

Unravelling long-term behaviour in historic geophysical data sets

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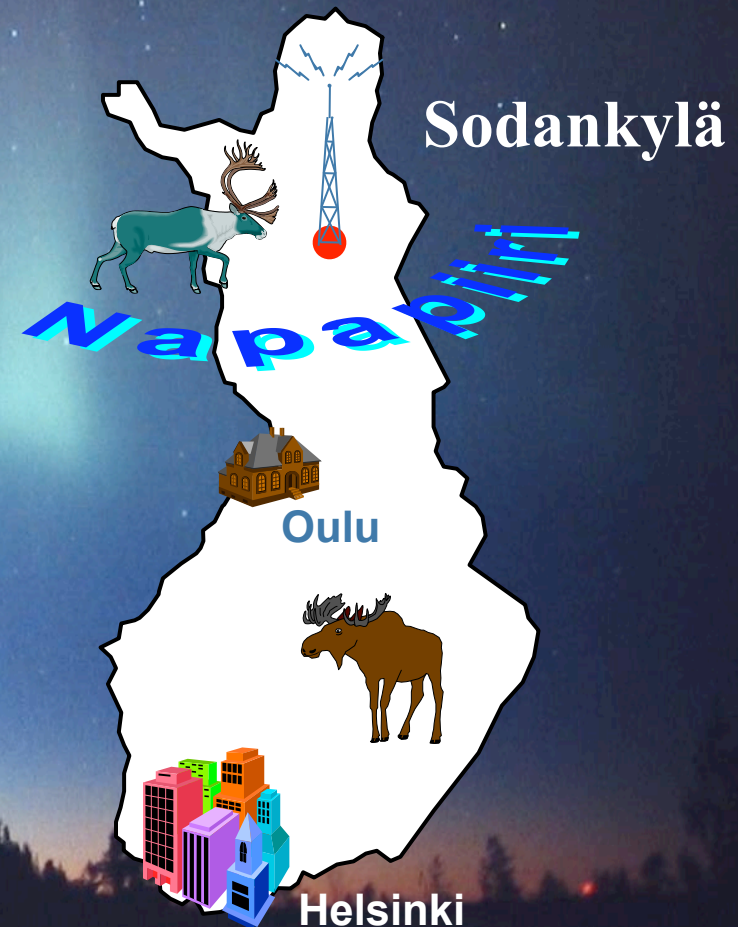
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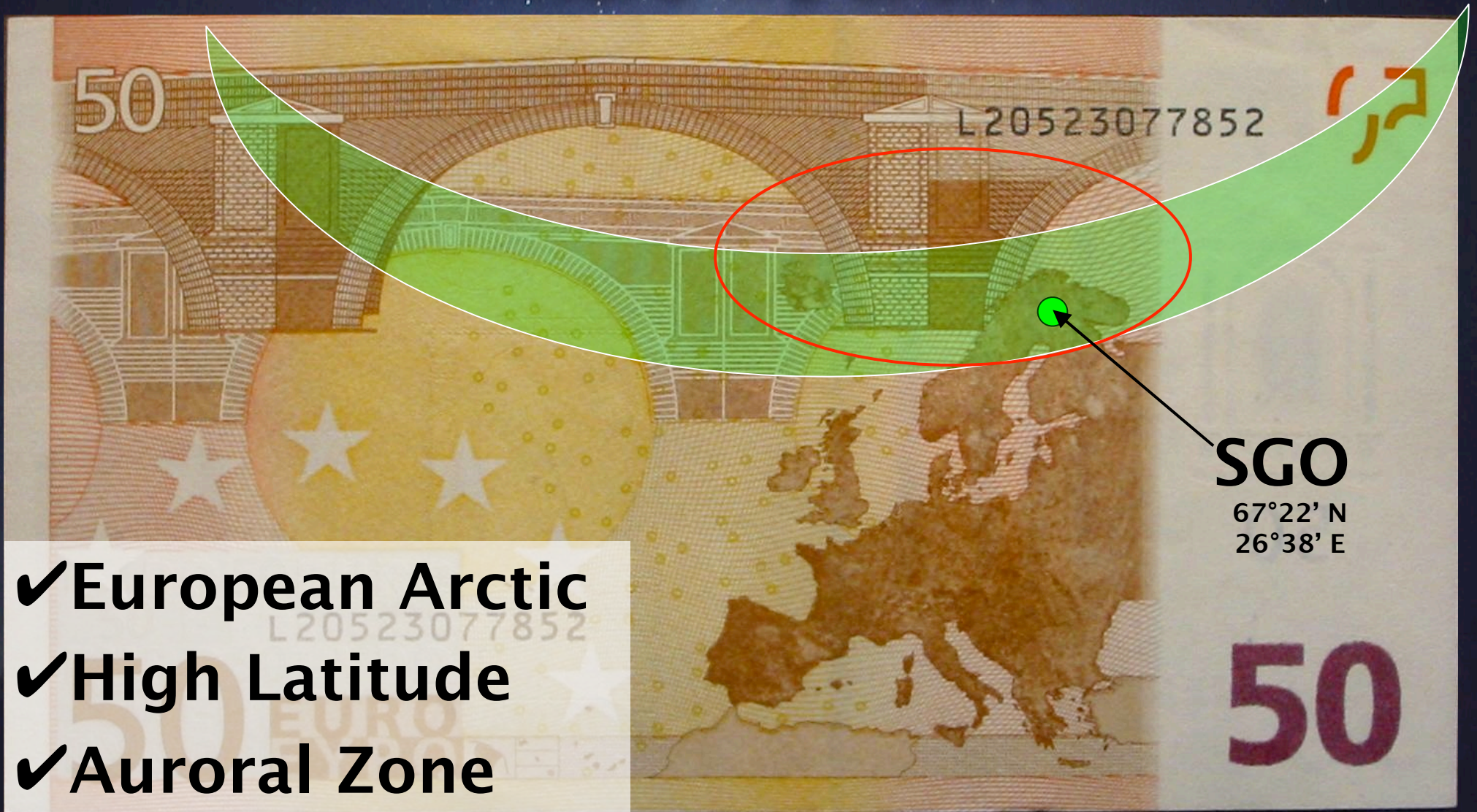
(2010-07-28)

Sodankylä Geophysical Observatory

- ➔ First observations during the International Polar Year 1882/83.
- ➔ SGO established 1913.
- ➔ Finland independent from Russia in 1917.
- ➔ SGO part of University of Oulu since 1997.
- ➔ Oldest scientific research institute in Northern Finland.



Where we are...



- ✓ European Arctic
- ✓ High Latitude
- ✓ Auroral Zone

SGO
67°22' N
26°38' E

Greenhouse Cooling

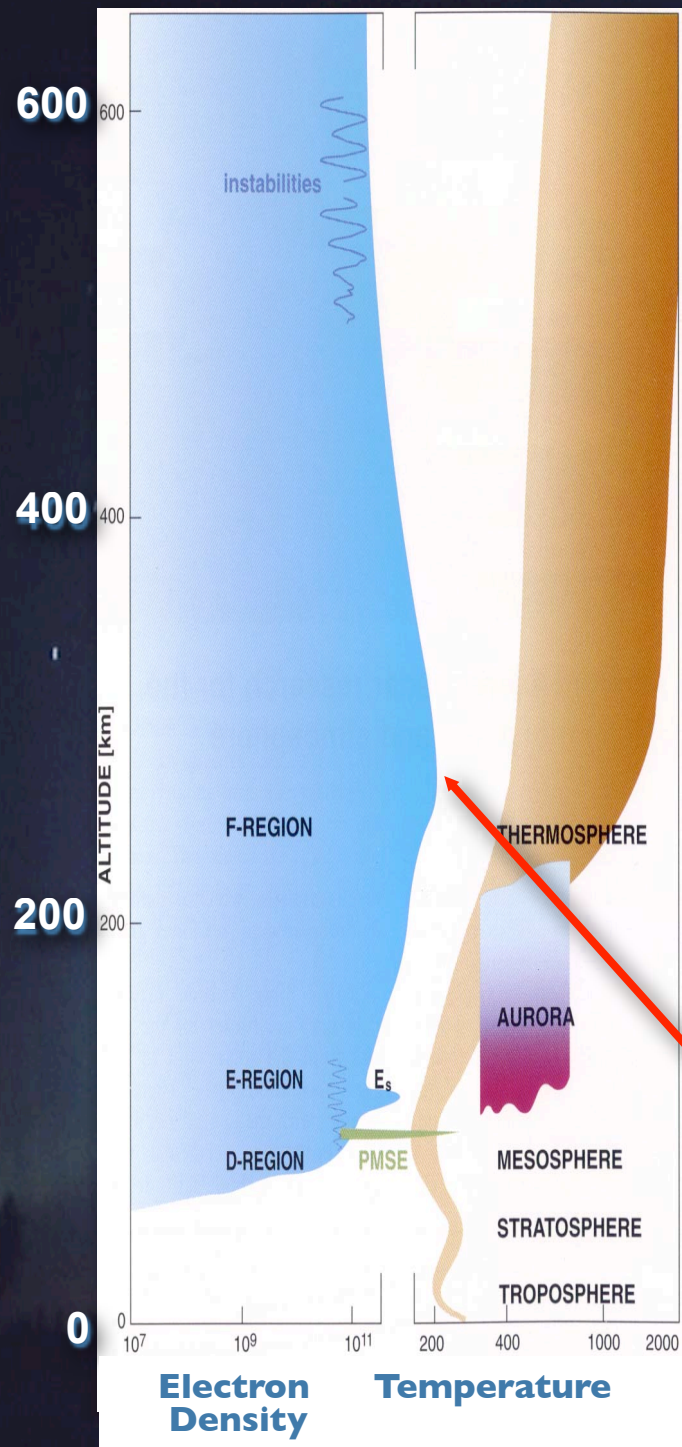
Doubling of $[\text{CO}_2]$ and $[\text{CH}_4]$

cools

Mesosphere by **10 K** and
Thermosphere by **50 K**.

Atmosphere shrinks.

Layer of maximum electron
density **lowers** by 15-20 km.



Greenhouse high up?

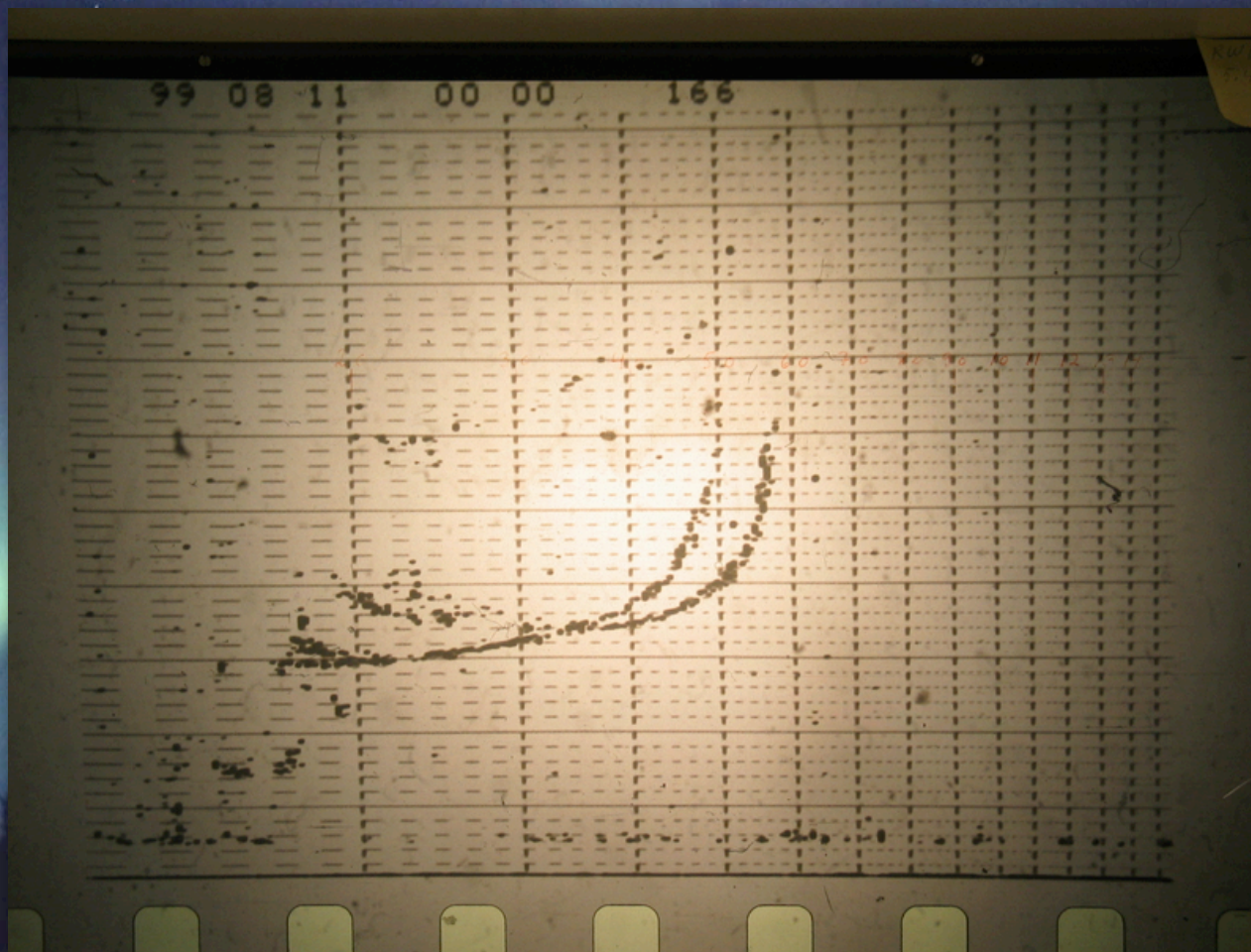
- Model results, assuming doubling of CO₂ and CH₄:
- Stratopause cools by 8 K, stratosphere by 15 K.
(Brasseur & Hitchman, 1988)
- Mesosphere and thermosphere cool by 10 K and 50 K, respectively.
(Roble & Dickinson, 1989)
- F2-layer peak (hmF2) lowers by 15-20 km.
(Rishbeth, 1990)
- Riometer absorption decreases.
(Serafimov & Serafimova, 1992)
- Stratopause cools by 14 K, mesosphere by 8 K, thermosphere by 50 K.
(Akmaev & Fomichev, 1998)

Sodankylä Ionosonde

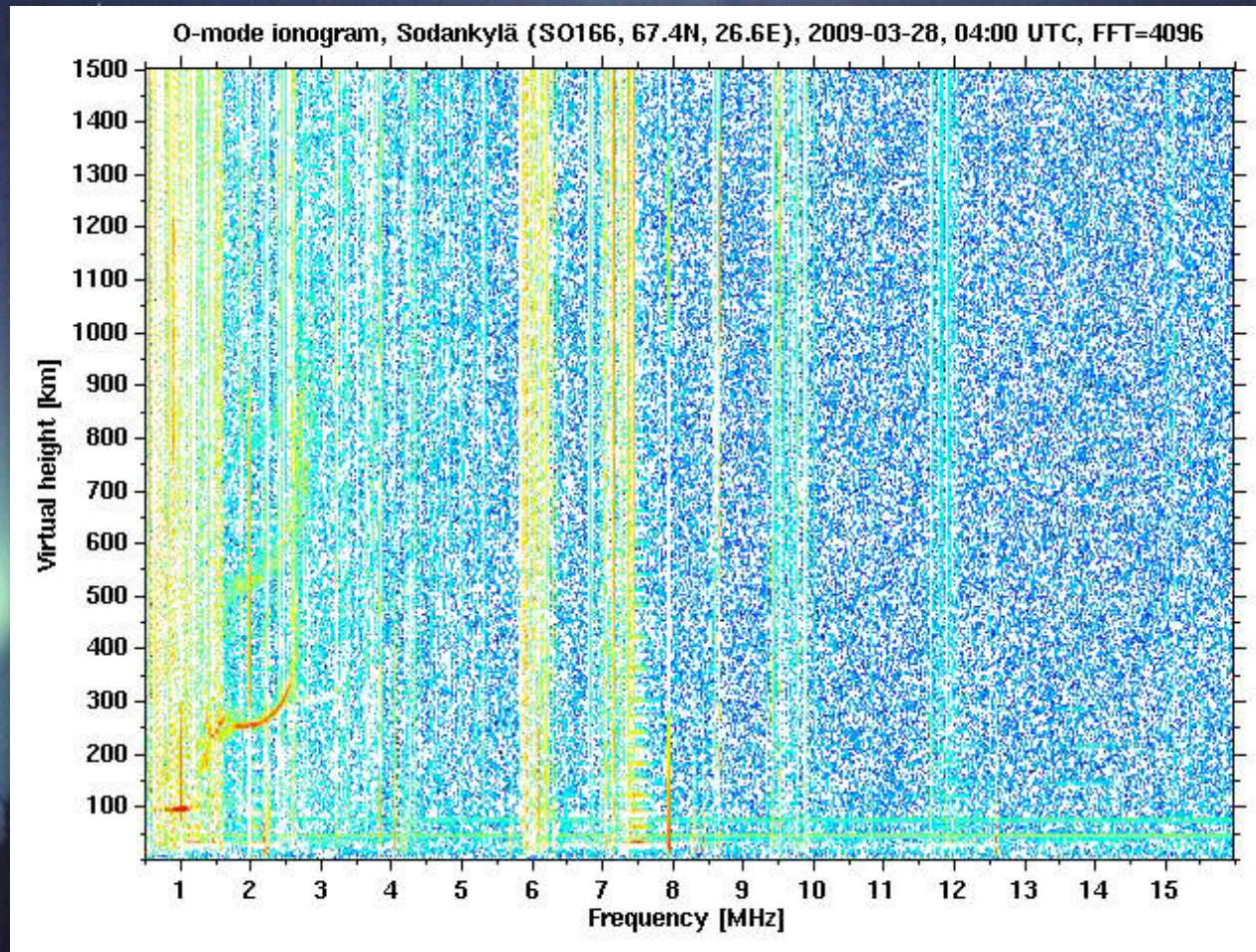
- ▶ Sodankylä ionosonde measurements began 1st August 1957.
- ▶ Until Nov 2005: 1 sounding per 30 min.
- ▶ Until Mar 2007: 1 sounding per 10 min.
- ▶ IPY (Apr '07-Mar '08): 1 sounding per minute.
- ▶ April 2008: we forgot to turn off IPY mode.
- ▶ Millionth ionogram: May 2007, at lunchtime.
- ▶ High data quality: first 800.000+ ionograms were analysed by the very same person!



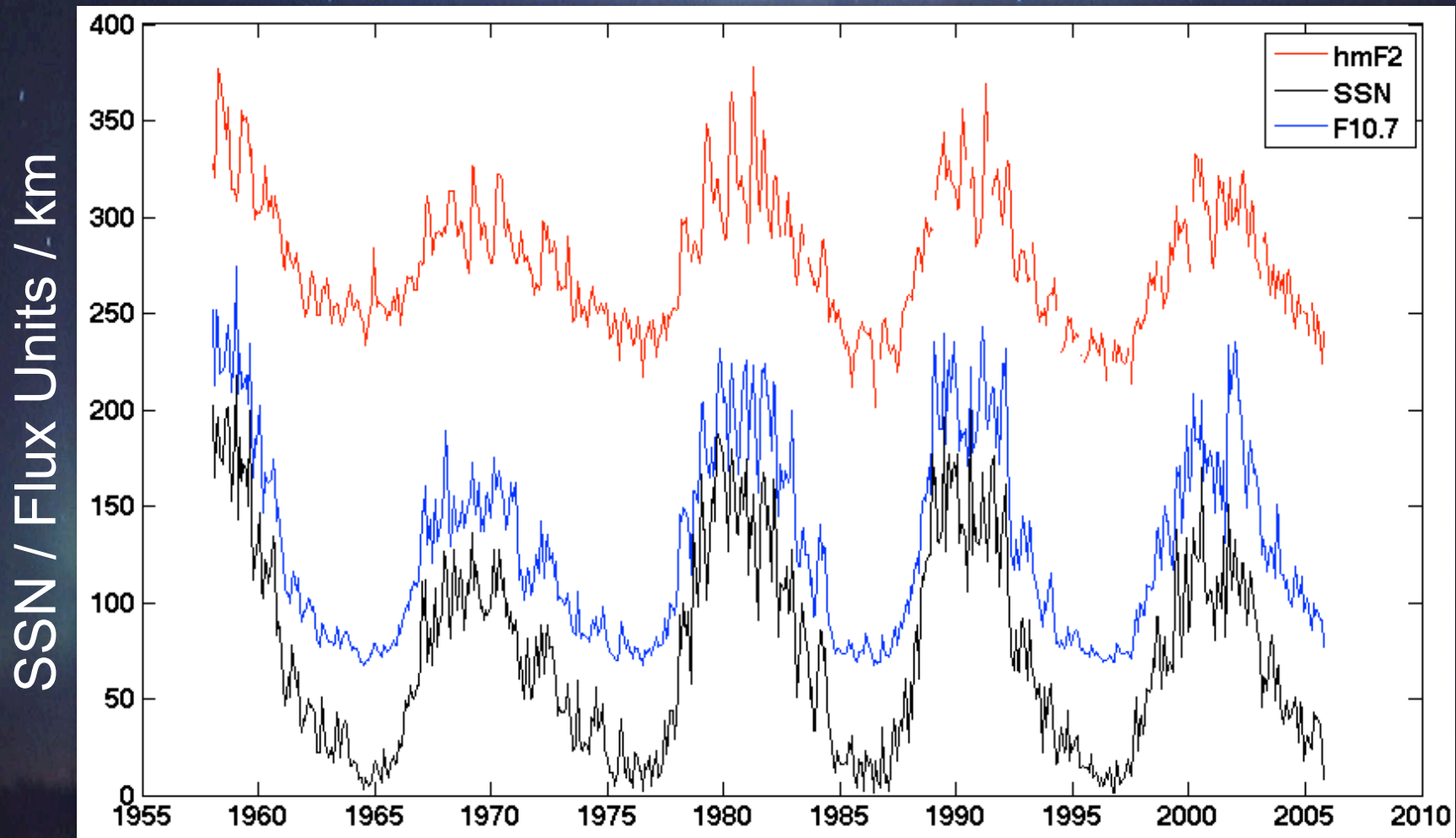
Sodankylä Ionosonde



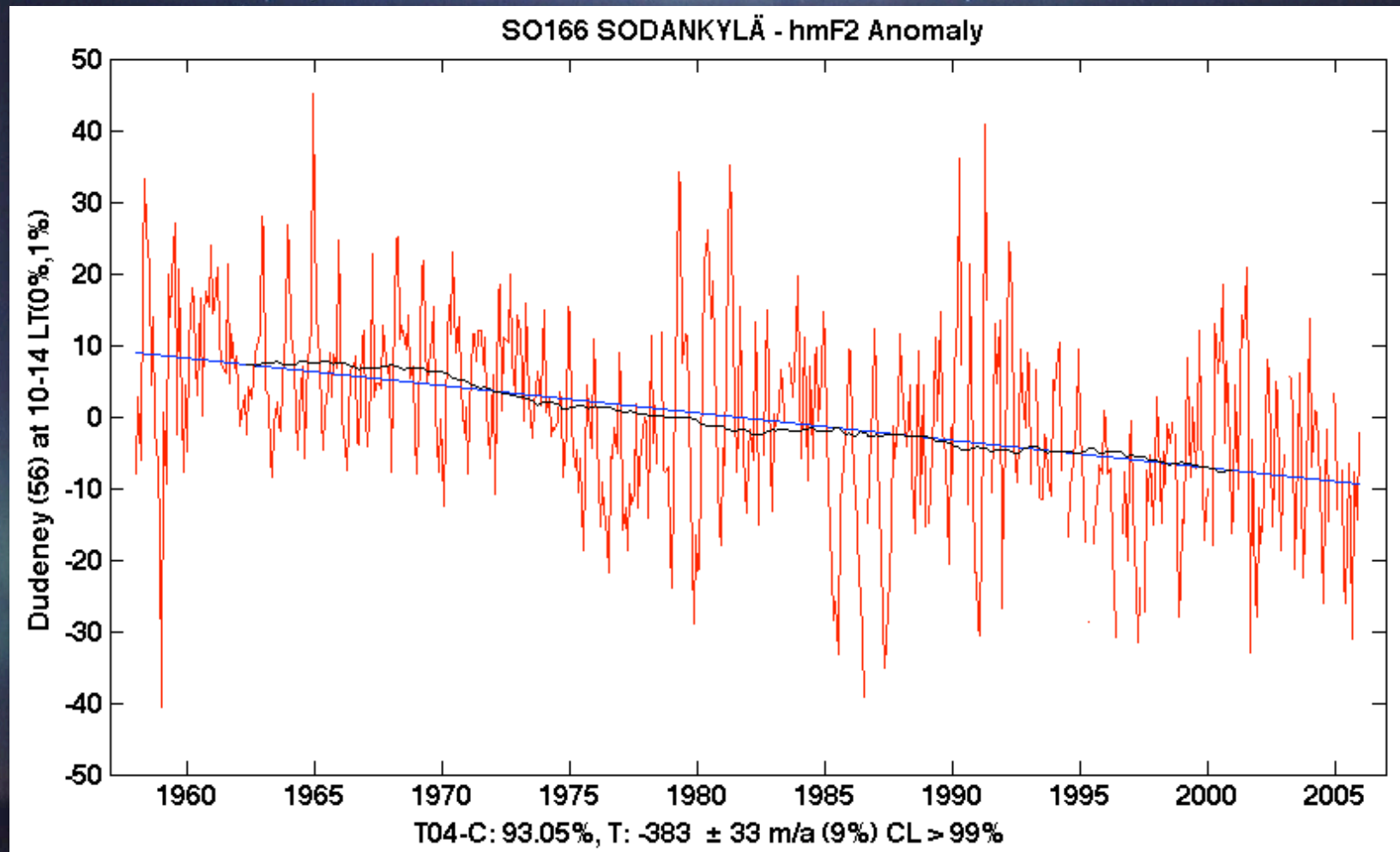
Sodankylä Ionosonde



Sodankylä hmF2 & Solar Activity



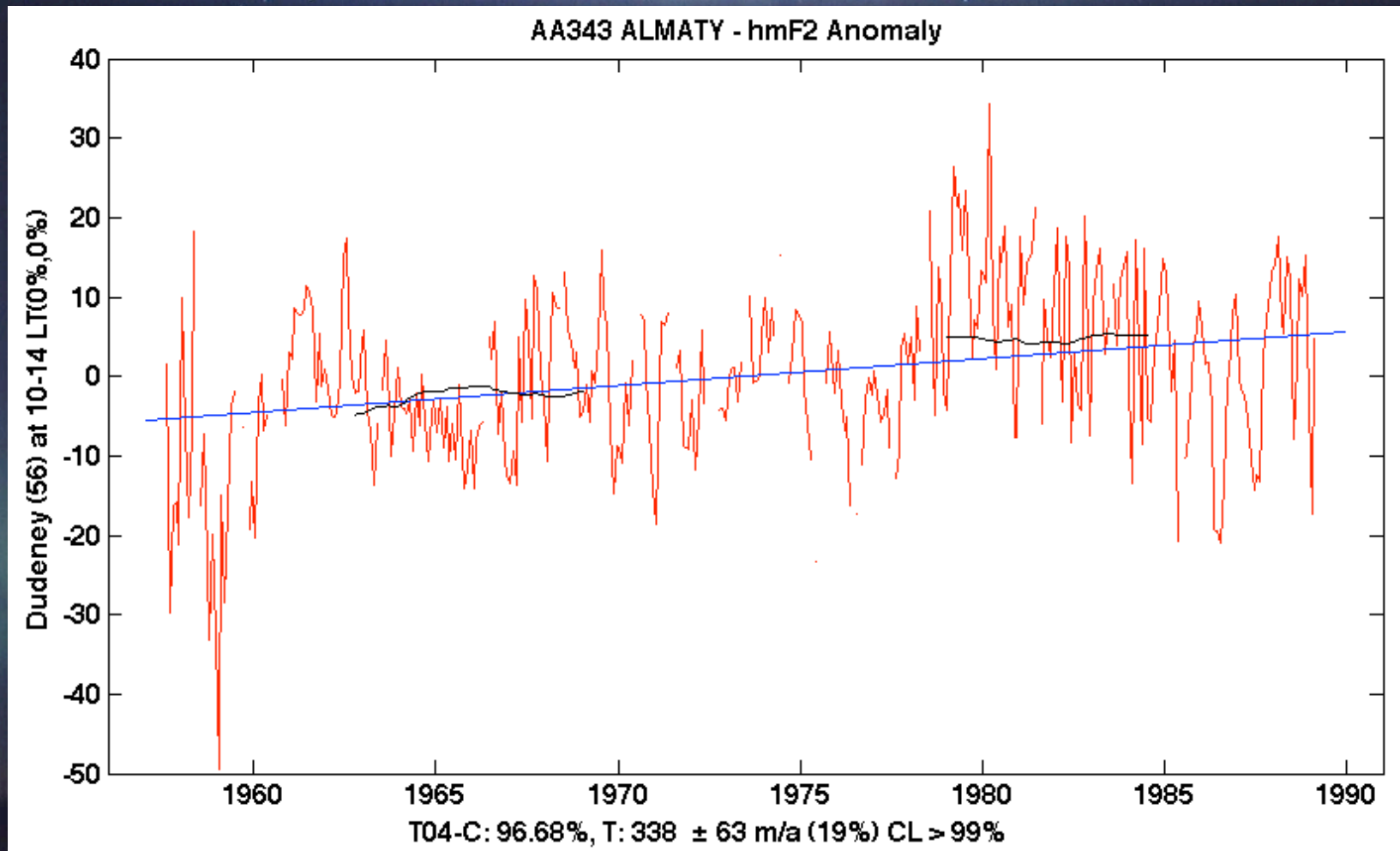
Sodankylä hmF2 Trend



Conclusion

- The enhanced greenhouse effect is clearly visible in the ionosphere.

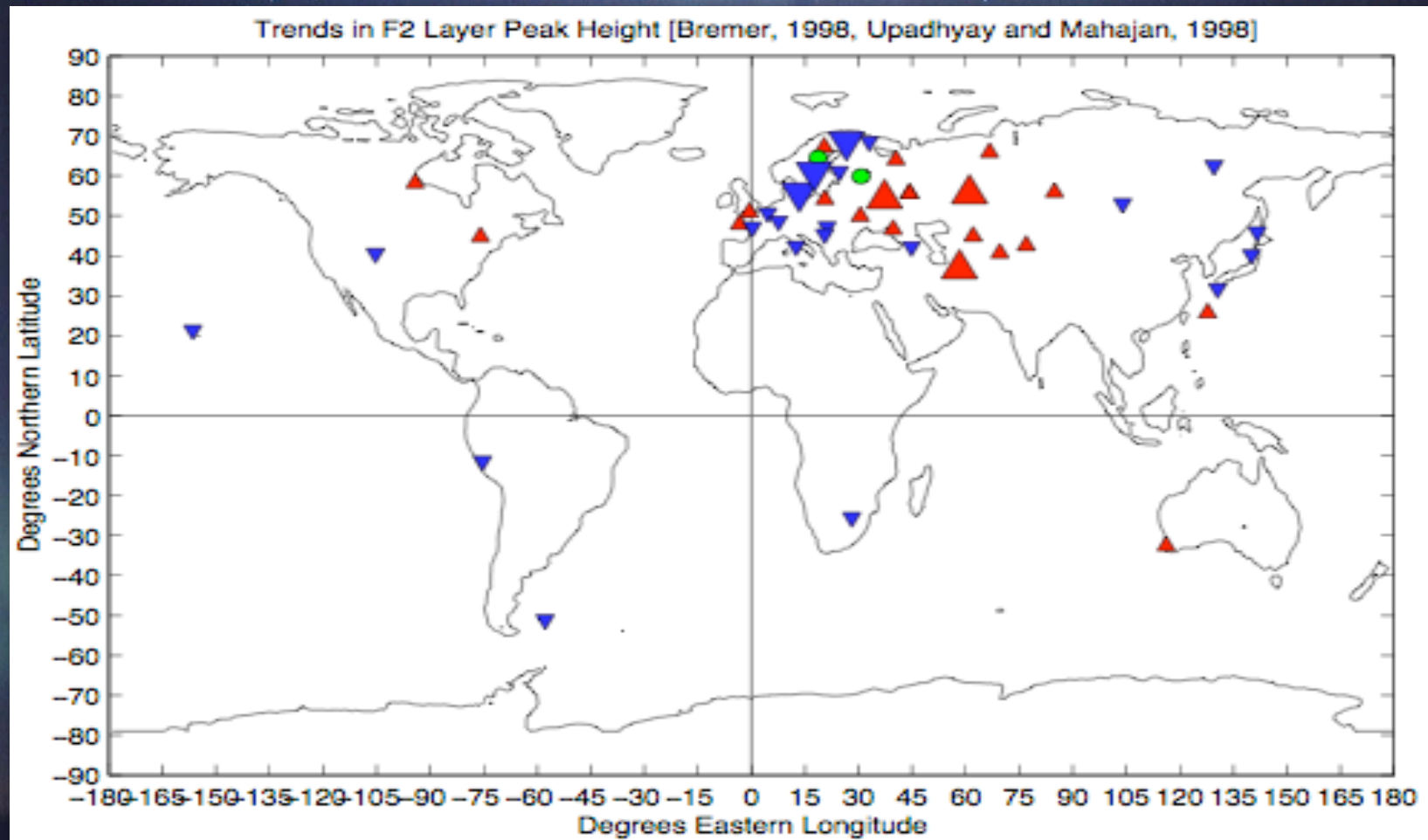
Almaty hmF2



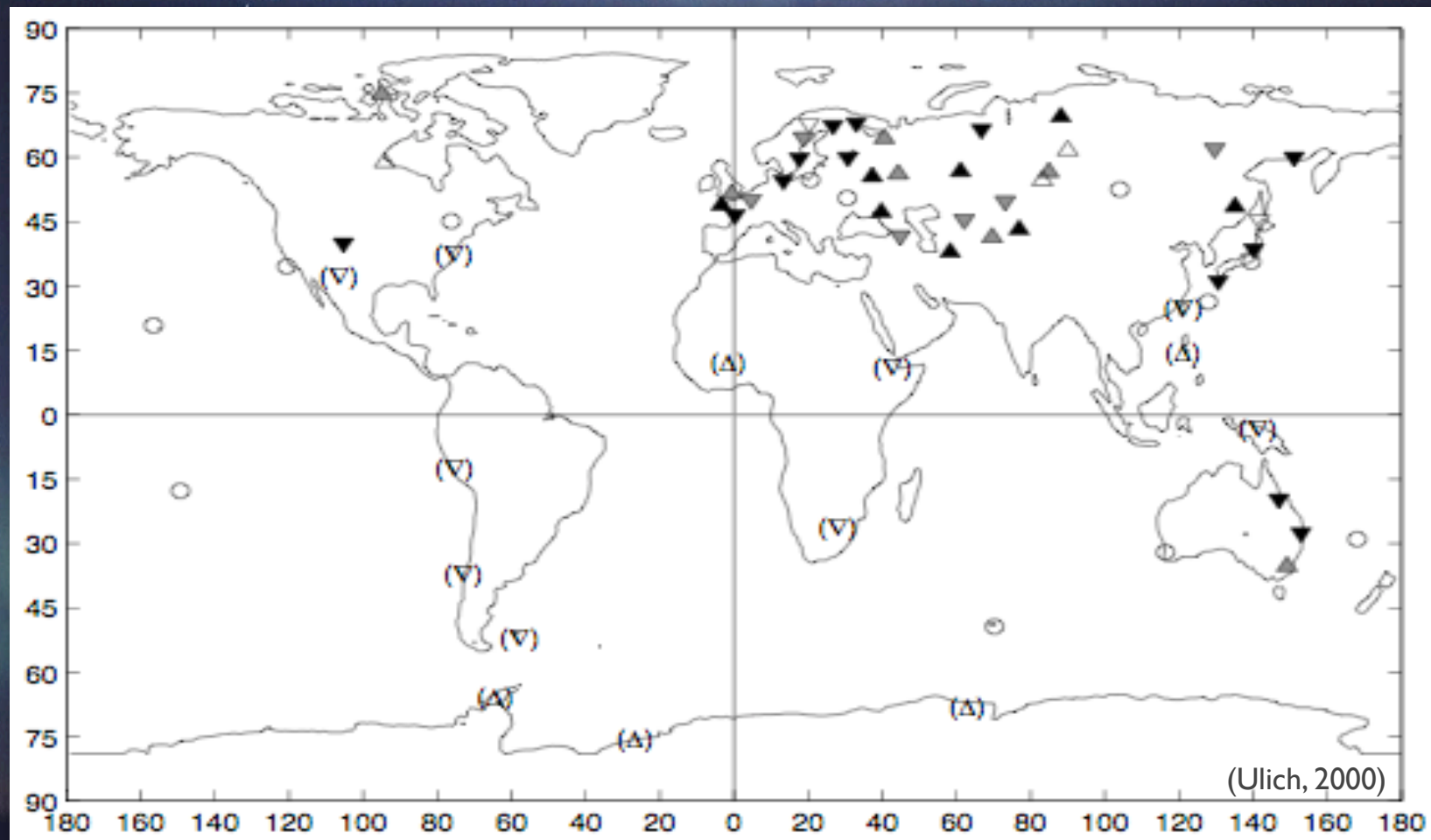
Conclusion

- Obviously, my data set is better than yours.
- The enhanced greenhouse effect is clearly visible in the ionosphere.

hmF2 Trends



Global hmF2 Trends



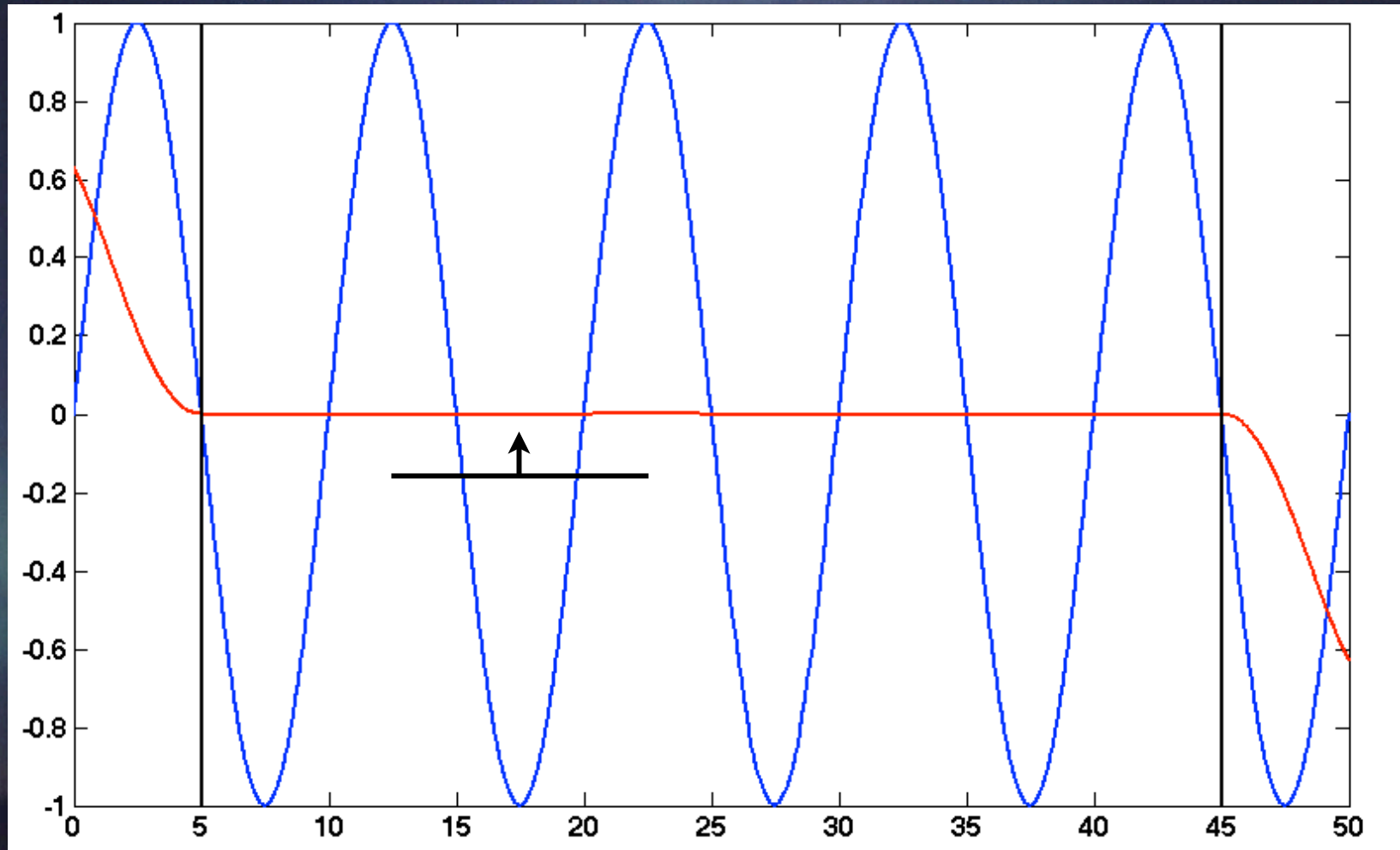
Conclusion

- What the ... ???

Problems

- Data resolution (h, 3-h, day, month(?), ...)
- Low-pass filtering or polynomial fitting...

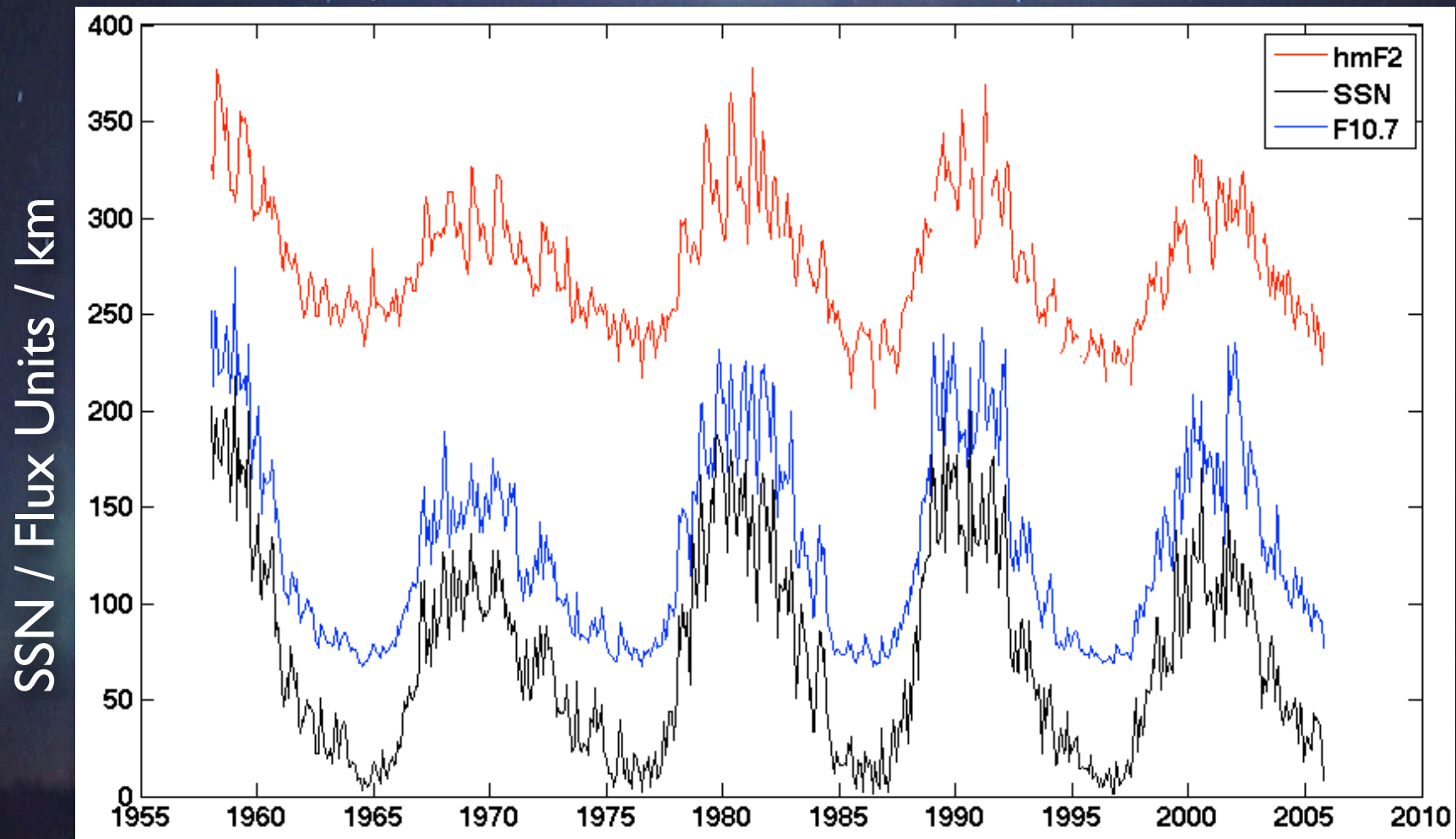
Running Mean Filter



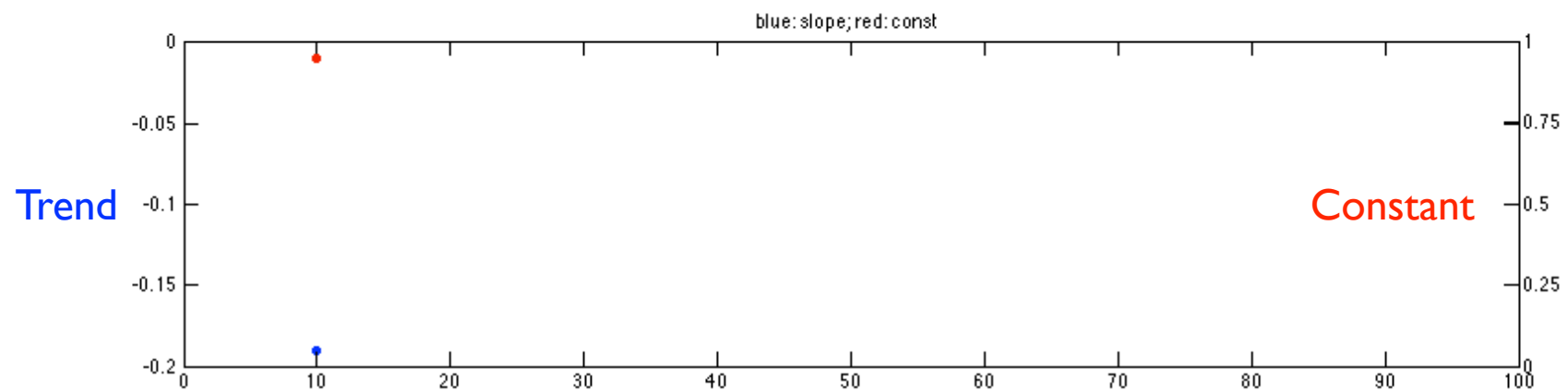
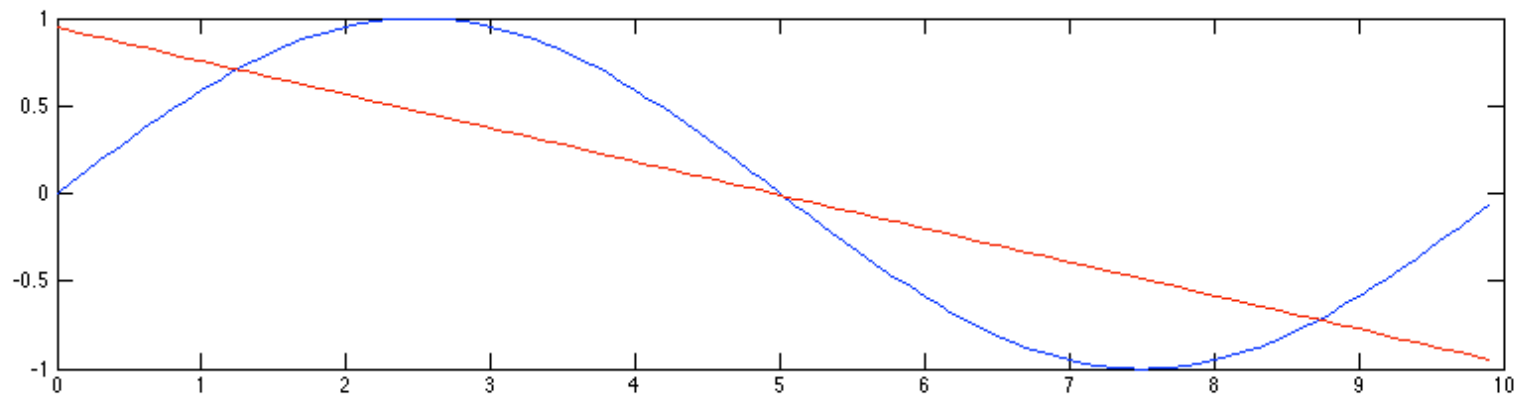
Problems

- Data resolution (h, 3-h, day, month(?), ...)
- Low-pass filtering or polynomial fitting...
- Removal of underlying (cyclic) variability:
 - Choice of proxy (sinusoid, SSN, Group SSN, F10.7 (adj./obs.), Ly- α , Mg II, E10.7, ...)
 - Resolution of proxy: compatibility with data

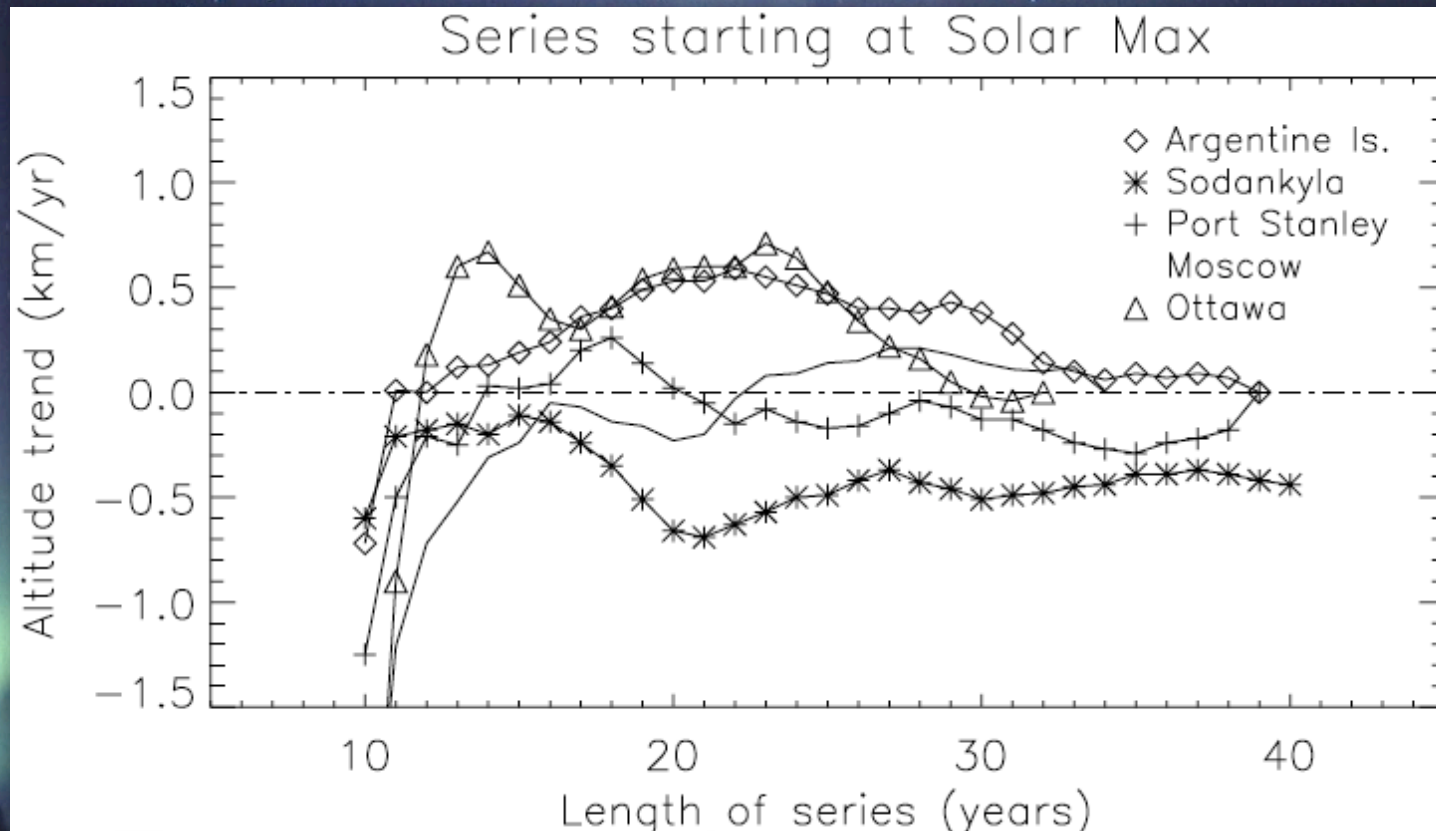
Sodankylä hmF2 & Solar Activity



Ringing



Ringing

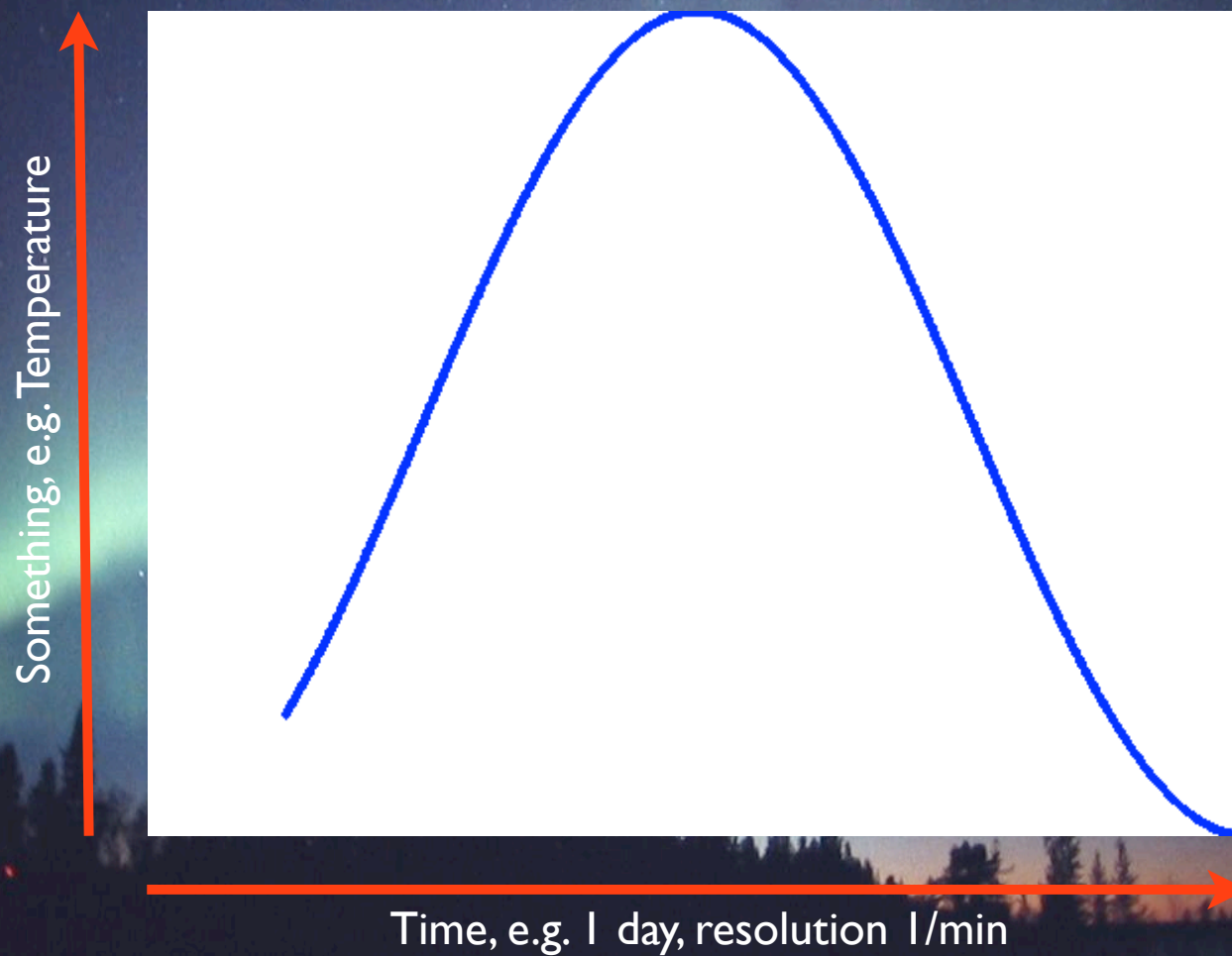


The ringing idea was first introduced by Jarvis et al., 2002. The plots shown here are from a follow-up paper by Clilverd et al., 2003.

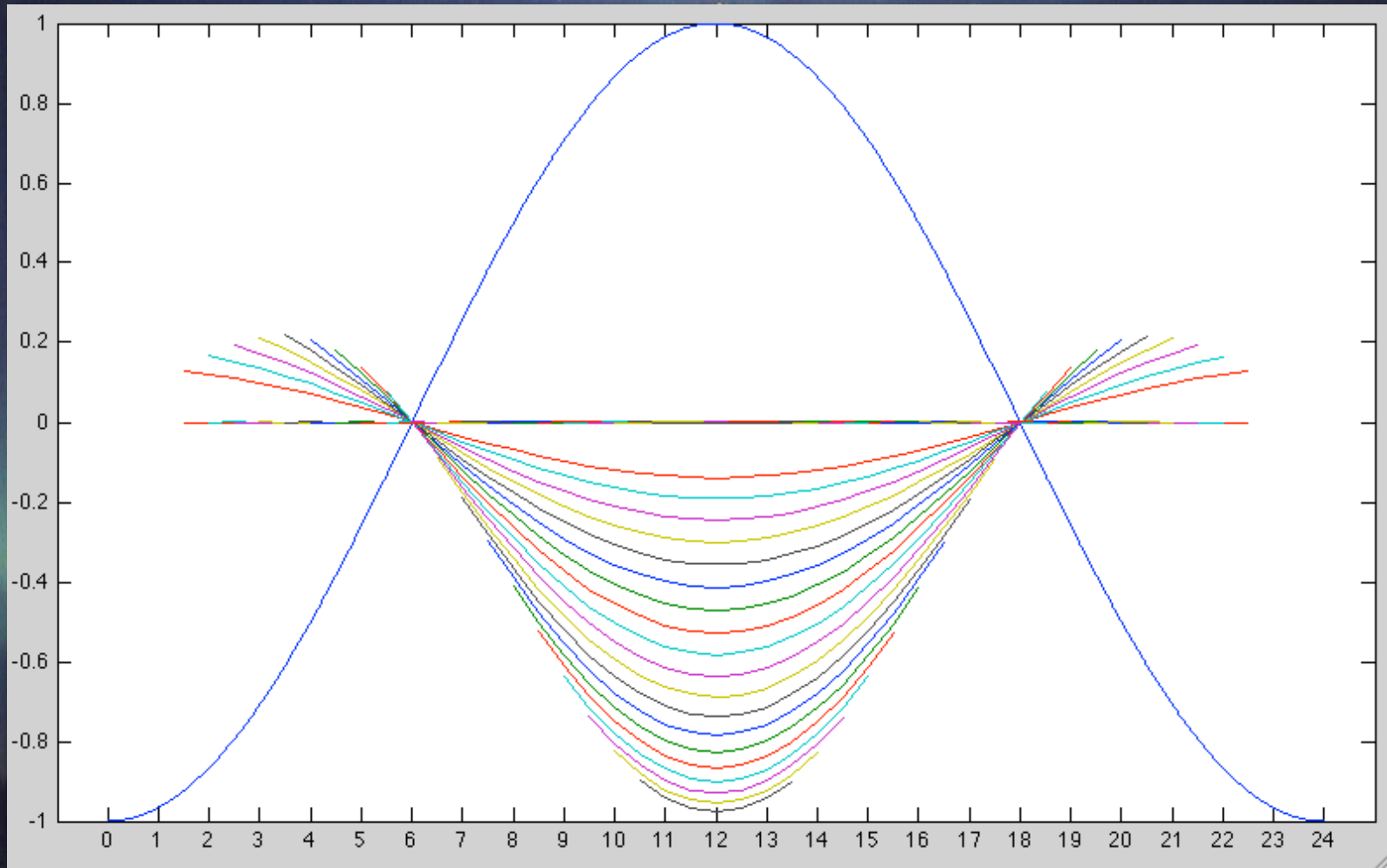
Problems

- Data resolution (h, 3-h, day, month(?), ...)
- Low-pass filtering or polynomial fitting...
- Removal of underlying (cyclic) variability: ...
- Data gaps

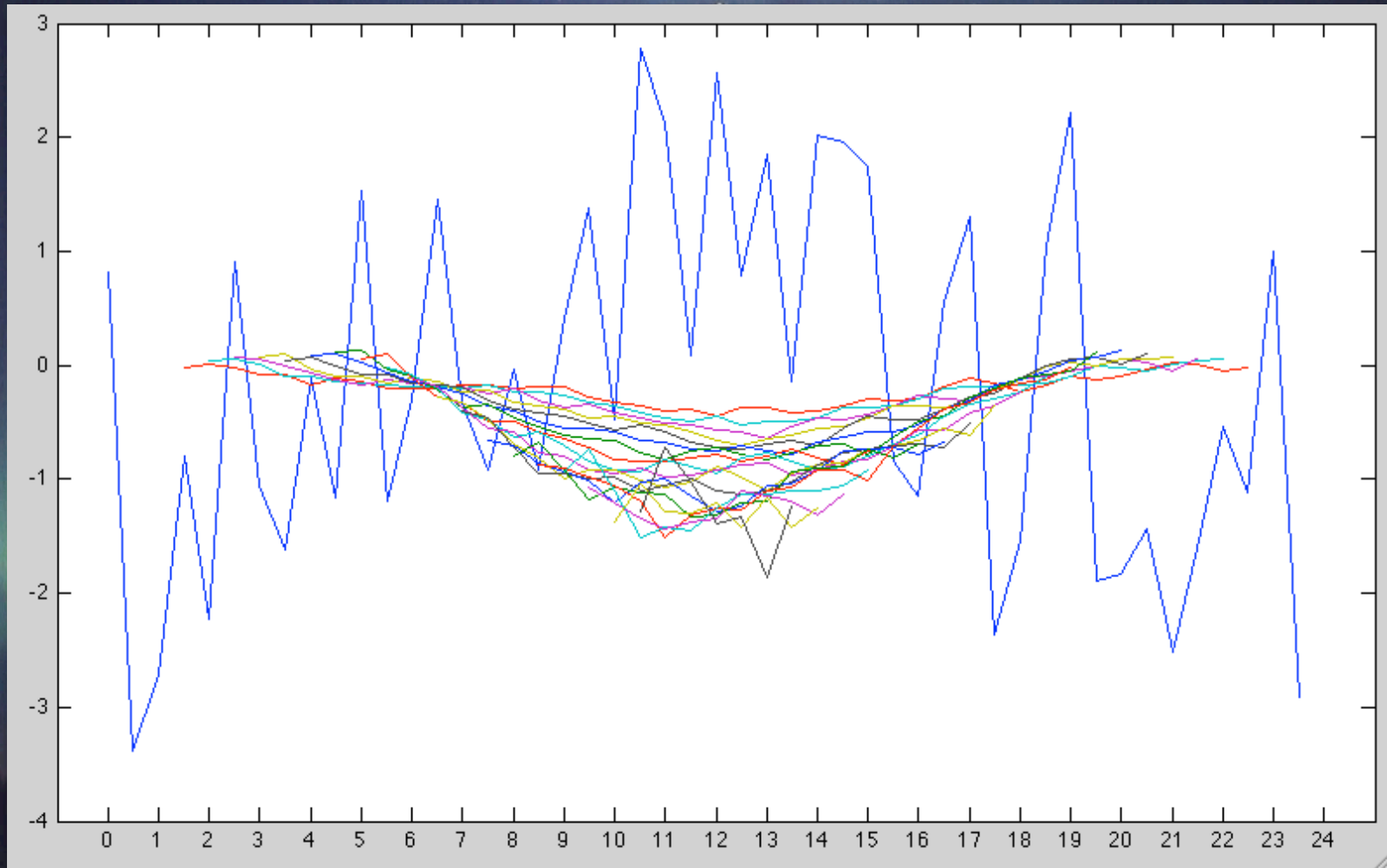
Example: Data Gaps



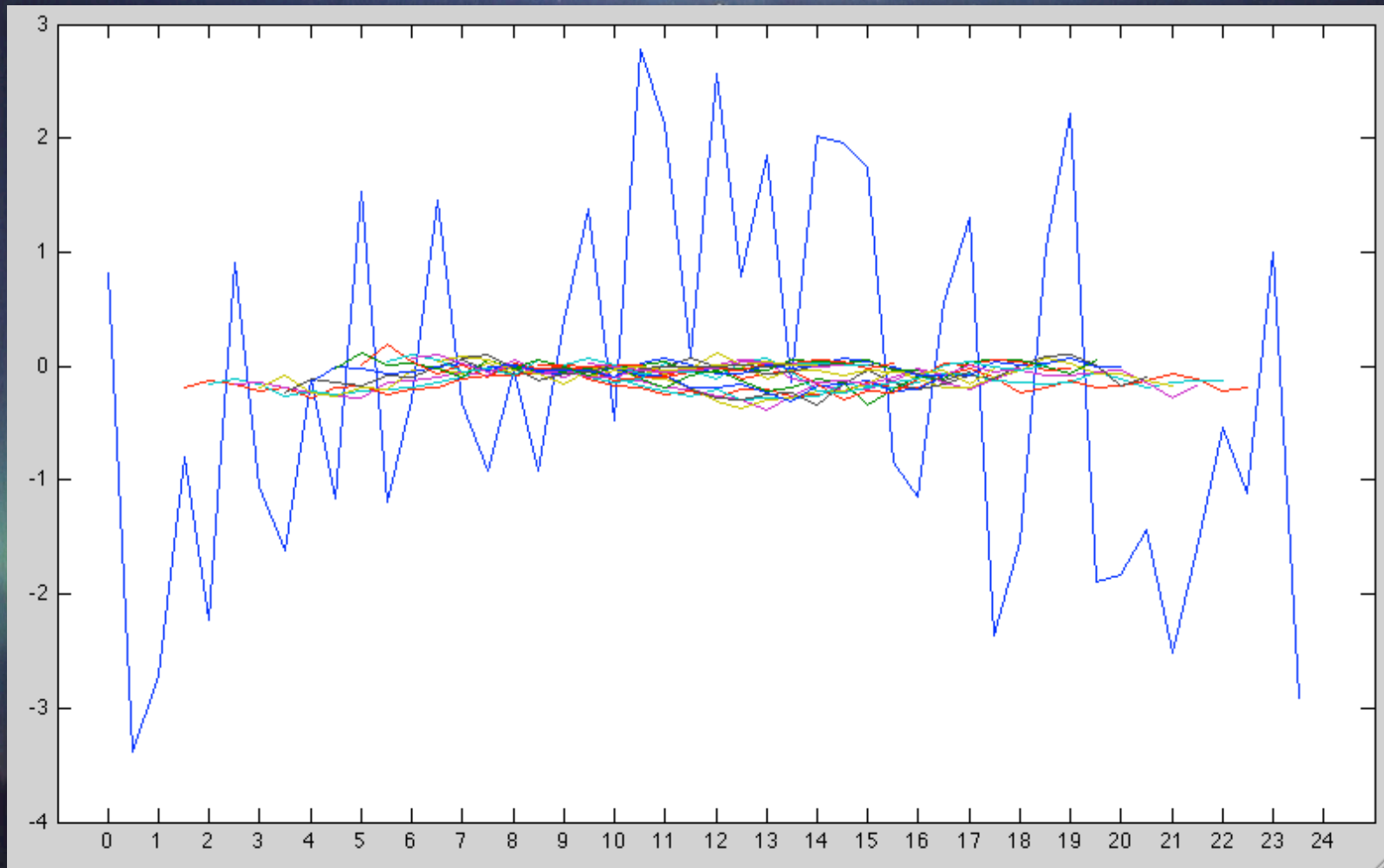
Data Gaps



Data Gaps



Data Gaps



Problems

- Data resolution (h, 3-h, day, month(?), ...)
- Low-pass filtering or polynomial fitting...
- Removal of underlying (cyclic) variability: ...
- Data gaps
- Measurement errors
- Mathematics of trend detection
 - stepwise or multi-parameter fit
 - error propagation

Making models

- Base functions of the model(s) are, e.g.:

$$m_i = \varepsilon_i \quad \rightarrow \text{measurement errors}$$
$$+ x_1 \quad \rightarrow \text{constant}$$
$$+ x_2 t_i \quad \rightarrow \text{sampling times}$$
$$+ x_3 F_{10.7}(t_i) \quad \rightarrow \text{solar activity}$$
$$+ x_4 A_p(t_i) \quad \rightarrow \text{geomagnetic activity}$$
$$+ x_5 \sin(2\pi t_i)$$
$$+ x_6 \cos(2\pi t_i) \quad \rightarrow \text{annual variation}$$
$$+ x_7 \sin(4\pi t_i)$$
$$+ x_8 \cos(4\pi t_i) \quad \rightarrow \text{semi-annual variation}$$
$$+ \dots$$

Modelling the data

The ionospheric property of interest is function of time and a number of other parameters. The model of the data is therefore

$$m(t) = \mathcal{F}(t, x_1, \dots, x_M)$$

where

$$\mathcal{F}(t, x_1, \dots, x_M) = \sum_{i=1}^M x_i f_i(t)$$

The actual measurements m_i observed at time t_i are equal to the model plus some measurement error ε_i

$$m_i = \mathcal{F}(t_i, x_1, \dots, x_M) + \varepsilon_i$$

Inverse problem I

This can be expressed as a matrix equation. Usually there are many more data points than unknowns x_i and the problem is over-determined:

$$\begin{pmatrix} m_1 \\ m_2 \\ \vdots \\ m_N \end{pmatrix} = \begin{pmatrix} f_1(t_1) & f_2(t_1) & \cdots & f_M(t_1) \\ f_1(t_2) & f_2(t_2) & \cdots & f_M(t_2) \\ \vdots & \vdots & \ddots & \vdots \\ f_1(t_N) & f_2(t_N) & \cdots & f_M(t_N) \end{pmatrix} \cdot \begin{pmatrix} x_1 \\ x_2 \\ \vdots \\ x_M \end{pmatrix} + \begin{pmatrix} \varepsilon_1 \\ \varepsilon_2 \\ \vdots \\ \varepsilon_N \end{pmatrix}$$

In other words:

$$\mathbf{m} = \mathbf{A} \cdot \mathbf{x} + \boldsymbol{\varepsilon}$$

Inverse problem II

Measurements and theory are weighted by the measurement errors:

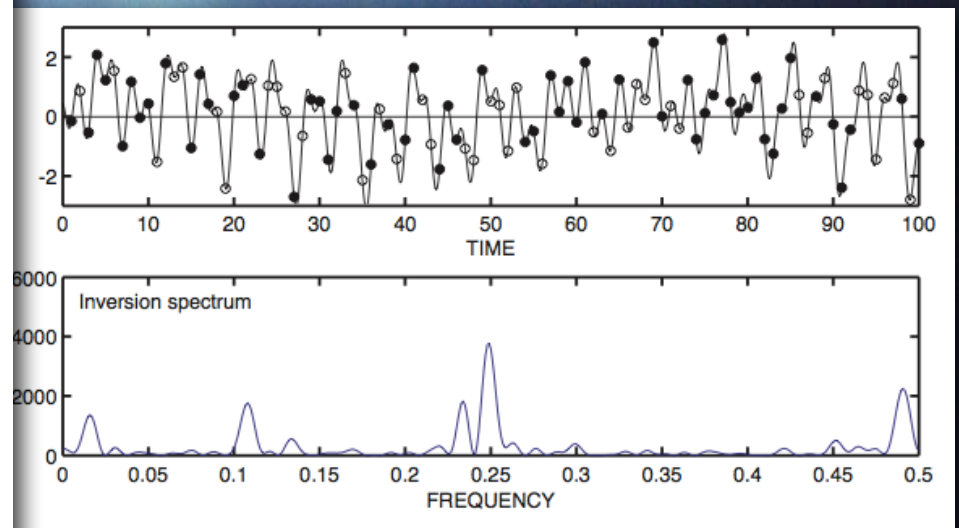
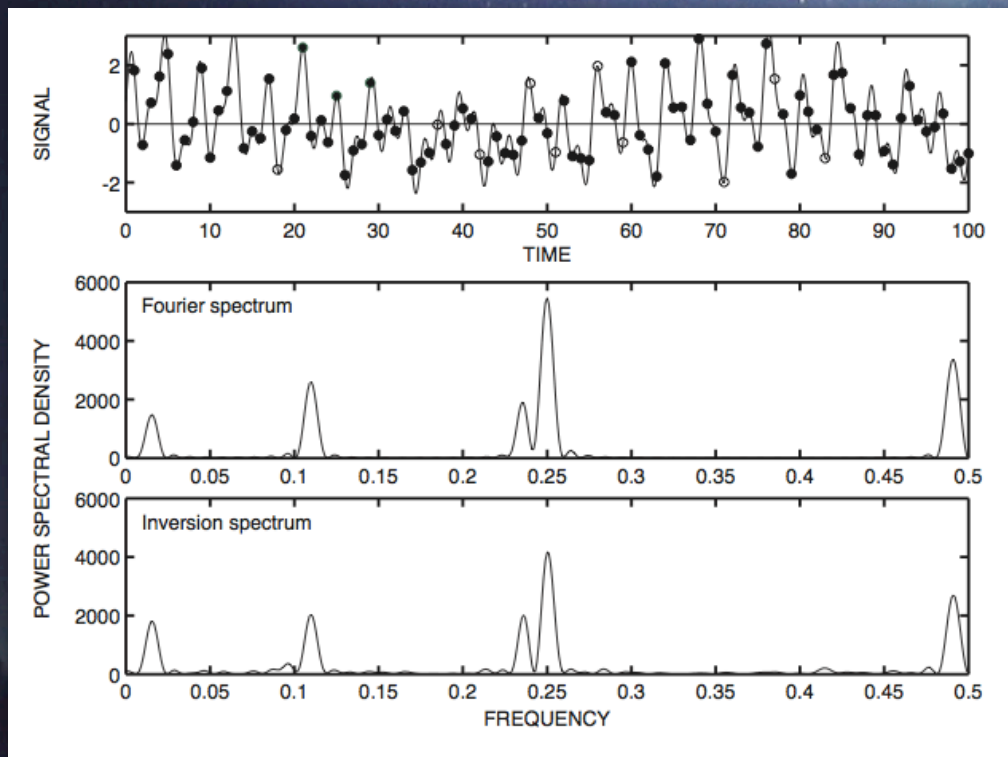
$$B_{ij} := \frac{A_{ij}}{\varepsilon_i} \quad \text{and} \quad b_i := \frac{m_i}{\varepsilon_i}$$

The solution is the vector \mathbf{x} , which minimises the following expression:

$$\chi^2 = |\mathbf{B} \cdot \mathbf{x} - \mathbf{b}|^2$$

We are left with a general least squares problem. Solving this results in the most probable solution for \mathbf{x} .

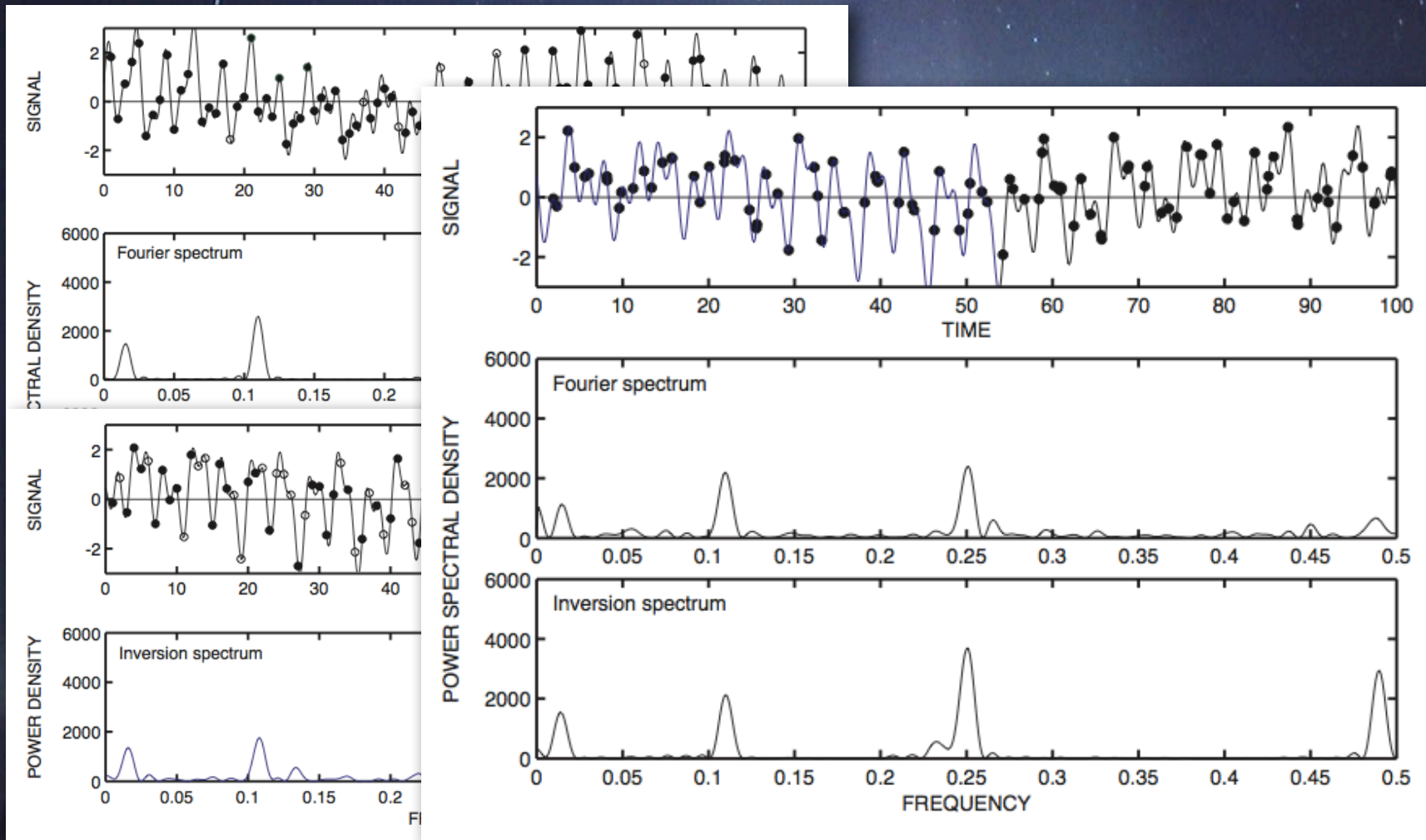
Signal Spectrum by Stochastic Inversion



Left: 100 pts for Fourier, 90 for inversion.
Above: 59 pts.

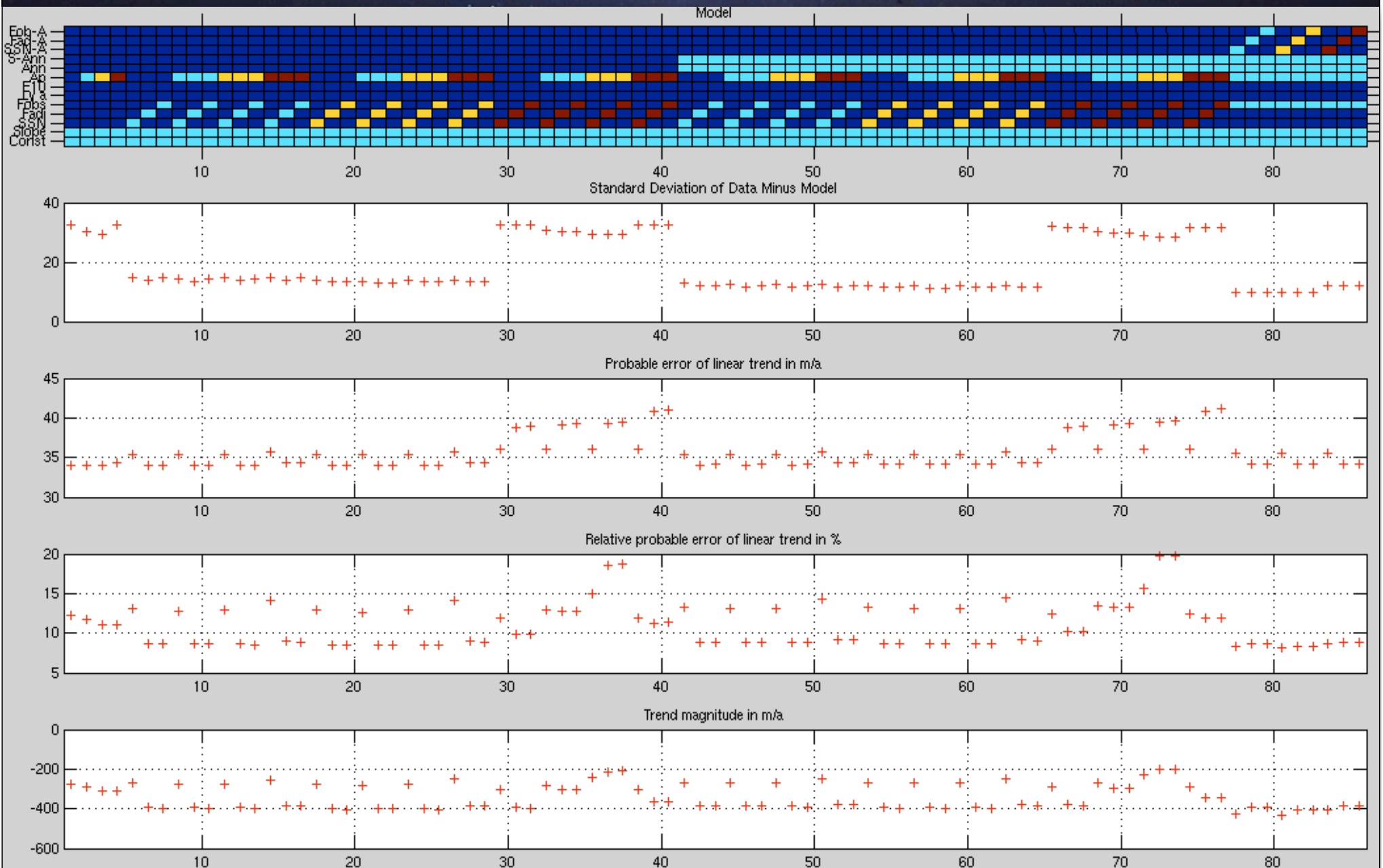
T. Nygrén and Th.Ulich, Calculation of signal spectrum by means of stochastic inversion, Ann. Geophys., 28, 1409-1418, 2010.

Signal Spectrum by Stochastic Inversion



T. Nygrén and Th.Ulich, Calculation of signal spectrum by means of stochastic inversion, Ann. Geophys., 28, 1409-1418, 2010.

Sodankylä F2-layer peak height hmF2



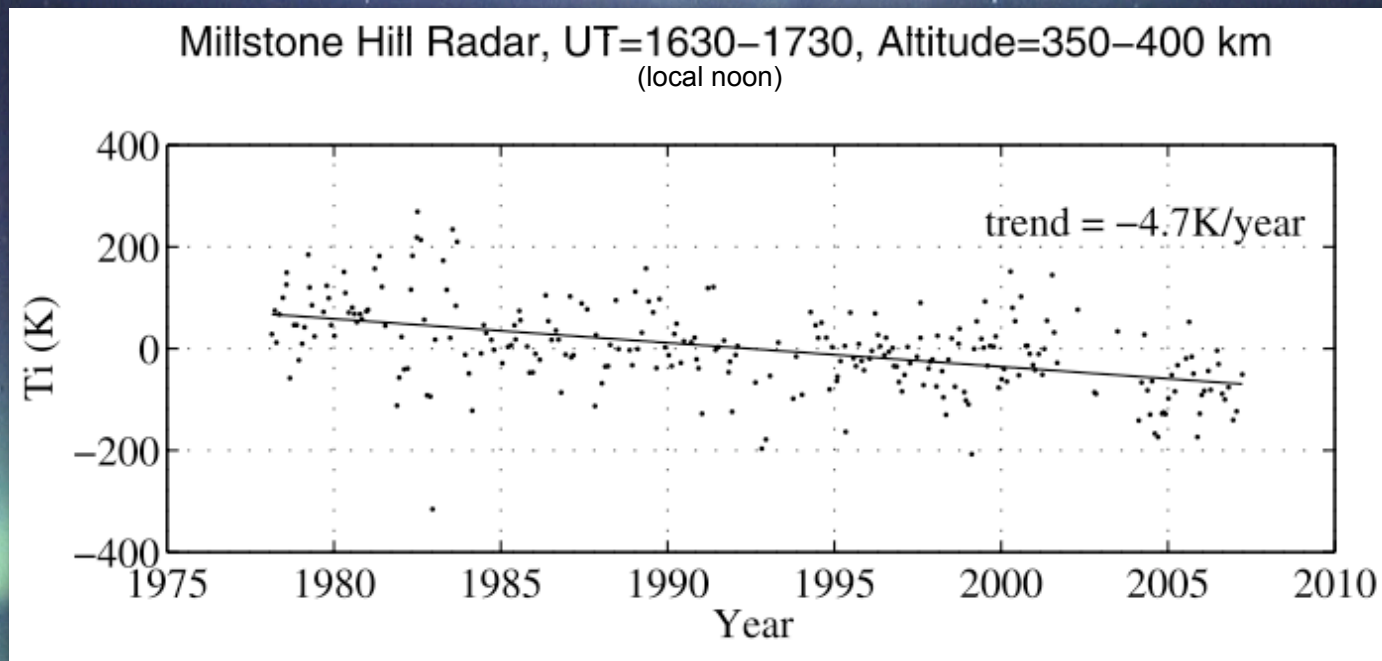
Conclusion

- This is pointless?
I don't think so...
...yet!

Trends in other Observations

Height in km	Method	Parameter	Trend per Year	Reference
75	Sounding rocket	Temperature	-0.6 K	Kokin and Lysenko, 1994
70	Sounding rocket	Temperature	-0.7 K	Golitsyn et al., 1996
60-70	Lidar	Temperature	-0.4 K	Hauchecorne et al., 1991
60	Sounding rocket	Temperature	-0.4 K	Golitsyn et al., 1996
60	Sounding rocket	Temperature	-0.33 K	Keckhut et al., 1999
50-60	Lidar	Temperature	-0.25 K	Aikin et al., 1991
50	Sounding rocket	Temperature	-0.25 K	Golitsyn et al., 1996
40	Sounding rocket	Temperature	-0.1 K	Golitsyn et al., 1996
30-60	Sounding rocket	Temperature	-0.17 K	Dunkerton et al., 1998
30-50	Sounding rocket	Temperature	-0.17 K	Keckhut et al., 1999
30	Sounding rocket	Temperature	-0.1 K	Golitsyn et al., 1996
25	Sounding rocket	Temperature	-0.1 K	Golitsyn et al., 1996
25	Sounding rocket	Temperature	-0.11 K	Keckhut et al., 1999

Direct F-Region Temperature



Long-term temperature trends in the ionosphere above Millstone Hill

J. M. Holt¹ and S. R. Zhang¹

GEOPHYSICAL RESEARCH LETTERS, VOL. 35, L05813, doi:10.1029/2007GL031148, 2008

Conclusion

(the last one, I promise!)

- Definitely, there's long-term change in the ionosphere and thermosphere!
- The enhanced greenhouse effect is probably a part of it.
- Other (unknown?) processes are involved.
- Do not discontinue long-term measurements of our (space) environment!
- Solution in modelling?
- We don't understand what's going on. Find it out!

Conclusion

(I was lying!)

Ionosondes, originally deployed for monitoring ionospheric conditions for HF radio communication and for studying short-term events, are becoming useful in an environmental context.

They provide long-term measurements of our environment!



Photo: Sally Ulich



KIITOS!