	%	2	Phase/Amp transmitter modulation
3	Max Pulse Length	10	ms	3	Arbitrary Tx/Rx polarisation
4	Shortest Pulse Length	0.5 (75)	μs (m)	4	One Core Active site
5	Peak Power	10	MW	5	4 Remote Passive sites
6	Centre frequency	233	MHz	6	Remote active site with power ~ 1 MW
7	3 dB Bandwidth	± 2.5	MHz	7	Electronic scanning and beam-forming
8	Modulation	Phase/Amp	-	8	5 phased steered array antennas
9	Polarization	Arbitrary	-	9	6-10 outlying antenna arrays at active site
Antennas				10	Aperture synthesis imaging
10	Type	Phased Array	-	11	High duty cycle
11	Antennas per site	$\sim 10,000$	-	12	Better sensitivity by a factor $> 20\times$
12	3 dB Bandwidth	± 15	MHz	13	Unmanned operation
13	Gain	~ 50	dBi	14	Remote operation via internet
14	beam pointing resolution	0.625°		15	Robust reliability in arctic environment
15	Transverse resolution at 100 km better than	50	m	16	Uninterrupted continuous operations
16	Grating-lobe free radiation pattern: zenith all azimuths	40°		17	Common Programs at low duty cycle
17	Power-aperture product	> 100	GWm^2	18	Special Programs to respond to pre-defined/unusual geophysical events
				19	Validated archival database
				20	Restricted real-time data

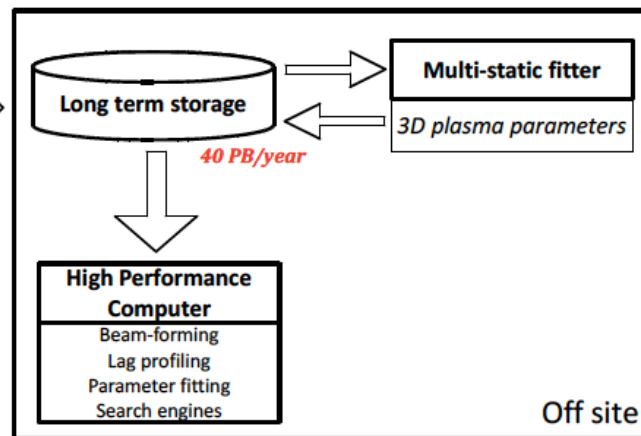


e-Infrastructure/e-Science



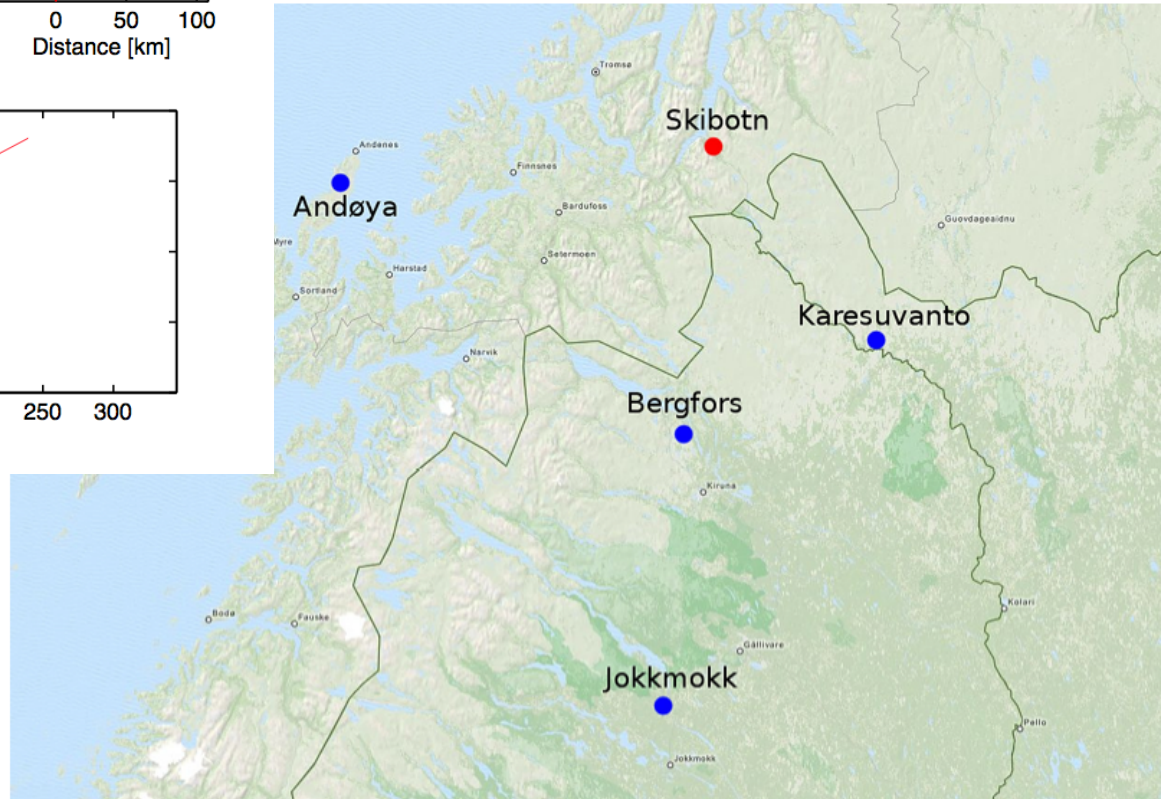
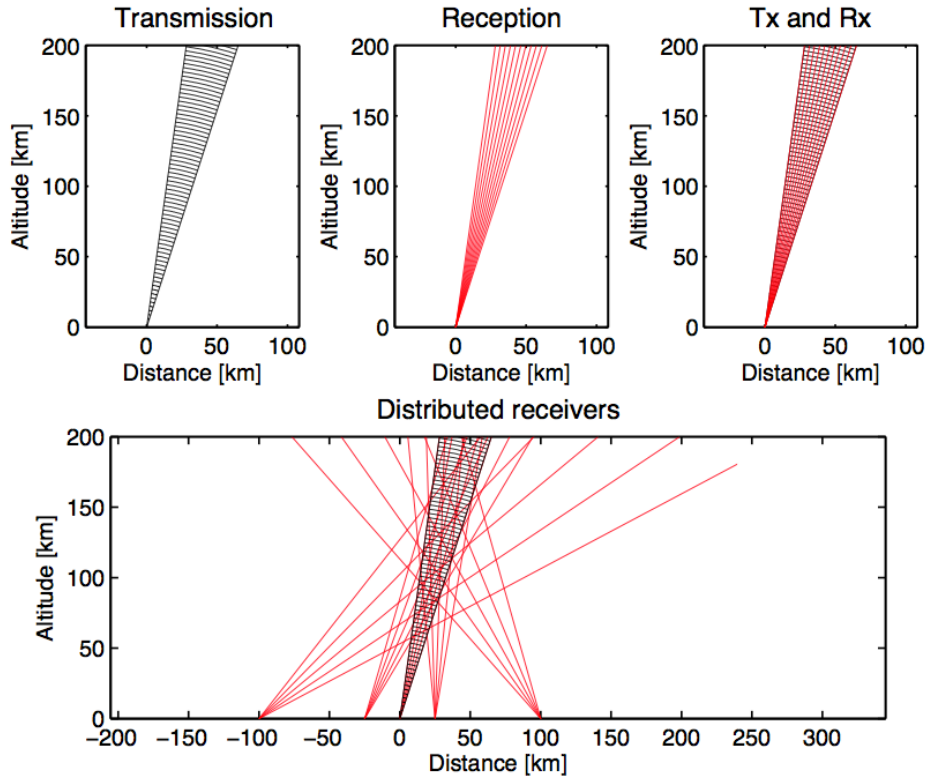
Data Flow EISCAT_3D

Antennas	Bandwidth from each antenna: 30 MHz Sampling from each antenna: 60 MS/s Data-production per antenna (2 Bytes per sample): 120 MB/s
Antenna groups	100 antennas per group forming beam with two polarisations (pol.) Data-production per group: $2 \times 120 \text{ MB/s} = 240 \text{ MB/s}$ Assuming 16,000 antennas on one site, there will be 160 groups Total data production: $160 \times 240 \text{ MB/s} = 38.4 \text{ GB/s} = 138 \text{ TB/h}$ Storage in ring buffer (up to three days of full time data): 10 PB
Beam-forming	Antenna groups on each site are combined into 100 beams Data decimated to 1 MHz Data per beam: $2 \times 1 \text{ MHz} \times 2 \text{ B/sample} \times 2 \text{ pol.} = 8 \text{ MB/s}$ Total volume of beam-formed data: $100 \times 8 \text{ MB/s} = 800 \text{ MB/s}$
Correlator	Correlated data is reduced by a factor 10 compared to raw data Correlated data use two bands with 10 MHz bandwidth Beam: $2 \text{ bands} \times 2 \times 10 \text{ MHz} \times 2 \text{ B/sample} \times 2 \text{ pol.} / 10 = 16 \text{ MB/s}$ Total volume of correlated data: $100 \times 16 \text{ MB/s} = 1.6 \text{ GB/s}$
Fitter	Data produced about 200 GB/yr
Data per year	Assume system running at 10% capacity over one year: Beam-formed data: $0.1 \times 365 \times 24 \times 3600 \times 800 \text{ MB} = 2.5 \text{ PB/yr}$ Correlated data: $0.1 \times 365 \times 24 \times 3600 \times 1.6 \text{ GB} = 5 \text{ PB/yr}$ Fitted data: 200 GB/yr. Total data to store for one site: 8 PB/yr





Multistatic Phased Array





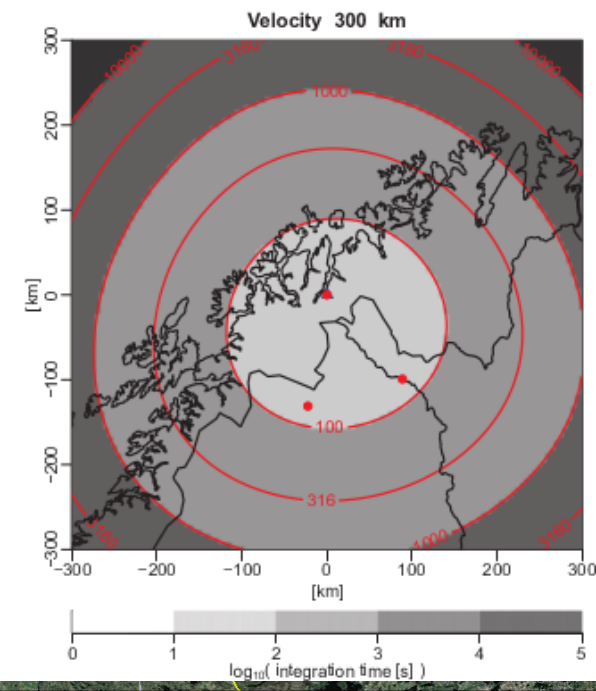
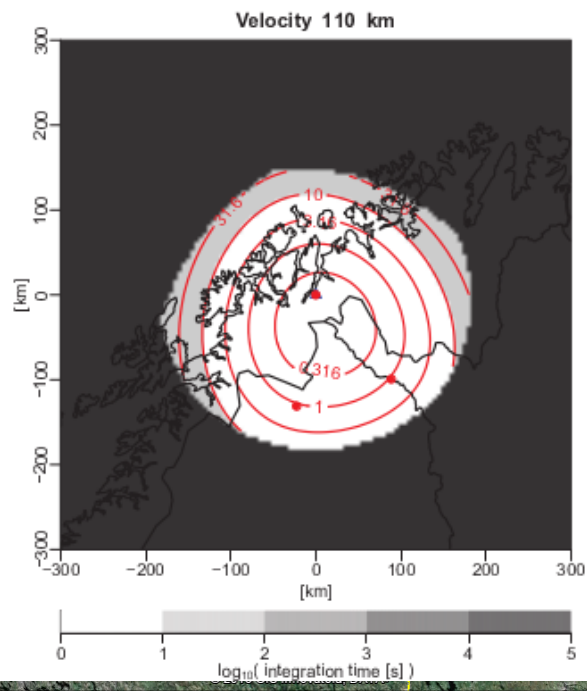
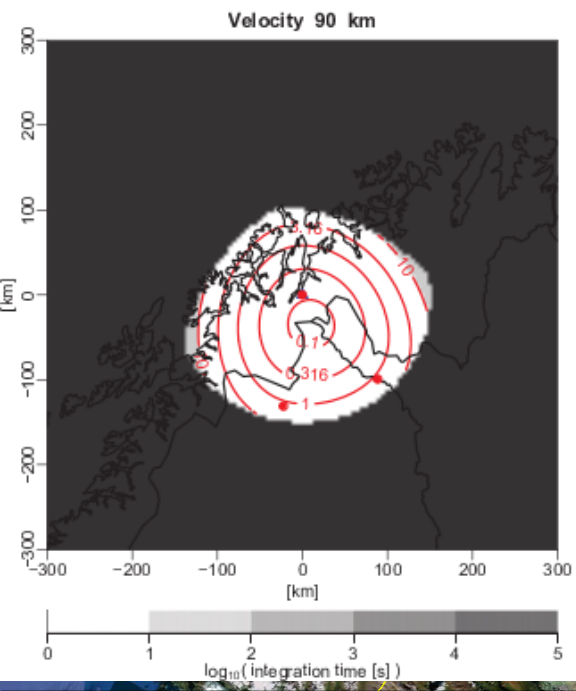
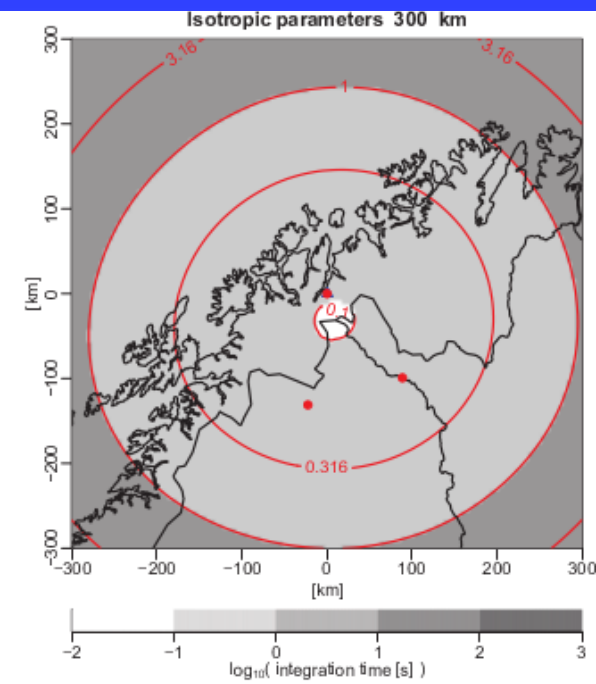
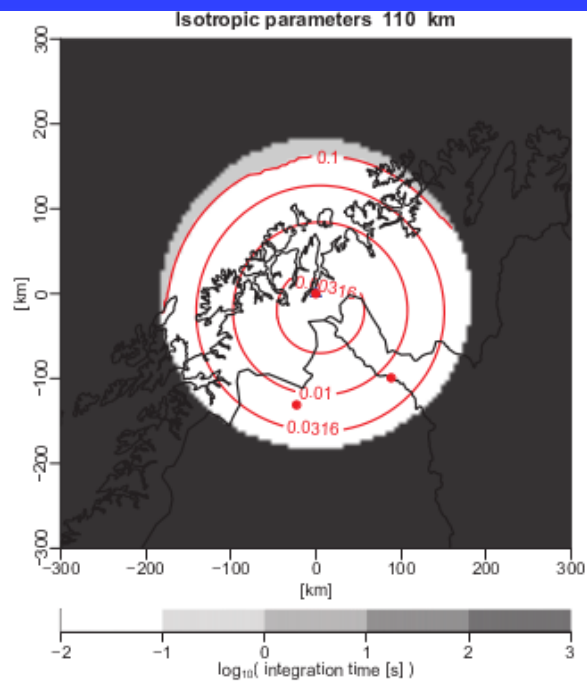
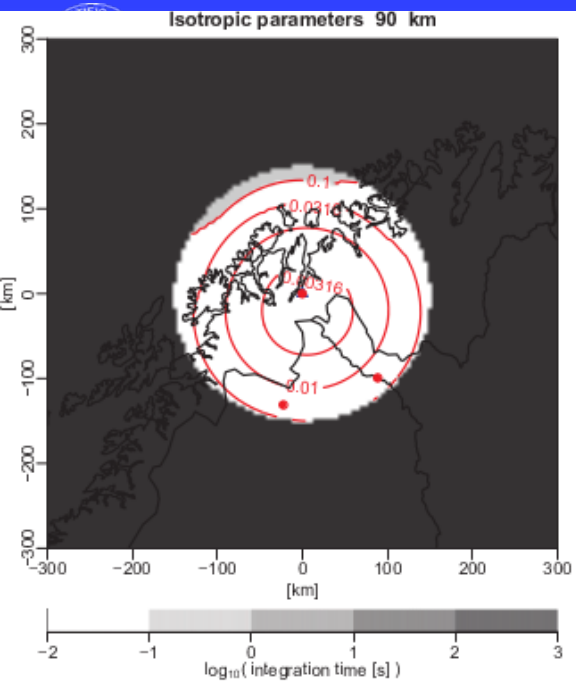
Staged Implementation

After Stage 4, total of 5 sites

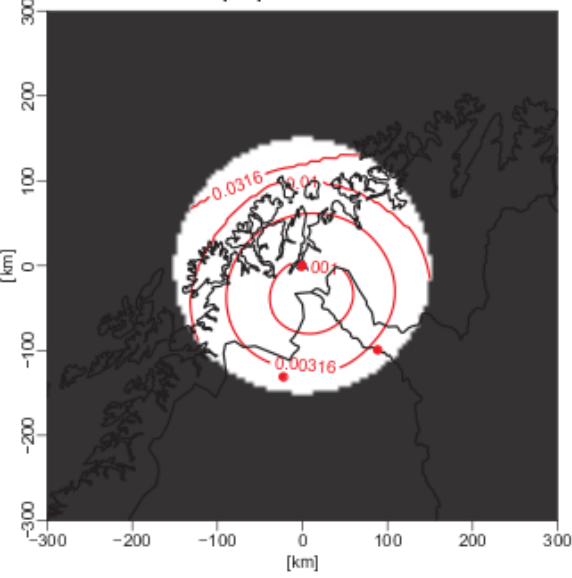
Core with 10 MW transmitter (Skibotn, Norway)

4 receive-only (-mostly) sites: Berfgors, Sweden;
Karesuvanto, Finland; Andøya, Norway;
Jokkmokk, Sweden

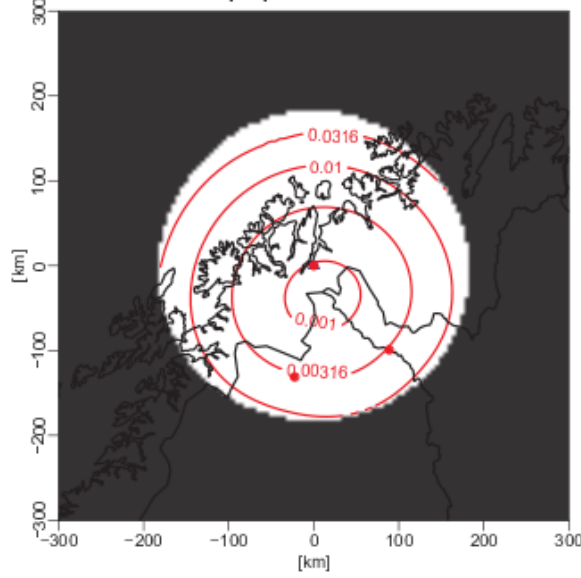
Construction plan ordered to provide useful
capabilities at each stage



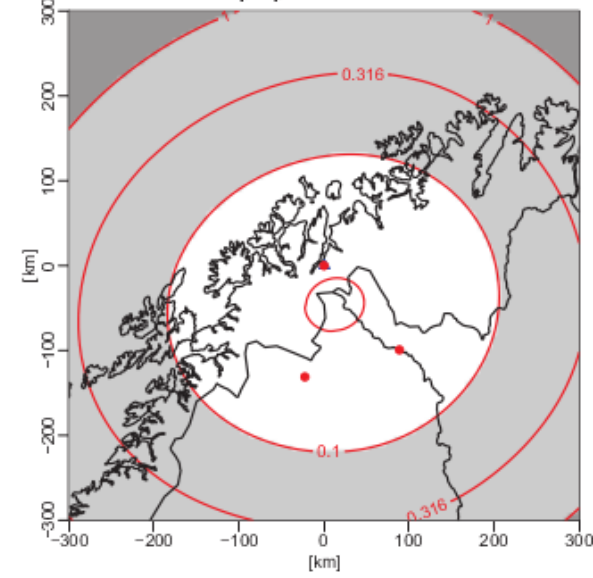
Isotropic parameters 90 km



Isotropic parameters 110 km



Isotropic parameters 300 km

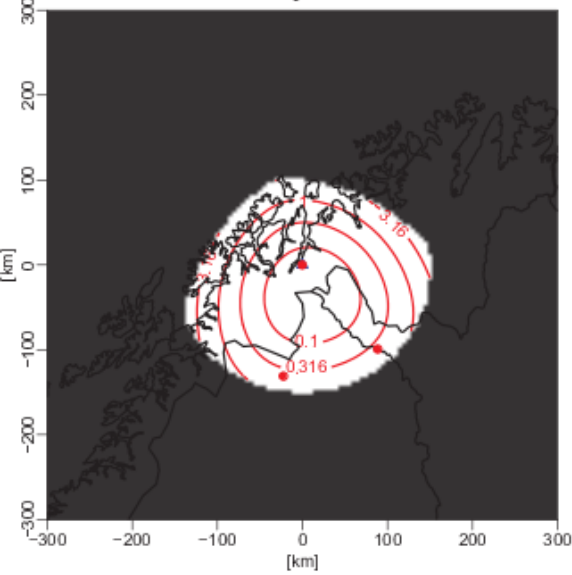


$\log_{10}(\text{integration time [s]})$

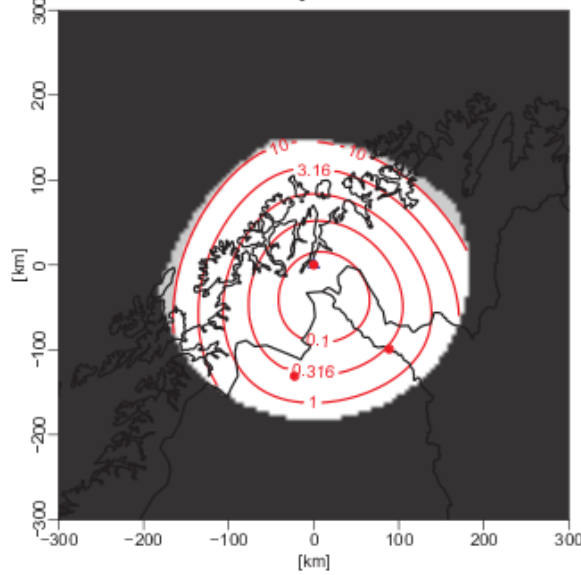
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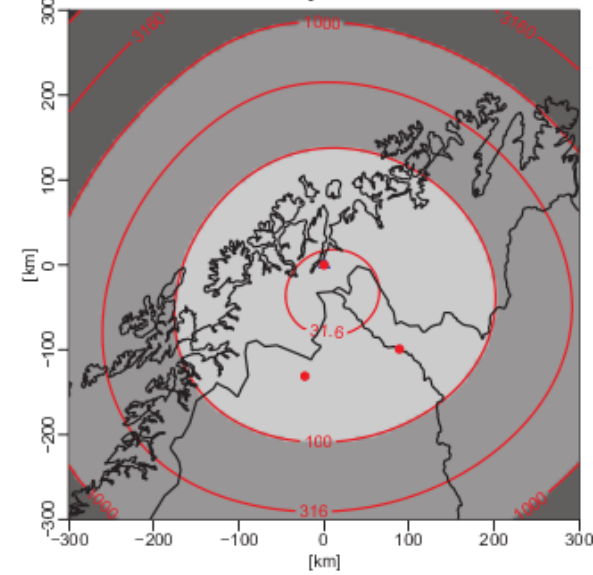
Velocity 90 km



Velocity 110 km



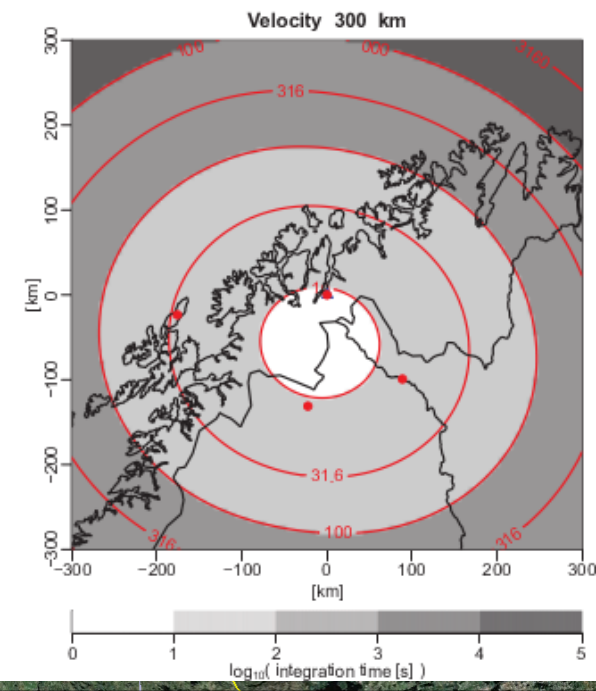
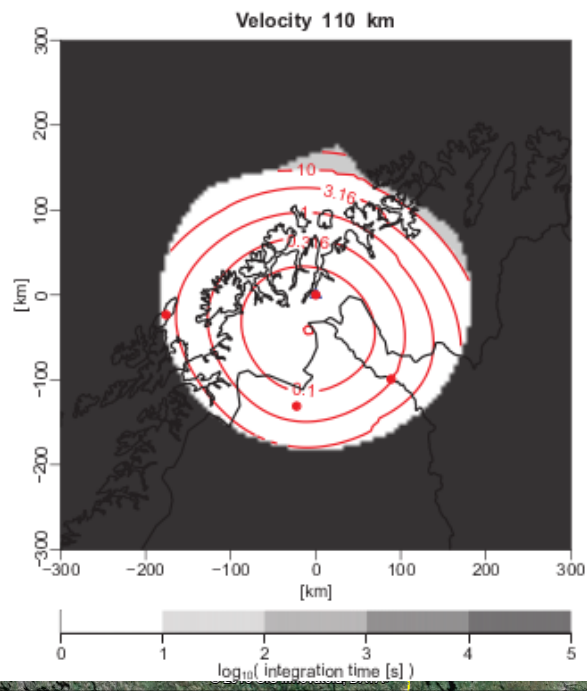
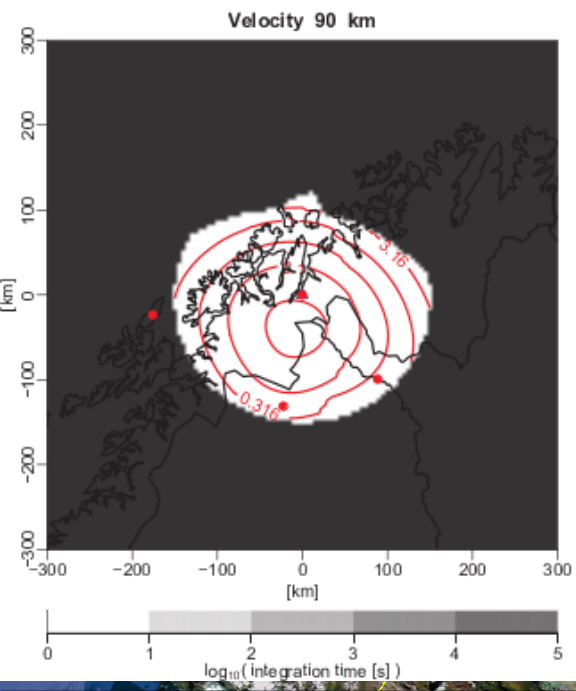
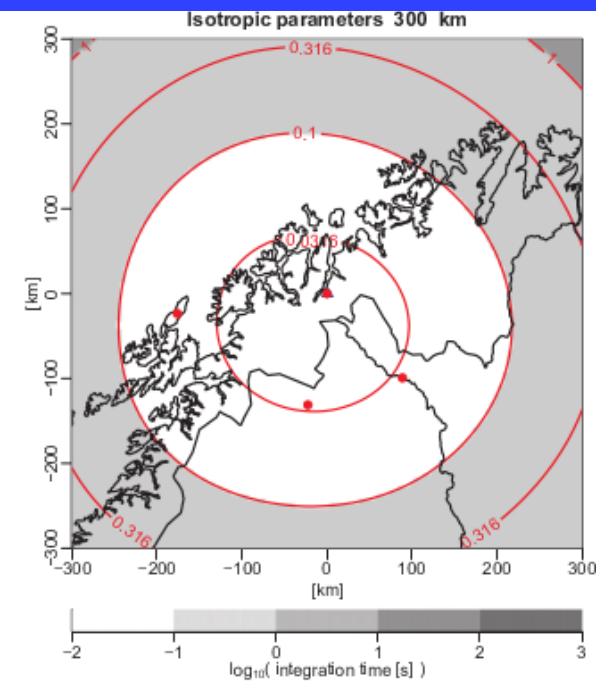
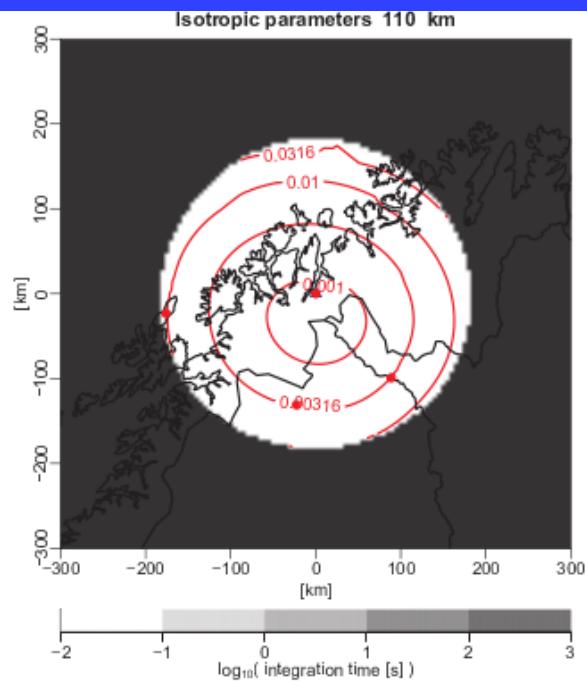
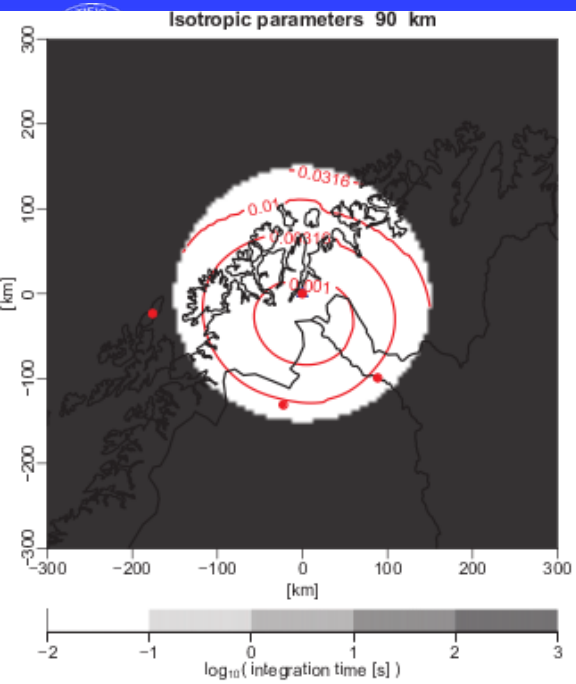
Velocity 300 km

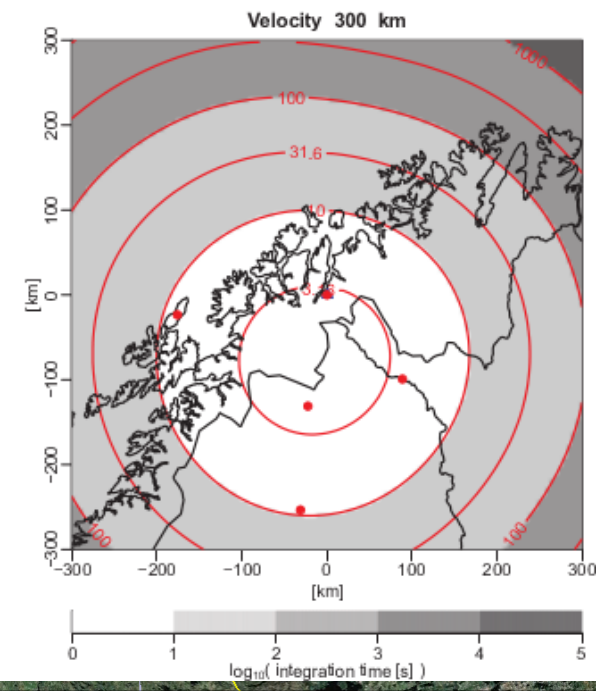
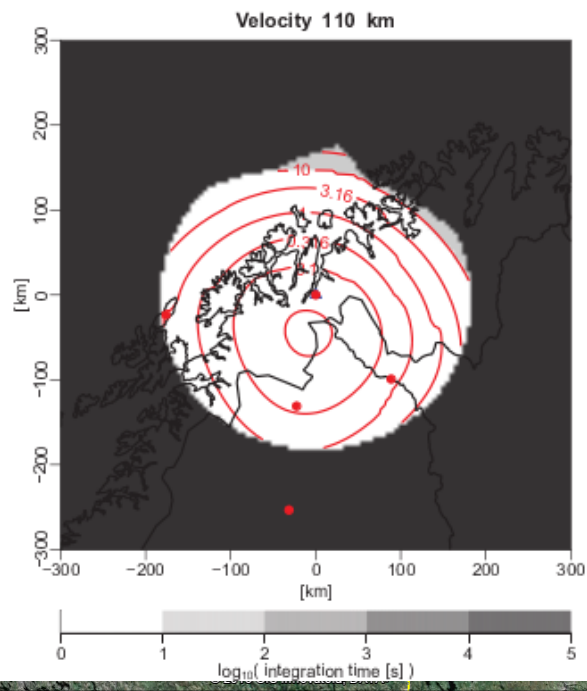
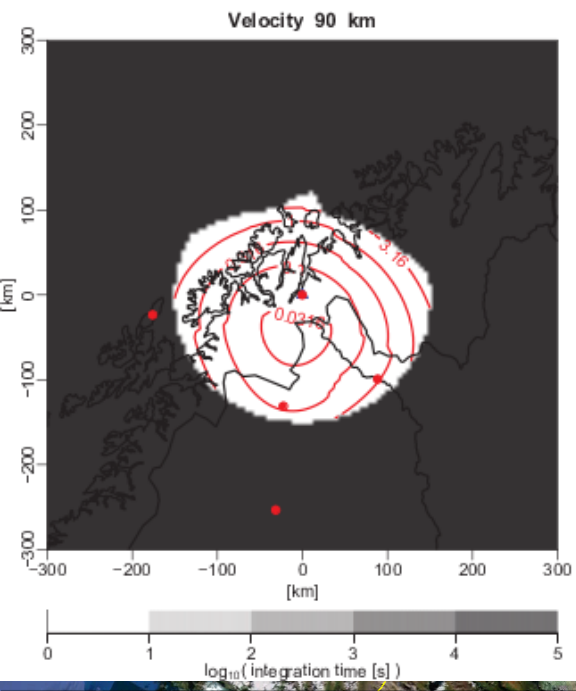
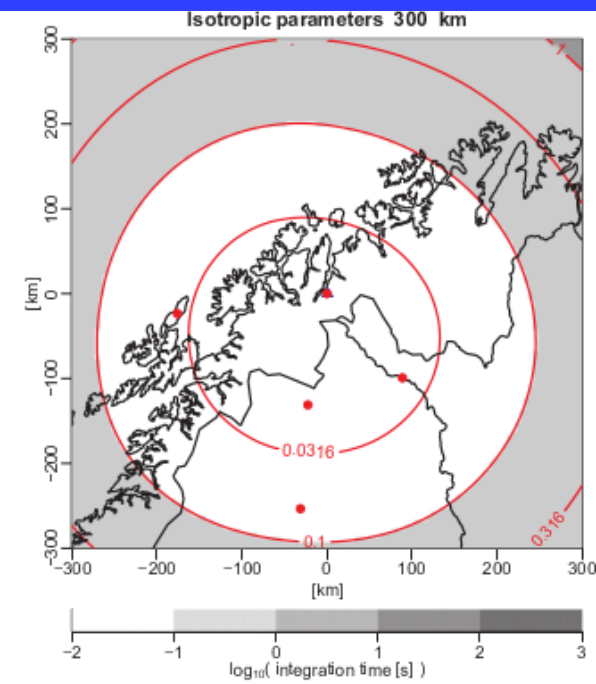
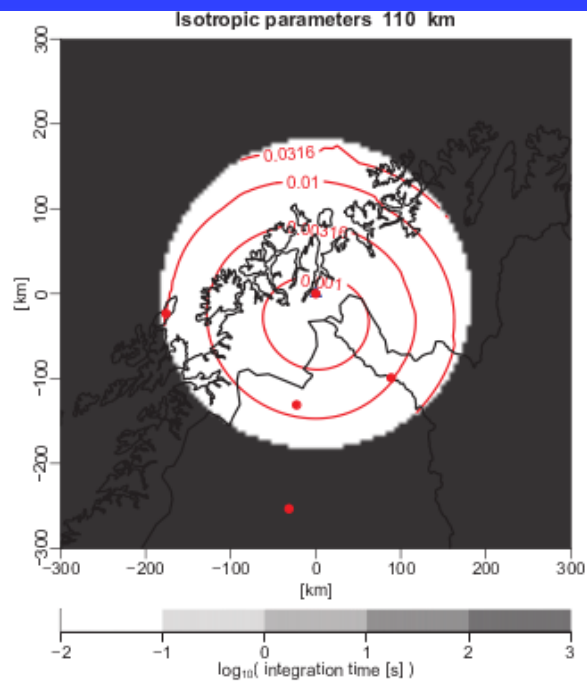
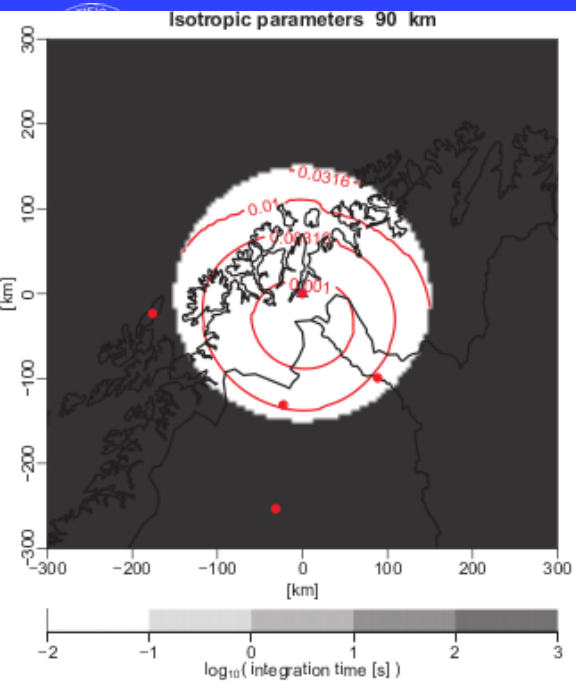


$\log_{10}(\text{integration time [s]})$

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$\log_{10}(\text{integration time [s]})$







	Stage 1	Stage 2	Stage 3	Stage 4
1. Atmospheric physics and global change				
a. Vertical coupling between the atmospheric layers				
i. Troposphere	X	X	X	X
ii. Stratosphere	P	X	X	X
iii. Mesosphere	P	P	X	X
b. Turbulence and waves in the mesosphere and lower thermosphere				
i. Polar mesospheric summer echo (PMSE) interferometry	P	X	X	X
ii. MAARSY collaborations (Andøya, Norway)	P	P	X	X
iii. Polar mesospheric winter echo (PMWE) interferometry	P	X	X	X
2. Space and plasma physics				
a. Multiple scale interactions in ionosphere-magnetosphere plasmas				
i. Aurora (tens of meters, fractions of a second) density and velocity	P	X	X	X
ii. Magnetospheric-driven convection (localized & transient)		P	X	X
iii. Substorm processes over a large spatial region		P	P	X
b. Plasma turbulence and active experiments				
i. Naturally enhanced ion acoustic lines (NEIALs)	P	X	X	X
ii. Heating experiments (volumetric imaging of entire pump beam)	P	P	X	X
iii. Self heating in the EISCAT_3D beam		X	X	X
3. Inflow and outflow of matter in the Earth's atmosphere				
i. Meteoroid-atmosphere interactions	P	X	X	X
ii. Meteoric smoke particles (dusty plasmas)	P	X	X	X
iii. Ion upflow	P	X	X	X
4. Space debris, near-earth objects and space weather				
i. Space debris and meteoroid orbital parameters with wide coverage	P	P	X	X
ii. Near-earth objects (NEOs)	P	X	X	X
iii. Continuous measurements for space weather monitoring	P	P	P	X
5. Radio astronomy				
i. Epoch of reionization (EoR)	(P)	(P)	P	X
ii. Exoplanet detection via auroral kilometric radiation (AKR)-like emissions	(P)	(P)	P	X
iii. Trans-ionospheric imaging support	P	X	X	X

Table 1: Summary

1. Atmospheric physics and global change
a. Vertical coupling between the atmospheric layers
 i. Troposphere
 ii. Stratosphere
 iii. Mesosphere
b. Turbulence and waves in the mesosphere and lower thermosphere
 i. Polar mesospheric summer echo (PMSE) interferometry
 ii. MAARSY collaborations (Andøya, Norway)
 iii. Polar mesospheric winter echo (PMWE) interferometry
2. Space and plasma physics
a. Multiple scale interactions in ionosphere-magnetosphere plasmas
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5. Radio astronomy
 i. Epoch of reionization (EoR)
 ii. Exoplanet detection via auroral kilometric radiation (AKR)-like emissions
 iii. Trans-ionospheric imaging support

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 ull coverage at stage 3.
 will be important due
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 ge from Stage 4 will
 better background
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 cially useful for
 rring the effects of
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 3.
 ery useful for NEOs at
 ture the most active
 most. The capabilities
 s are difficult to assess
 Vector measurements
 lity growth reasons).
 ne topic

Blank – insuffi

Blank – insufficient capabilities X – full capabilities P – partial capabilities (P) – partial capabilities



EISCAT_3D Project Status

Fp6 Design Phase completed 2009

Fp7 Preparatory Phase on track for 2014 completion

Norway: application submitted Autumn 2012, revised August 2013, next submission Autumn 2014 (with clarifications)

Sweden: application submitted Spring 2013, clarifications submitted

Finland: on roadmap

Japan: on Master Plan 2014

U.K.: recently responded to a request for planning documents

Working to increase the size of the association (Affiliates)



EISCAT Scientific Association

Questions?



Stages Timeline

Ver 2014-02-27 EISCAT and EISCAT_3D combined, simulated price-level 2014 and thereafter inflated

EISCAT 2014 2015 2016 2017 2018 2019 2020 2021 2022 2023 2024

Investment budget, EISCAT_3D system only

Assuming fixed price contracts. Contingency handles currency and inflation changes

Staged Implementation: First stage only (Core + two receive arrays: Bergfors Sweden (as originally planned) and Kaaresuvanto area (new location))

EISCAT_3D construction phases	Planning	Stage 1			Timing depends on funding, example only			
					Stage 2	Stage 3	Stage 4	
Stage 1								
Core array, Norway	Prep -> Site est		Build		Commissioning - operations			
First receive array, Sweden		Site est	Build		Commissioning - operations			
Second receive array, Finland		Site est	Build		Commissioning - operations			
Op centre established, location to be decided			op centre in place - operations					
Stage 2								
Core array - 5 -> 10 MW, Norway							Installation	
Stage 3								
Third receive array							Build	
Stage 4								
Fourth receive array							Build	

End of first stage